

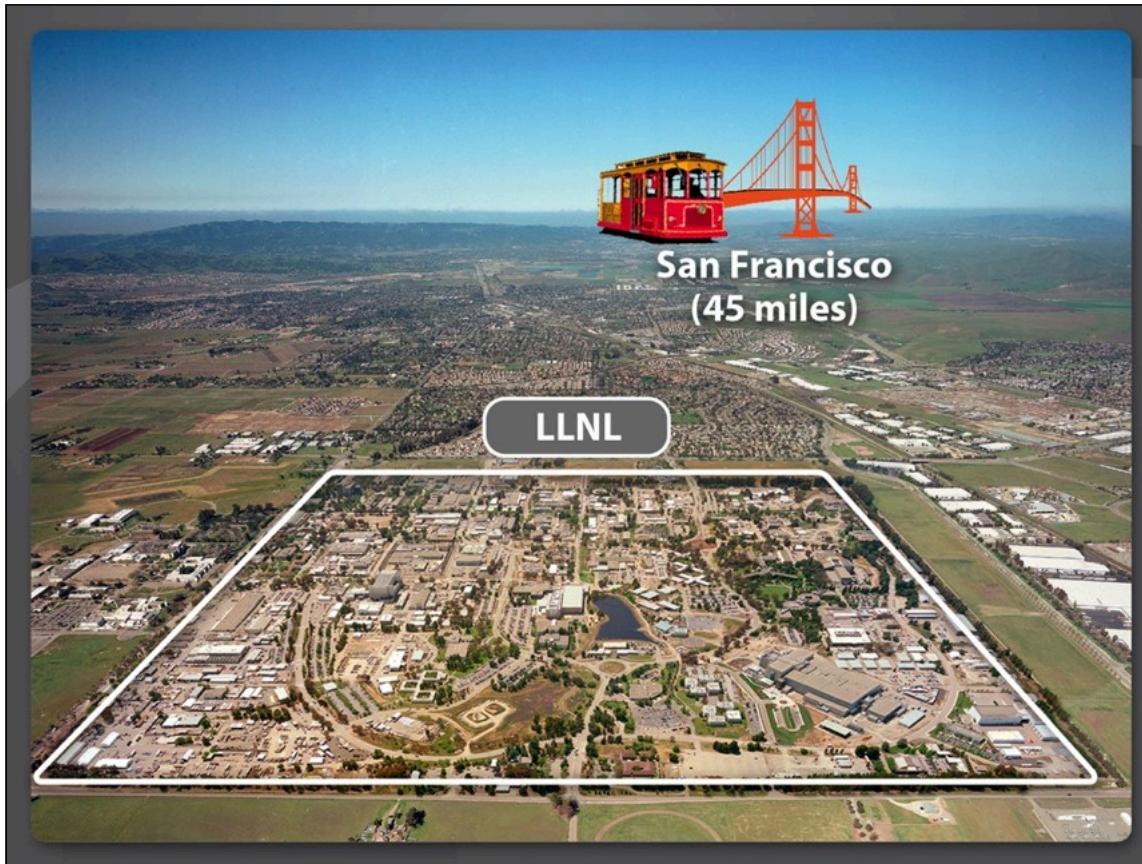
## Beyond the Synchrotron: The Dawn of Laser-Compton X-ray and Gamma-ray Science

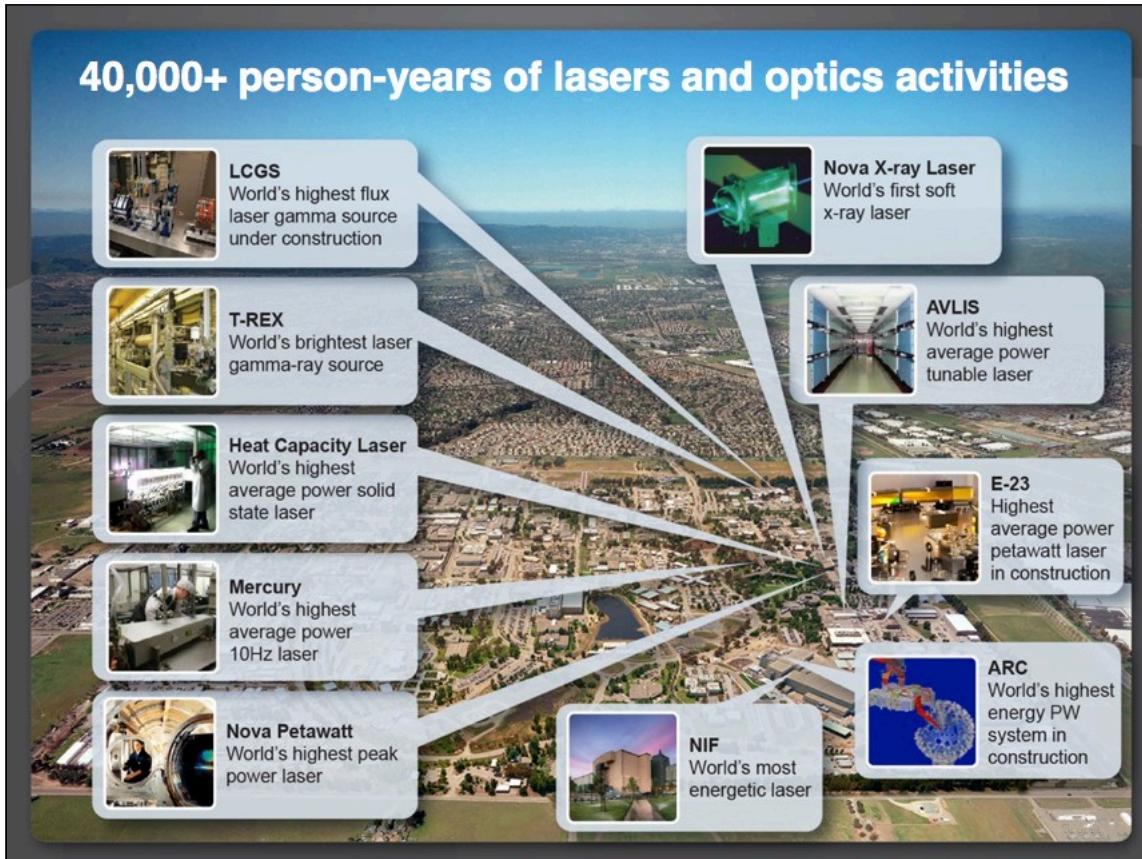
Norman Rostoker Distinguished Lecture  
University of California, Irvine



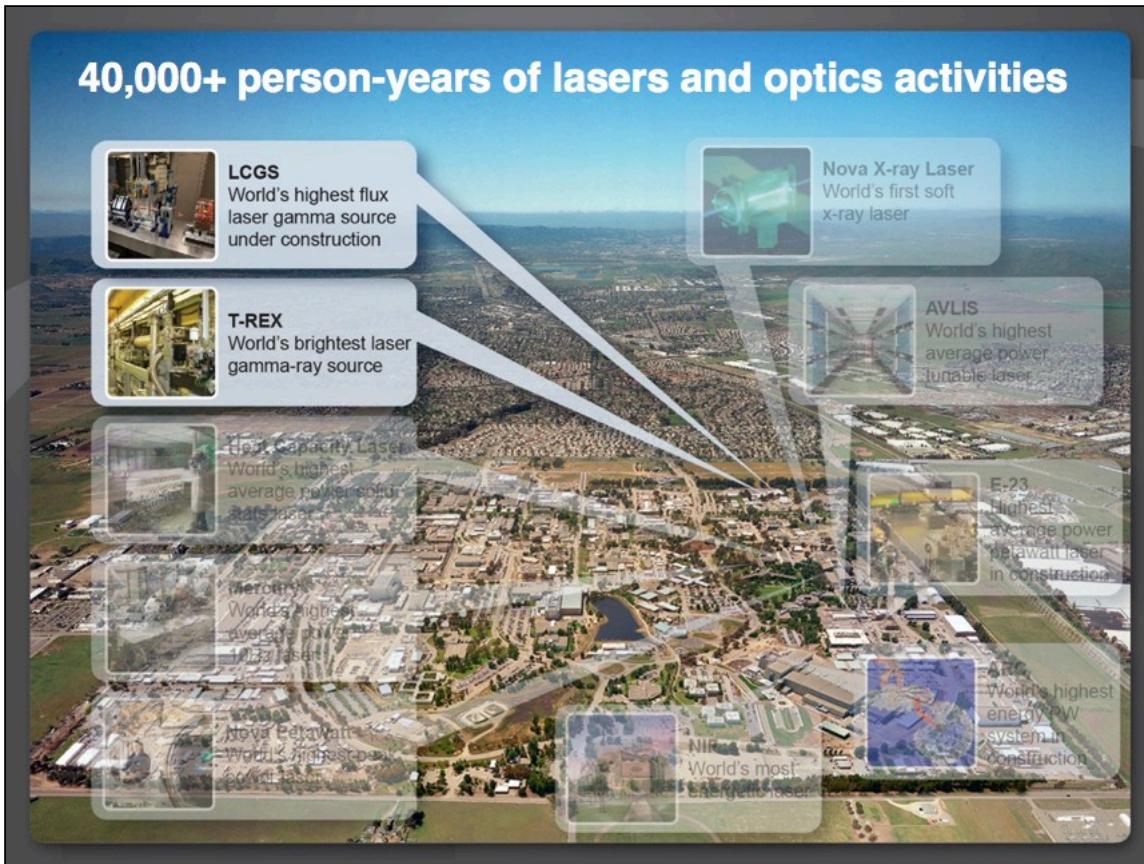
Dr. C. P. J. Barty  
Chief Technology Officer  
National Ignition Facility & Photon Science Directorate  
Lawrence Livermore National Laboratory  
Livermore, California  
April 10, 2014

This work performed under the auspices of the U.S. Department of Energy by Lawrence Livermore National Laboratory under Contract DE-AC52-07NA27344

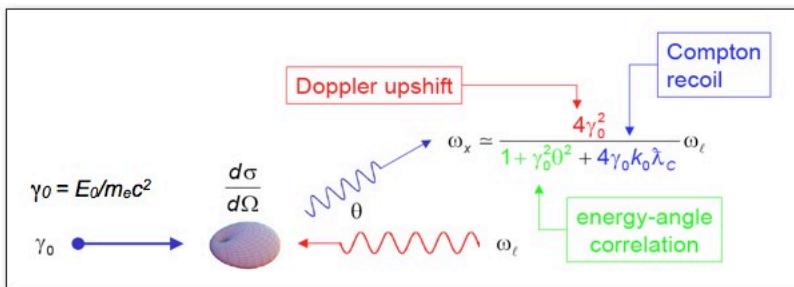
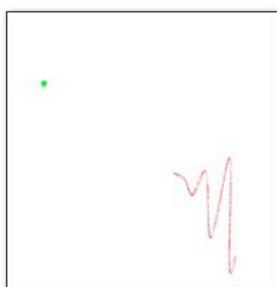




## 40,000+ person-years of lasers and optics activities



Laser Compton back scattering off of high energy electrons can produce tunable, high energy photons



Energy-momentum conservation yields  $\sim 4\gamma^2$  Doppler upshift

## 1965 First Light: 8 photons

PHYSICAL REVIEW

VOLUME 138, NUMBER 6B

21 JUNE 1965

### High-Energy Photons from Compton Scattering of Light on 6.0-GeV Electrons\*

CARLO BEMPORAD†, RICHARD H. MILBURN, AND NOBUYUKI TANAKA  
*Department of Physics, Tufts University, Medford, Massachusetts*

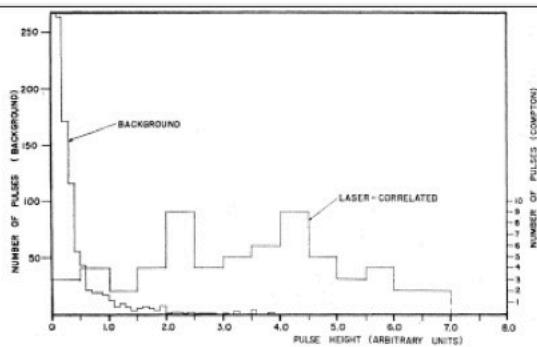
AND

MIRCEA FOTINO

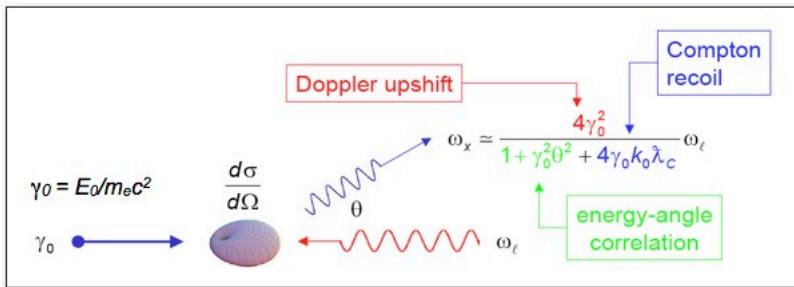
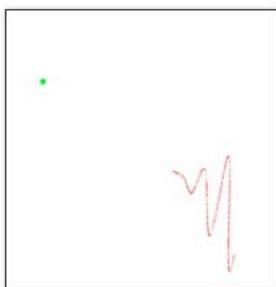
*Cambridge Electron Accelerator, Harvard University, Cambridge, Massachusetts*  
 (Received 28 January 1965; revised manuscript received 1 March 1965)

Compton scattering of optical photons on 6.0-GeV electrons has been observed at the Cambridge Electron Accelerator. A giant-pulsed ruby-laser burst of 0.2 J, impinging upon a 2-mA circulating electron current, was observed to yield about 8 scattered photons per pulse. These photons acquire, through a twofold Doppler shift, energies of hundred of MeV, and are expected to retain to a high degree the polarization of the laser beam. The observed yield is compatible with predictions based upon the theory of Compton scattering.

FIG. 6. Čerenkov-counter pulse-height distributions. Laser-correlated pulses are associated with second laser pulse. (2.3 units = 3 GeV).

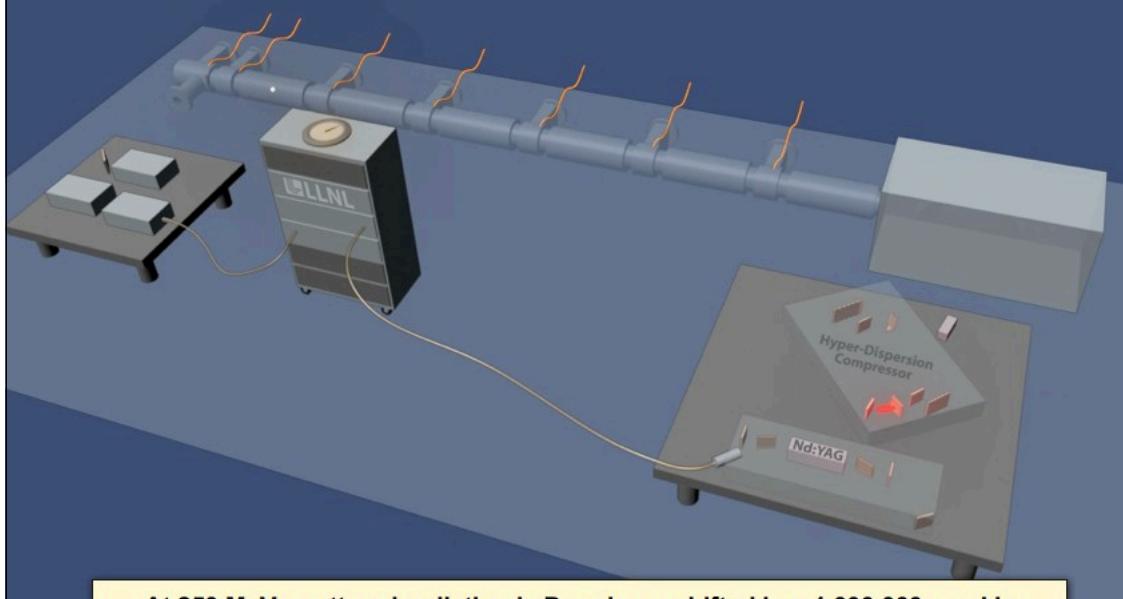


### Laser Compton back scattering off of high energy electrons can produce tunable, high energy photons



Thomson cross section is very small  $\sim 6 \times 10^{-25} \text{ cm}^2$

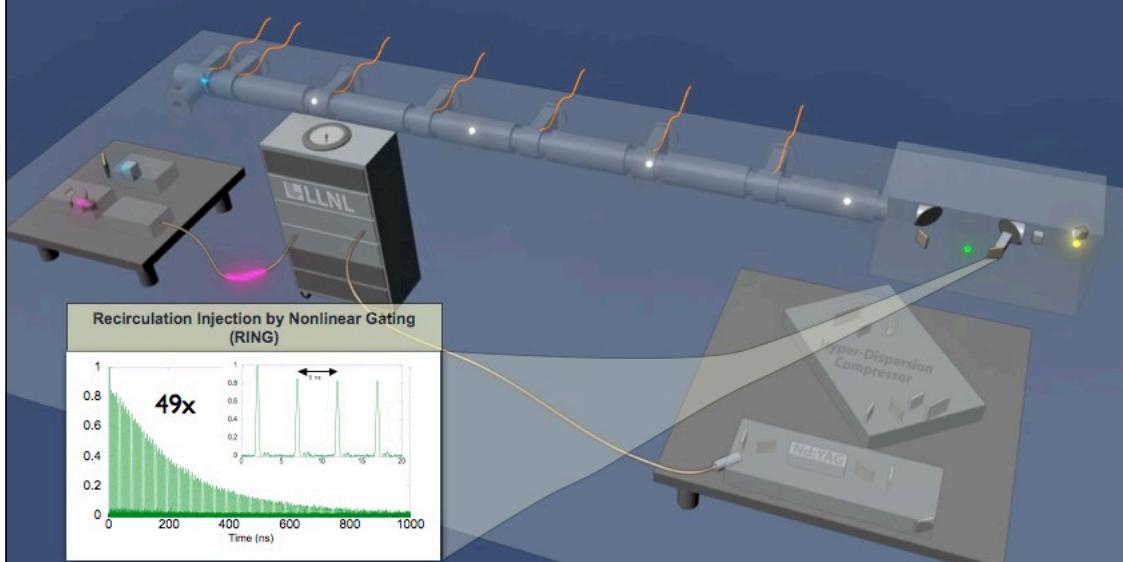
**High-flux, laser-Compton scattering arrangements aim to produce high photon & electron densities at a common focus**



**At 250 MeV, scattered radiation is Doppler upshifted by ~1,000,000x and is forwardly-directed in a narrow, polarized, tunable, laser-like, gamma beam**

US patent #8,068,522 Barty - Hyperdispersion Chirped Pulse Amplification and Compression

**High-flux, laser-Compton scattering arrangements perturb the laser pulse energy very little during the interaction**



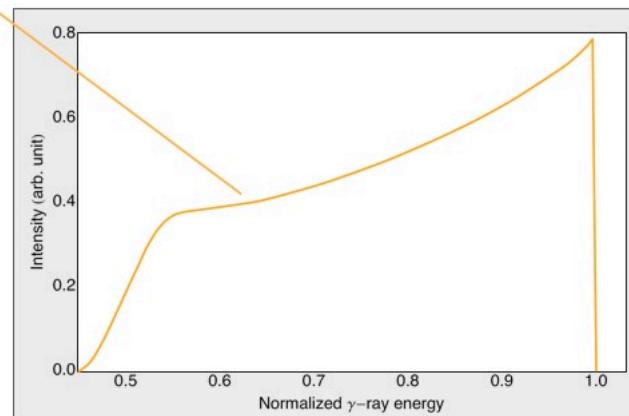
**Recirculation can give > 50x increase in Compton photon production for "free"  
RING positioning requirements are 10,000x less stringent than 'cavity' schemes**

Shverdin, M. Y., I. Jovanovic, V. A. Semenov, S. M. Betts, C. Brown, D. J. Gibson, R. M. Shuttlesworth, F. V. Hartemann, C. W. Siders and C. P. J. Barty., "High-power picosecond laser pulse recirculation." Optics Letters 35(13): 2224-2226. (2010)

**Overall Compton scattering is broadband, but it is highly angle correlated and is 'narrowband' on axis**



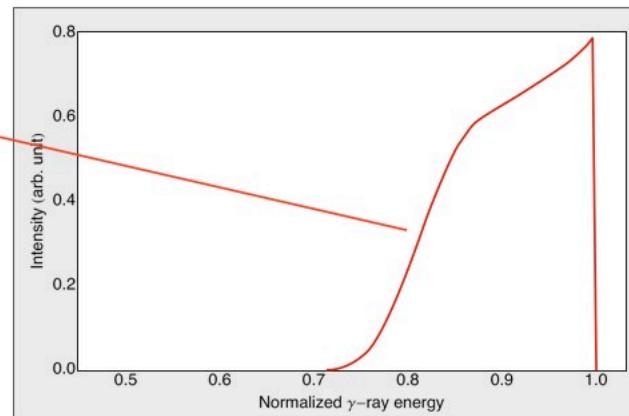
$$\Delta\Omega ; \pi \left( \frac{1}{\gamma} \right)^2 \text{ few mrad}$$



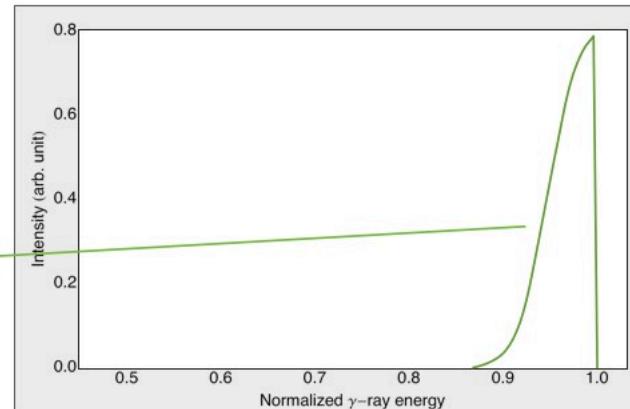
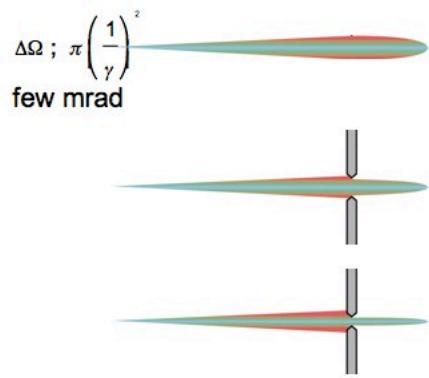
**Overall Compton scattering is broadband, but it is highly angle correlated and is 'narrowband' on axis**



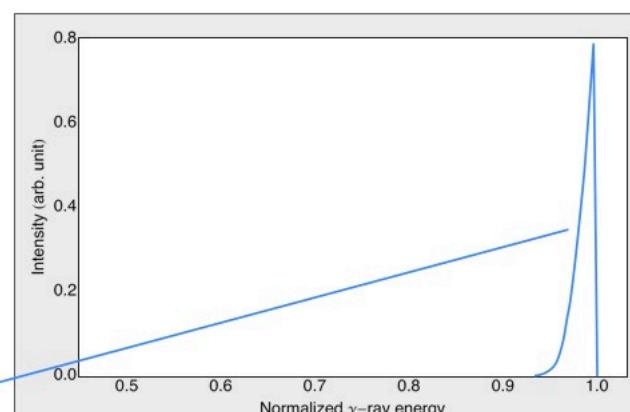
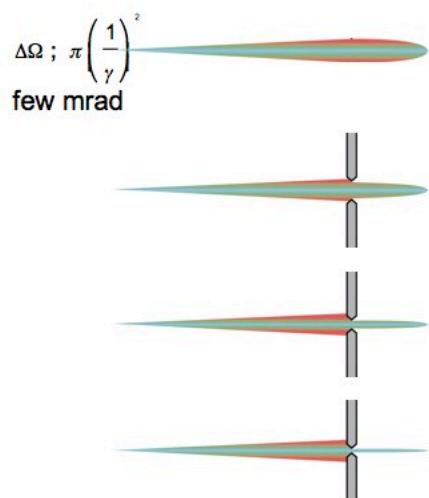
$$\Delta\Omega ; \pi \left( \frac{1}{\gamma} \right)^2 \text{ few mrad}$$



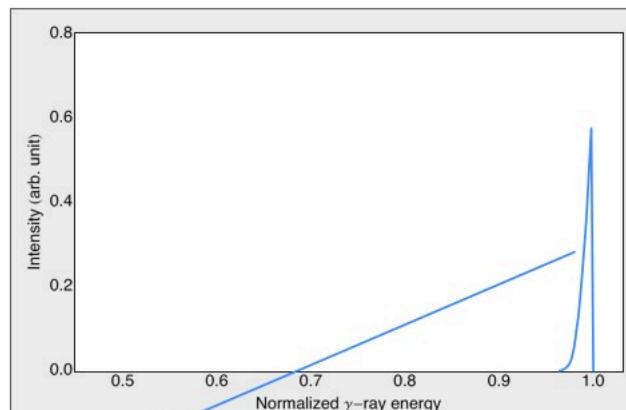
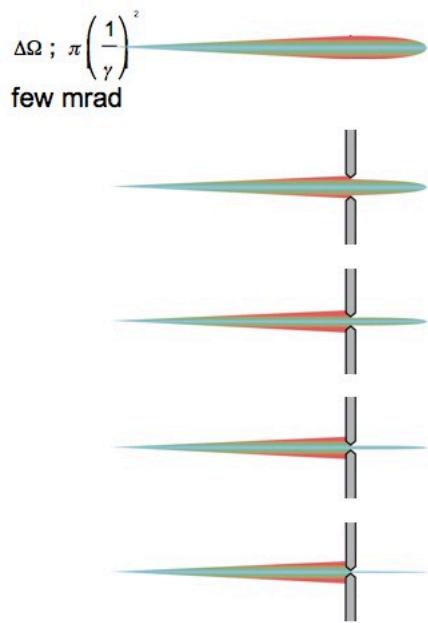
**Overall Compton scattering is broadband, but it is highly angle correlated and is 'narrowband' on axis**



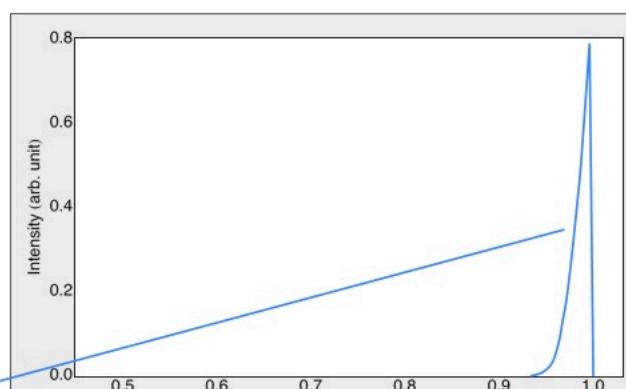
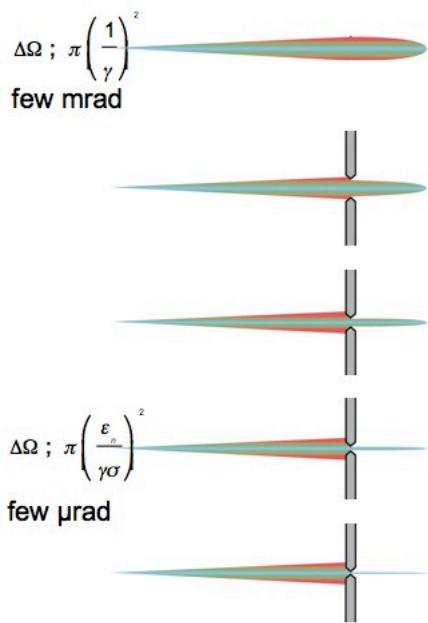
**Overall Compton scattering is broadband, but it is highly angle correlated and is 'narrowband' on axis**



**Overall Compton scattering is broadband, but it is highly angle correlated and is 'narrowband' on axis**

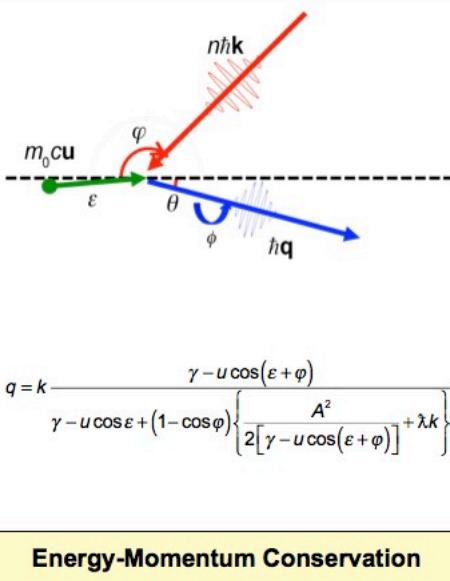


**Overall Compton scattering is broadband, but it is highly angle correlated and is 'narrowband' on axis**



**"Mono-Energetic Gamma-rays" - MEGa-rays**

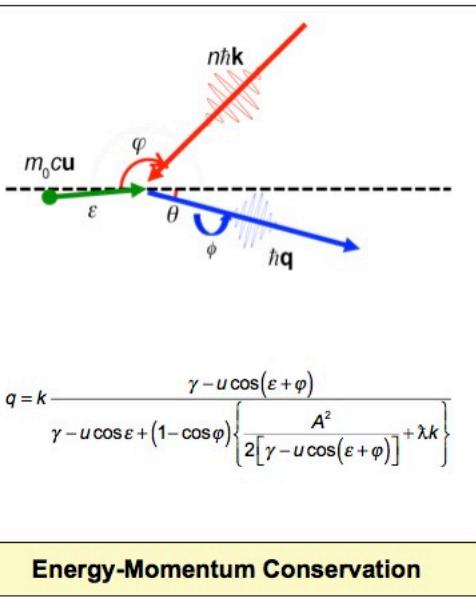
## Many factors contribute to the minimum possible bandwidth\*



- **Detection Aperture**  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta \theta^2$
- **Laser Bandwidth**  $\frac{\Delta q}{q} \approx \frac{\Delta k}{k}$
- **Laser Focal Spot**  $\frac{\Delta q}{q} \approx -\frac{1}{4} \Delta \varphi^2$
- **Nonlinear Radiation Pressure**  $\frac{\Delta q}{q} \approx -\frac{\Delta A^2}{1 + A^2}$
- **Electron Energy Spread**  $\frac{\Delta q}{q} \approx 2 \frac{\Delta \gamma}{\gamma}$
- **Electron Beam Emittance**  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta \varepsilon^2$

\* order-of-magnitude estimated contributions based on 2013 LLNL technology and optimized laser-Compton interaction geometry

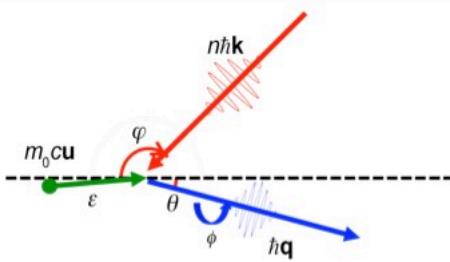
## Many factors contribute to the minimum possible bandwidth\*



- **Detection Aperture**  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta \theta^2$
- **Laser Bandwidth**  $\frac{\Delta q}{q} \approx \frac{\Delta k}{k}$  O( $10^{-4}$ )
- **Laser Focal Spot**  $\frac{\Delta q}{q} \approx -\frac{1}{4} \Delta \varphi^2$
- **Nonlinear Radiation Pressure**  $\frac{\Delta q}{q} \approx -\frac{\Delta A^2}{1 + A^2}$
- **Electron Energy Spread**  $\frac{\Delta q}{q} \approx 2 \frac{\Delta \gamma}{\gamma}$
- **Electron Beam Emittance**  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta \varepsilon^2$

\* order-of-magnitude estimated contributions based on 2013 LLNL technology and optimized laser-Compton interaction geometry

## Many factors contribute to the minimum possible bandwidth\*



$$q = k \frac{\gamma - u \cos(\varepsilon + \varphi)}{\gamma - u \cos \varepsilon + (1 - \cos \varphi) \left\{ \frac{A^2}{2[\gamma - u \cos(\varepsilon + \varphi)]} + \lambda k \right\}}$$

### Energy-Momentum Conservation

- Detection Aperture  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta \theta^2$

- Laser Bandwidth  $\sim 10 \text{ ps}$   $\frac{\Delta q}{q} \approx \frac{\Delta k}{k}$   $O(10^{-4})$

- Laser Focal Spot  $\sim 10 \text{ microns}$   $\frac{\Delta q}{q} \approx -\frac{1}{4} \Delta \varphi^2$   $O(10^{-4})$

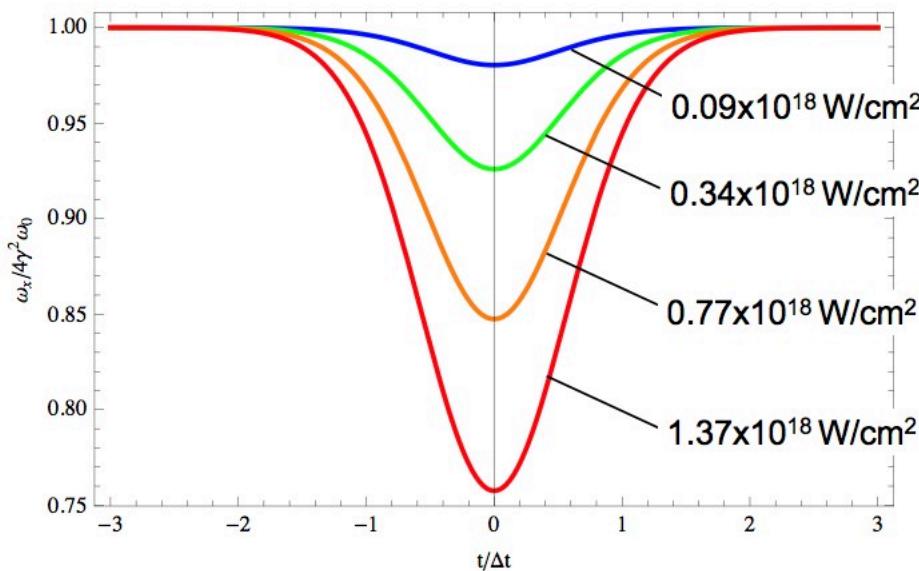
- Nonlinear Radiation Pressure  $\frac{\Delta q}{q} \approx -\frac{\Delta A^2}{1 + A^2}$

- Electron Energy Spread  $\frac{\Delta q}{q} \approx 2 \frac{\Delta \gamma}{\gamma}$

- Electron Beam Emittance  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta \varepsilon^2$

\* order-of-magnitude estimated contributions based on 2013 LLNL technology and optimized laser-Compton interaction geometry

## Radiation pressure can significantly change the velocity of the electron during scattering



# Nonlinear broadening at low intensity can also be an issue under certain long pulse conditions



PRL 105, 130801 (2010) PHYSICAL REVIEW LETTERS week ending 24 SEPTEMBER 2010

## Low-Intensity Nonlinear Spectral Effects in Compton Scattering

Frederic V. Hartemann, Félicie Albert, Craig W. Siders, and C. P. J. Barty  
Lawrence Livermore National Laboratory, Livermore, California, 94550, USA  
(Received 23 February 2010; published 24 September 2010)

Nonlinear effects are known to occur in Compton scattering light sources, when the laser normalized potential  $A$  approaches unity. In this Letter, it is shown that nonlinear spectral features can appear at arbitrarily low values of  $A$ , if the fractional bandwidth of the laser pulse  $\Delta\phi^{-1}$  is sufficiently small to satisfy  $\Delta^2\Delta\phi \ll 1$ . A three-dimensional analysis, based on a local plane wave, slow-varying envelope approximation, enables a study of these effects for realistic interactions between an electron beam and a laser pulse, and their influence on high-precision Compton scattering light sources.

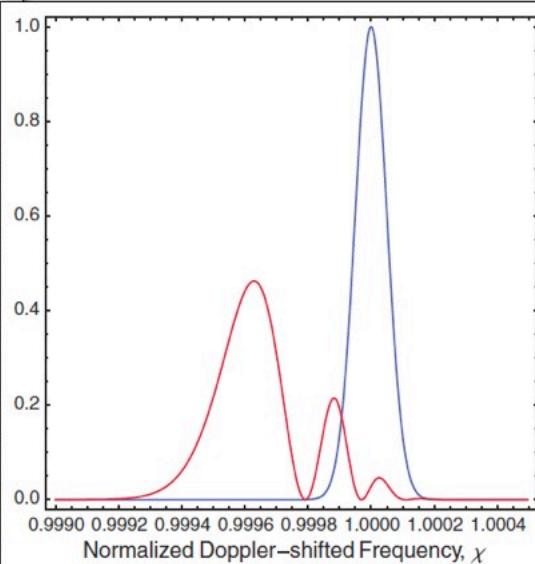
DOI: 10.1103/PhysRevLett.105.130801

PCNS numbers: 0785.-m, 41.60.Cc, 42.72.-g

Rapid advances in terawatt-class laser technology [1] and high-brightness, high-gradient electron accelerators [2] are enabling the development of a new type of light source based on Compton scattering [3], where relativistic electrons interact with a coherent photon field to generate bright, ultrafast, tunable x rays and γ rays [4,5]. These compact sources are a natural complement to larger-scale 3rd and 4th generation light sources [6], and provide a means to generate MeV-scale photons with unprecedented spectral brightness.

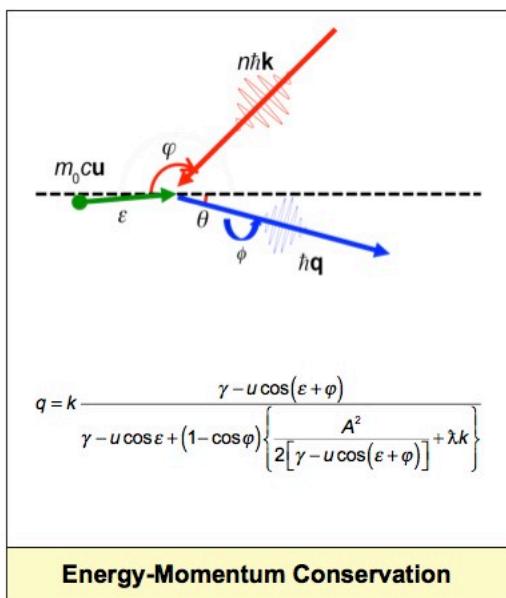
Among other important features, such as wide tunability and ultrafast pulse capability, Compton scattering x-ray and γ-ray sources offer the potential of generating highly-correlated, narrow band radiation in a very small solid angle. This characteristic is desirable for a number of applications, including nuclear resonance fluorescence [7] or protein crystallography [8]. Therefore, the focus of this work is the physical origin of spectral broadening mechanisms in Compton scattering, with a special emphasis on nonlinear effects and recoil, and their influence on the performance of high-precision Compton scattering light sources.

In this Letter, three novel results are presented: (1) A covariant form of the radiation formula is given, including a quantum correction term shown to yield the proper recoil for the interaction, along with a gauge invariant, covariant definition of the 4-polarization [9]. (2) We demonstrate that, while nonlinear effects are known to occur in light sources, when the laser potential  $A$  or normalized 4-potential  $A$  approaches unity, nonlinear spectral features can also appear at arbitrarily low values of  $A$ , if the fractional bandwidth of the laser pulse  $\Delta\phi^{-1}$  is sufficiently small and satisfies the condition  $\Delta^2\Delta\phi \ll 1$ . (3) A fully three-dimensional (3D) analysis of nonlinear effects in the



Hartemann, F. V., F. Albert, C. W. Siders and C. P. J. Barty. "Low-intensity nonlinear spectral effects in Compton scattering." *Physical Review Letters* 105(13): 130801. (2010)

# Many factors contribute to the minimum possible bandwidth\*



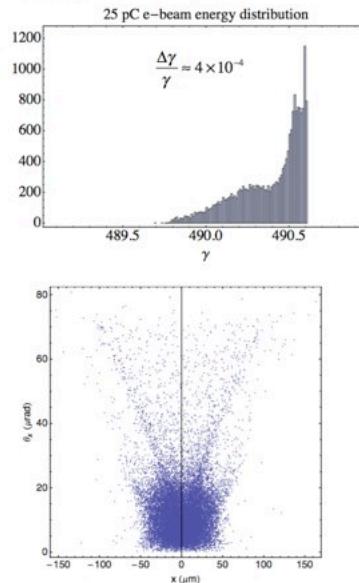
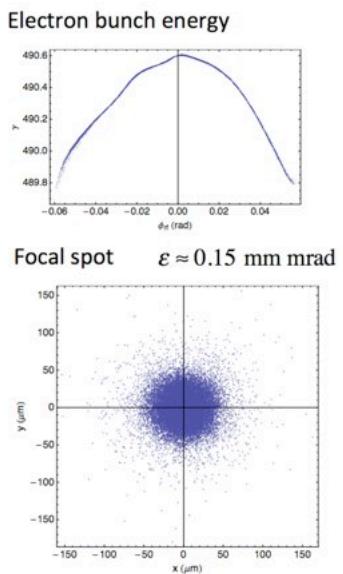
- **Detection Aperture**  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta\theta^2$
- **Laser Bandwidth**  $\sim 10 \text{ ps}$   $\frac{\Delta q}{q} \approx \frac{\Delta K}{K}$   $O(10^{-4})$
- **Laser Focal Spot**  $\sim 10 \text{ microns}$   $\frac{\Delta q}{q} \approx -\frac{1}{4} \Delta\varphi^2$   $O(10^{-4})$
- **Nonlinear Radiation Pressure**  $\frac{\Delta q}{q} \approx -\frac{\Delta A^2}{1 + A^2}$   $< 10^{-4}$
- **Electron Energy Spread**  $\frac{\Delta q}{q} \approx 2 \frac{\Delta\gamma}{\gamma}$
- **Electron Beam Emittance**  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta\epsilon^2$

\* order-of-magnitude estimated contributions based on 2013 LLNL technology and optimized laser-Compton interaction geometry

**Modern, high gradient accelerators can create very small emittance and energy spread at low charge**

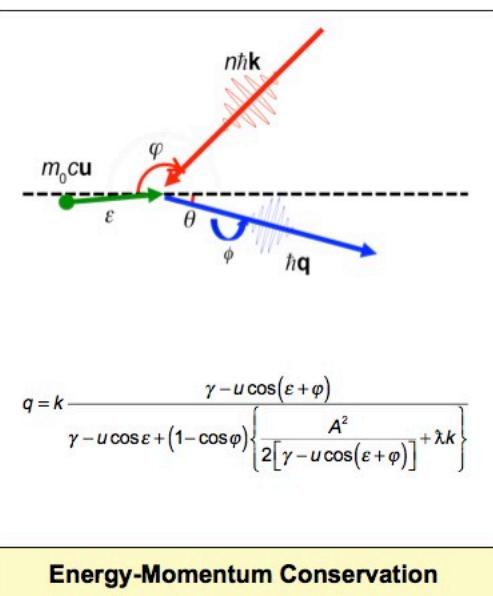


### 25 pC electron beam phase space



Marsh, R. A., F. Albert, S. G. Anderson, G. Beer, T. S. Chu, R. R. Cross, G. A. Deis, C. A. Ebbers, D. J. Gibson, T. L. Houck, F. V. Hartemann, C. P. J. Barty, et al. "Modeling and design of an X-band RF photoinjector." *Physical Review Special Topics: Accelerators and Beams* 15(10): 102001 (2012)

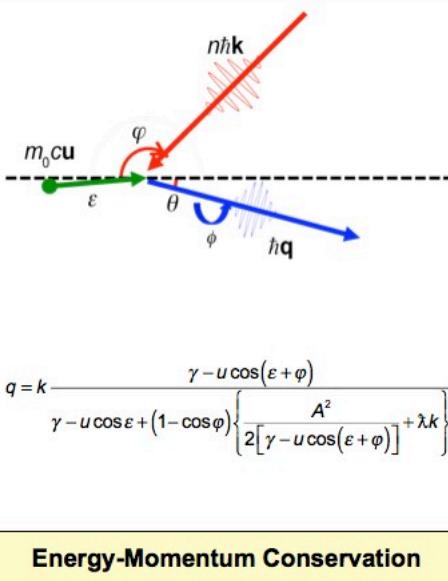
**Many factors contribute to the minimum possible bandwidth\***



- Detection Aperture  $\frac{\Delta q}{q} \simeq -\gamma^2 \Delta \theta^2$
- Laser Bandwidth  $\sim 10 \text{ ps}$   $\frac{\Delta q}{q} \simeq \frac{\Delta \lambda}{\lambda} \propto O(10^{-4})$
- Laser Focal Spot  $\sim 10 \text{ microns}$   $\frac{\Delta q}{q} \simeq -\frac{1}{4} \Delta \varphi^2 \propto O(10^{-4})$
- Nonlinear Radiation Pressure  $\frac{\Delta q}{q} \simeq -\frac{\Delta A^2}{1 + A^2} < 10^{-4}$
- Electron Energy Spread  $\frac{\Delta q}{q} \simeq 2 \frac{\Delta \gamma}{\gamma}$
- Electron Beam Emittance  $\frac{\Delta q}{q} \simeq -\gamma^2 \Delta \varepsilon^2$

\* order-of-magnitude estimated contributions based on 2013 LLNL technology and optimized laser-Compton interaction geometry

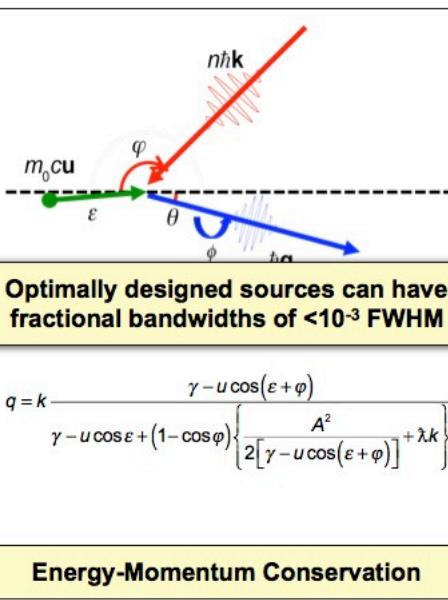
## Many factors contribute to the minimum possible bandwidth\*



- **Detection Aperture**  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta \theta^2 < 10^{-3}$
- **Laser Bandwidth**  $\sim 10 \text{ ps}$   $\frac{\Delta q}{q} \approx \frac{\Delta k}{k} O(10^{-4})$
- **Laser Focal Spot**  $\sim 10 \text{ microns}$   $\frac{\Delta q}{q} \approx -\frac{1}{4} \Delta \varphi^2 O(10^{-4})$
- **Nonlinear Radiation Pressure**  $\frac{\Delta q}{q} \approx -\frac{\Delta A^2}{1 + A^2} < 10^{-4}$
- **Electron Energy Spread**  $\frac{\Delta q}{q} \approx 2 \frac{\Delta \gamma}{\gamma} < 10^{-3}$
- **Electron Beam Emittance**  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta \varepsilon^2 < 10^{-3}$

\* order-of-magnitude estimated contributions based on 2013 LLNL technology and optimized laser-Compton interaction geometry

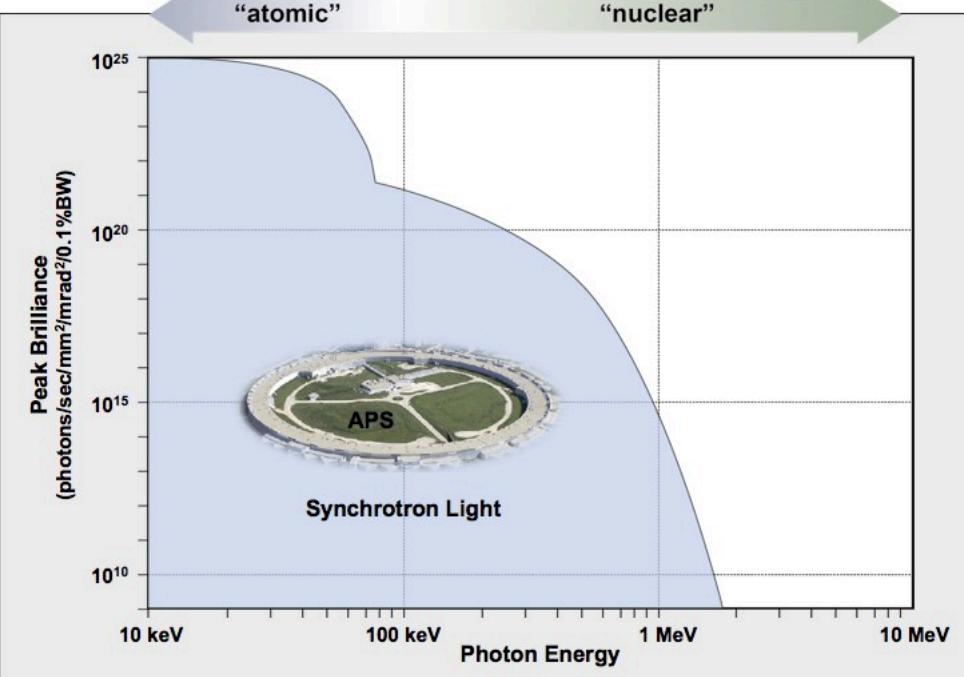
## Many factors contribute to the minimum possible bandwidth\*



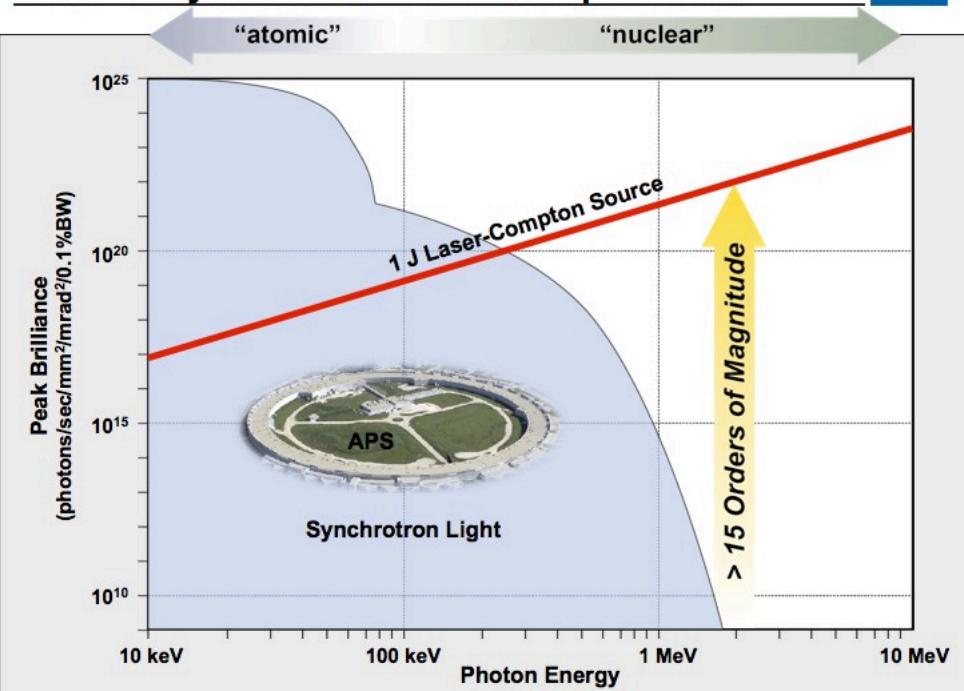
- **Detection Aperture**  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta \theta^2 < 10^{-3}$
- **Laser Bandwidth**  $\sim 10 \text{ ps}$   $\frac{\Delta q}{q} \approx \frac{\Delta k}{k} O(10^{-4})$
- **Laser Focal Spot**  $\sim 10 \text{ microns}$   $\frac{\Delta q}{q} \approx -\frac{1}{4} \Delta \varphi^2 O(10^{-4})$
- **Nonlinear Radiation Pressure**  $\frac{\Delta q}{q} \approx -\frac{\Delta A^2}{1 + A^2} < 10^{-4}$
- **Electron Energy Spread**  $\frac{\Delta q}{q} \approx 2 \frac{\Delta \gamma}{\gamma} < 10^{-3}$
- **Electron Beam Emittance**  $\frac{\Delta q}{q} \approx -\gamma^2 \Delta \varepsilon^2 < 10^{-3}$

\* order-of-magnitude estimated contributions based on 2013 LLNL technology and optimized laser-Compton interaction geometry

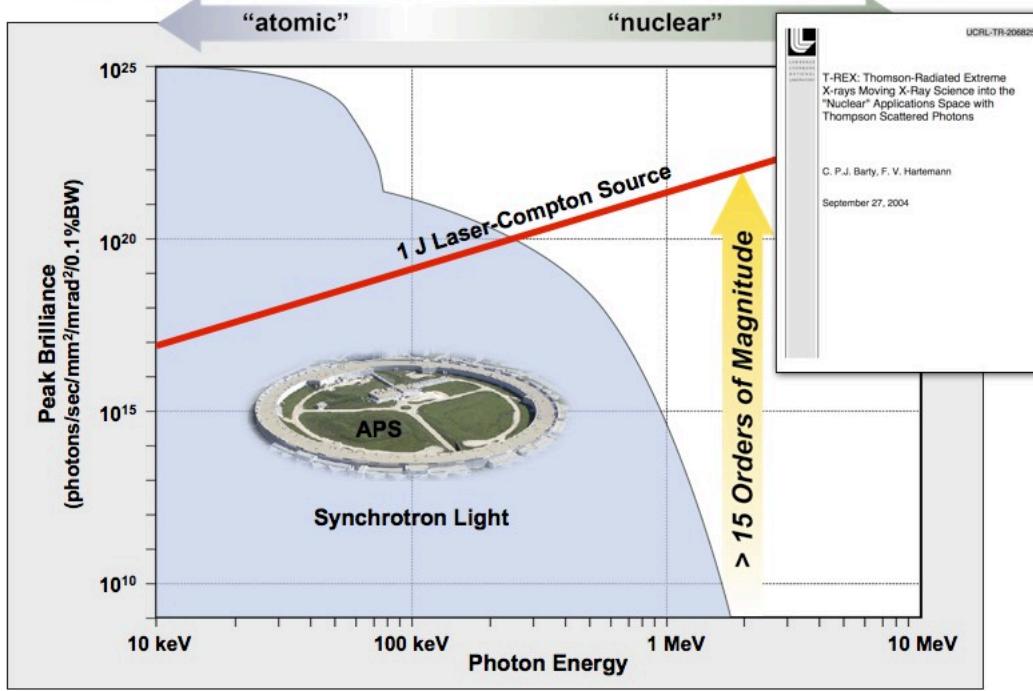
**Optimized laser-Compton sources enable “lab-scale” x-ray science and “nuclear” photonics**



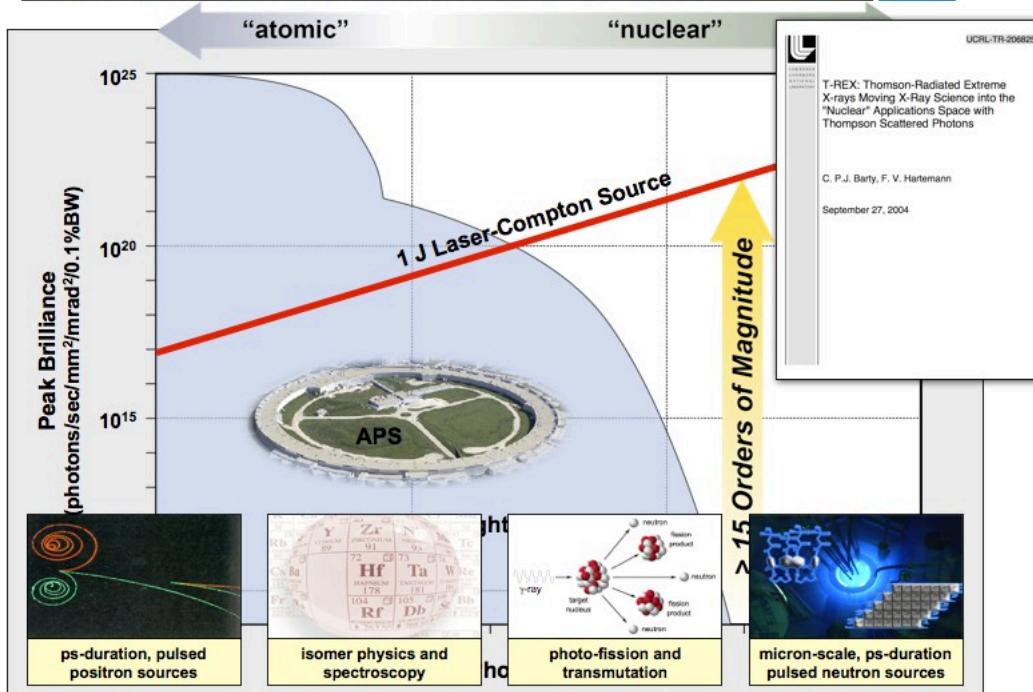
**Optimized laser-Compton sources enable “lab-scale” x-ray science and “nuclear” photonics**



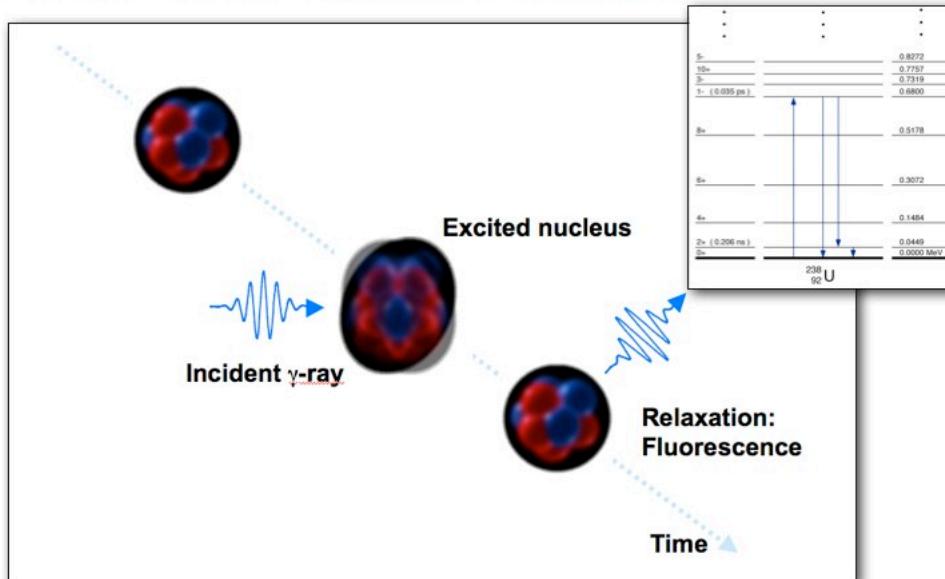
## Optimized laser-Compton sources enable “lab-scale” x-ray science and “nuclear” photonics



## Optimized laser-Compton sources enable “lab-scale” x-ray science and “nuclear” photonics

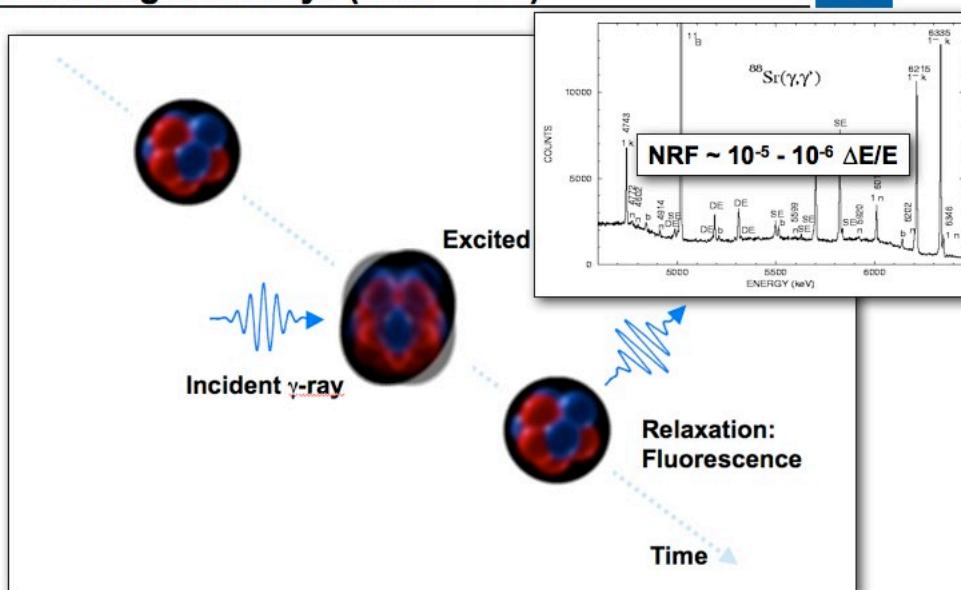


**Gamma-ray absorption & radiation by the nucleus is an “isotope-specific” signature of the material**



**Nuclear Resonance Fluorescence (NRF) is analogous to atomic resonance fluorescence but depends upon the number of protons AND the number of neutrons in the nucleus**

**Selective excitation of NRF is possible with narrow bandwidth gamma-rays ( $\Delta E/E \sim 10^{-3}$ )**

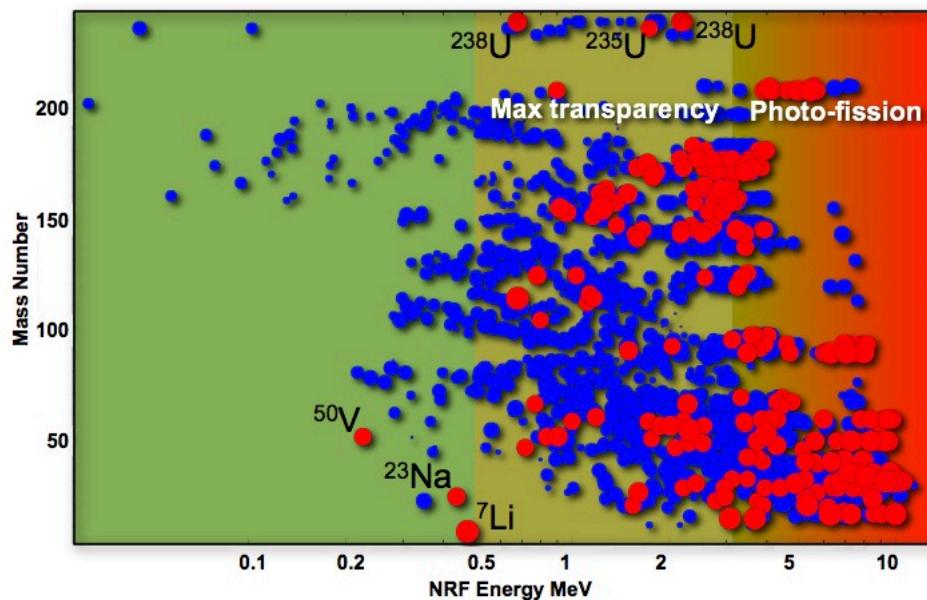


**Nuclear Resonance Fluorescence (NRF) is analogous to atomic resonance fluorescence but depends upon the number of protons AND the number of neutrons in the nucleus**

**NRF transitions are common and many have cross sections larger than the atomic background**

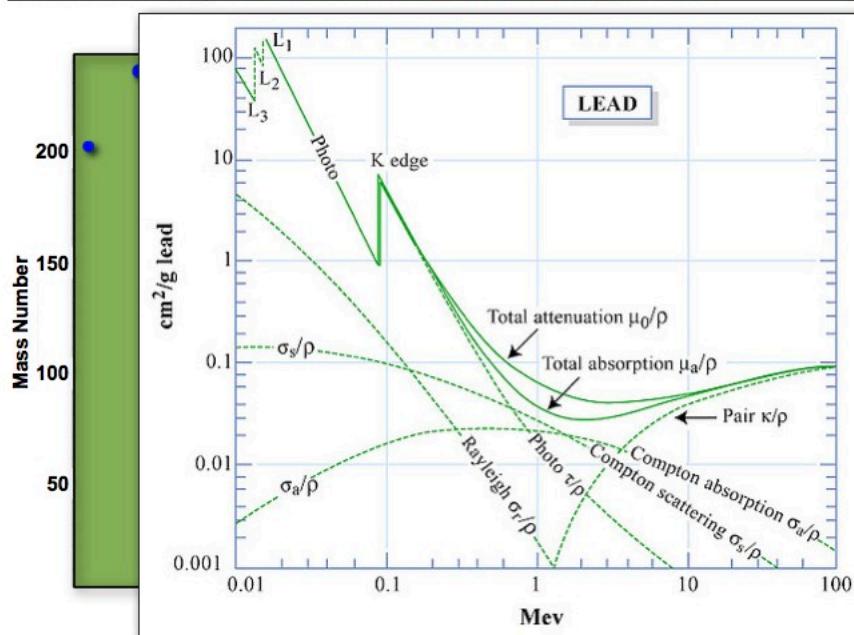


NRF/Atomic



Gammas in the 1 MeV to 3 MeV range are both highly penetrating and non-activating

**NRF transitions are common and many have cross sections larger than the atomic background**



Gammas in the 1 MeV to 3 MeV range are both highly penetrating and non-activating

## Bright, mono-energetic, gamma-ray sources enable new possibilities: "Inverse Density" radiography



- X-ray absorption is proportional to electron density/atomic number
- IF the x-ray can penetrate the high-Z material, it is not stopped by the low-Z material
- Low-Z materials are effectively shielded by dense, high-Z material



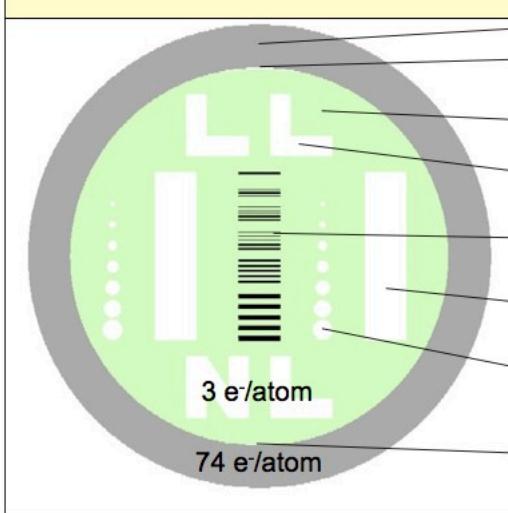
Precision imaging of low density material features inside of high density components is not feasible with conventional x-ray radiography

Monte Carlo simulations using COG



## Imaging simulations (shielded low-Z material)

Shielded low-Z NRF test object (a)



2.2 cm Ø W shell (2 mm wall)

"Crown erosions" @ T & B  
— 125 µm (T), 50 µm (B)

Natural LiH salt core (92.5% <sup>7</sup>Li)

1 mm thick letters (void)

Resolution grids (absorber)  
— 40 - 4 lp / mm (T to B)

Step wedges (void)  
— 125 - 1000 µm thick (B to T)

Spherical voids  
— 125 - 1000 µm Ø (T to B)  
"Crown erosions" @ T & B  
— 125 µm (T), 50 µm (B)



## Imaging simulations (shielded low-Z material)

Shielded low-Z NRF test object (a)



2.2 cm Ø W shell (2 mm wall)

“Crown erosions” @ T & B

— 125 µm (T), 50 µm (B)

Natural LiH salt core (92.5%  $^7\text{Li}$ )

1 mm thick letters (void)

Resolution grids (absorber)

— 40 - 4 lp / mm (T to B)

Step wedges (void)

— 125 - 1000 µm thick (B to T)

Spherical voids

— 125 - 1000 µm Ø (T to B)

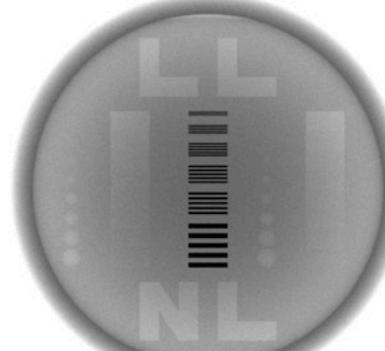
“Crown erosions” @ T & B

— 125 µm (T), 50 µm (B)



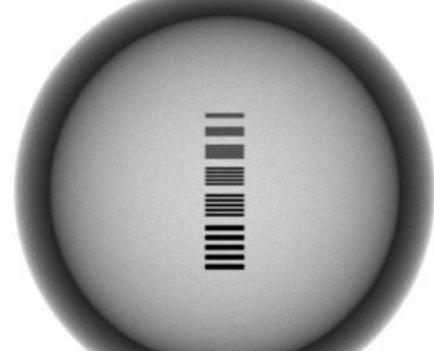
## Simulation of 0.478 MeV NRF image vs. 9 MeV e-Brem x-ray image (normalized\*)

0.478 MeV NRF Image Simulation  
(Enhanced  $^7\text{Li}$  Density)



High-Z shell:  
EDep ~ 0.117 MeV/ $\gamma$   
RDose ~ 1.060E-07  $\mu\text{Sv}/\gamma$

9 MeV e-Brem X-Ray  
Image Simulation



High-Z shell:  
EDep ~ 0.215 MeV/ $\gamma$   
RDose ~ 4.276E-07  $\mu\text{Sv}/\gamma$

\* The images have been normalized such that the open-field intensity in each case is ≈ 1 (arbitrary units)

Potential NRF-based Applications of Bright Gamma Sources are Numerous



**HEU Grand Challenge**  
*detection of shielded material*



**Nuclear Fuel Assay**  
*100 parts per million per isotope*



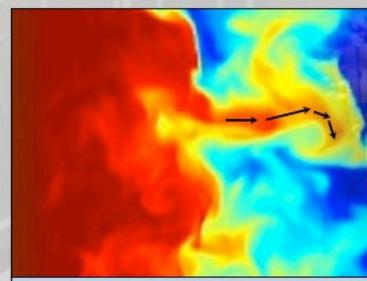
**Waste Imaging & Assay**  
*non-invasive content certification*



**Industrial NDE**  
*micron-scale & isotope specific*



**Medical Imaging**  
*low density & isotope specific*



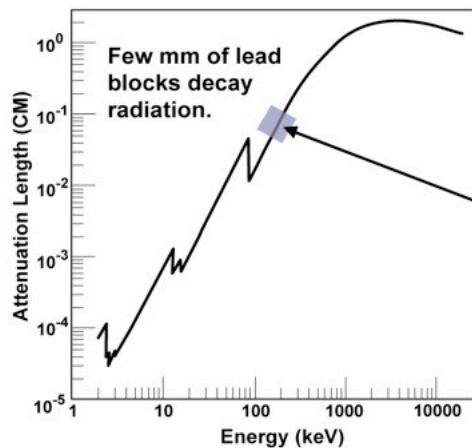
**Dense Plasma Science**  
*isotope mass, position & velocity*

US patent #7,564,241 Barty, Hartemann, McNabb & Pruet - detection, assay and imaging with laser-Compton gamma-rays

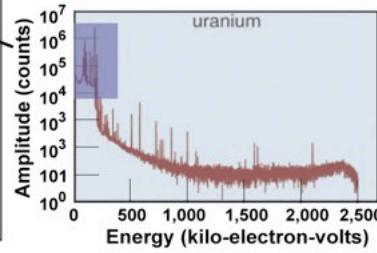
**Finding shielded Highly Enriched Uranium (HEU) is a grand challenge**



**Lead Attenuation**



Long half-life:  $7 \times 10^8$  y  
Low-energy signal: 186 keV

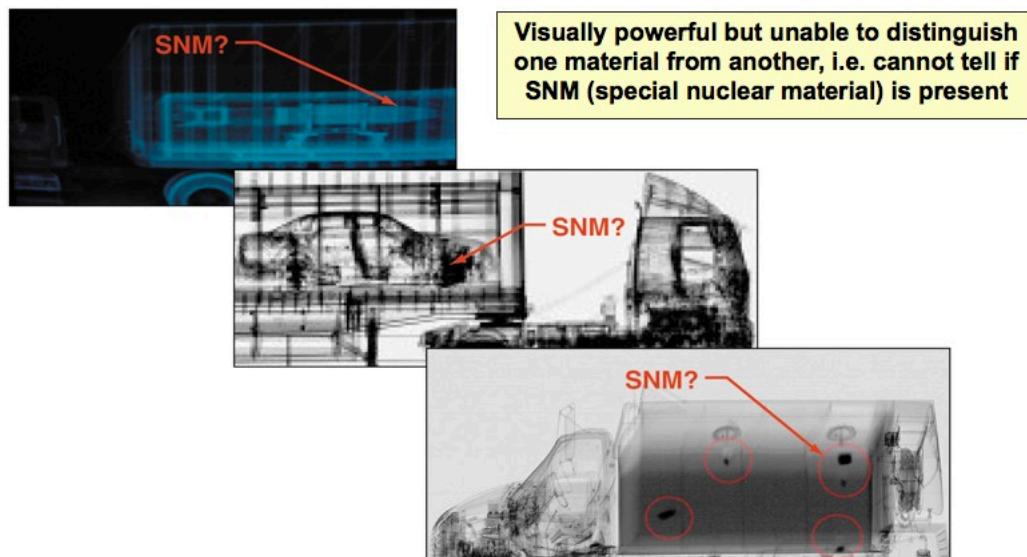


**Passive detection of decay signature is easily defeated by mm's of lead shielding**

- Worldwide sea-faring cargo container traffic: 48,000,000/year
- 3 containers shipped/received every 2 seconds!
- No present method reliably detects weapons grade Uranium

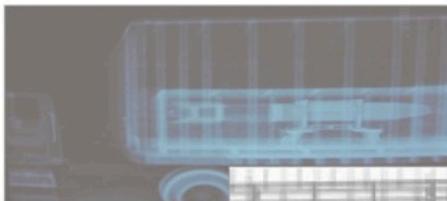


**Conventional MeV-range radiographic inspection  
can image through thick objects**



Visually powerful but unable to distinguish one material from another, i.e. cannot tell if SNM (special nuclear material) is present

## MEGa-ray beam absorption could be used to rapidly & safely distinguish SNM from background materials



Using MEGA-ray technologies currently under development, identification of a ~5mm thick piece of 235U would take << 1s



Numerical analysis suggests that active (but not "activating") SNM detection could be accomplished with MEGA-rays at doses that are 100x below allowable human limits

US patent #7,564,241 Barty, Hartemann, McNabb & Pruet - detection, assay and imaging with MEGA-rays

## MEGa-ray beam absorption could be used to rapidly & safely distinguish SNM from background materials



JOURNAL OF APPLIED PHYSICS 99, 123102 (2006)

Detecting clandestine material with nuclear resonance fluorescence  
J. Pruet,<sup>a</sup> D. P. McNabb,<sup>b</sup> C. A. Hagmann,<sup>c</sup> F. V. Hartemann,<sup>c</sup> and C. P. J. Barty  
<sup>a</sup>Los Alamos National Laboratory, 1000 E. Arroyo Laramie, California 93740  
<sup>b</sup>Lawrence Livermore National Laboratory, 7000 East Avenue, Livermore, California 94550  
<sup>c</sup>Battelle Seattle Research Center, 2000 University Street, Seattle, Washington 98103 (Received 19 March 2006)

We study the performance of a class of interrogation systems that utilize nuclear resonance fluorescence (NRF) to detect specific isotopes. In these systems the presence of a particular nuclide is inferred by observing the preferential attenuation of photons that strongly excite an

### MEGa-ray Bandwidth Matters

resonance transition. The performance of NRF-based interrogation systems may provide insights for efforts in light source development for applications related to national security and industry. © 2006 American Institute of Physics. [DOI: 10.1063/1200505]

#### I. INTRODUCTION

Conventional X-ray systems have found countless practical applications. These operate on the simple principle that some materials are more transparent to X-rays than others will preferentially attenuate and scatter photons. All of the important processes for X-rays are atomic in nature. This implies a potential for X-rays to be highly sensitive to specific isotopic content. Further photons with extremely large absorption cross sections and we cannot penetrate objects with large axial densities. Increasing the photon energy to higher values increases the probability that an electron's electrons results in better penetration. At these energies, though, the critical depth of a material is merely only a function of the atomic number. As a result, the X-ray's dose and probe sensitivity to atomic composition is lost until the threshold of the Compton edge is reached.

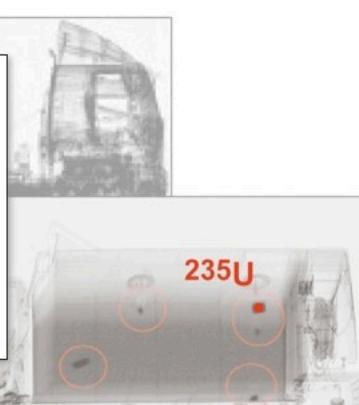
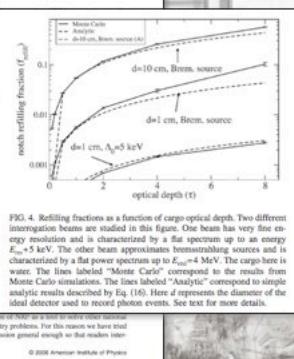
Here we consider the capabilities and performance of systems which can be viewed as the nuclear analog of conventional X-ray systems. These systems utilize a beam of impinging photons to induce and observe nuclear resonance fluorescence (NRF). As we discuss, they are capable of rapidly and accurately detecting specific isotopes, such as the isotopic composition of large well-shaped objects. For this reason, these systems have the potential to be used for detection relating to areas that include nonproliferation, waste identification, material characterization, and homeland security.

Most of the promise associated with NRF-based detection is a consequence of the unique nuclear resonance fluorescence, which refers to the absorption and reemission of photons by a nucleus. It is well known that the pattern of emission of a nucleus is unique. This provides a means for a unique fingerprint identifying that nucleus. In many cases even just a single well-measured transition is enough to iden-

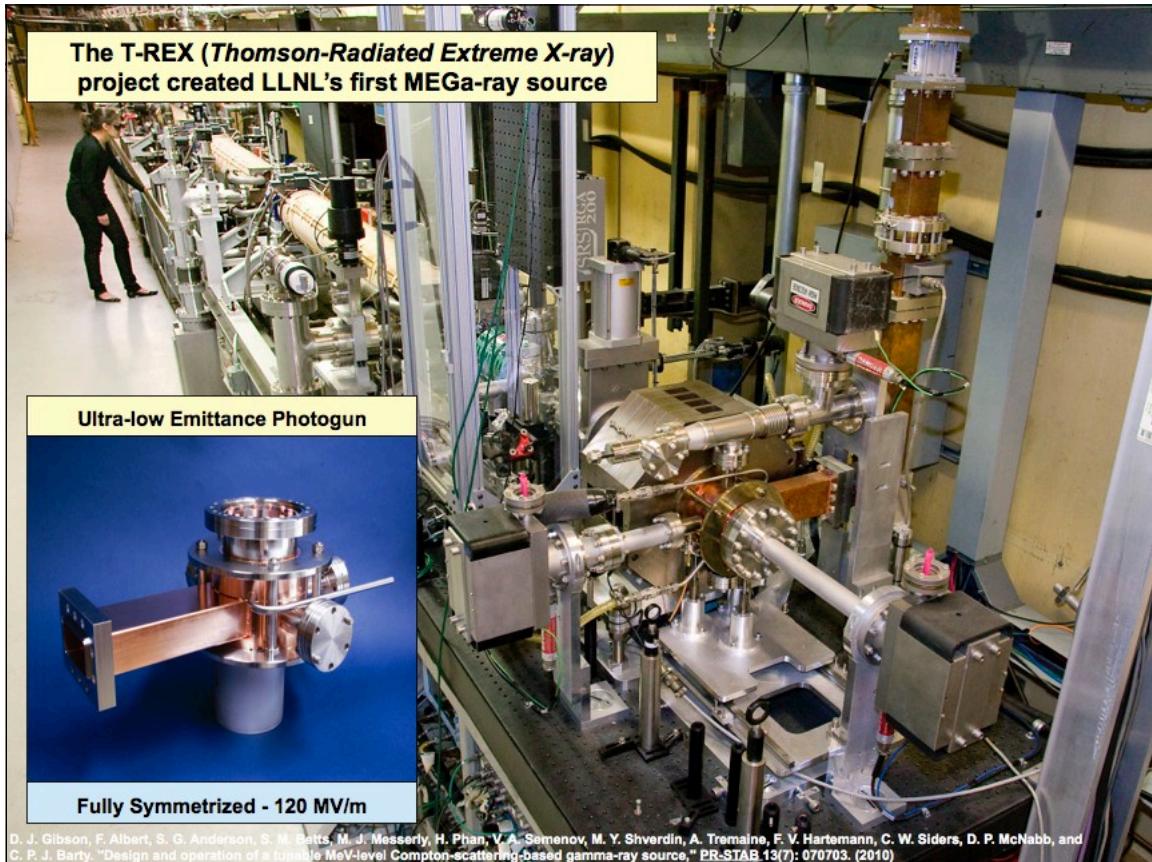
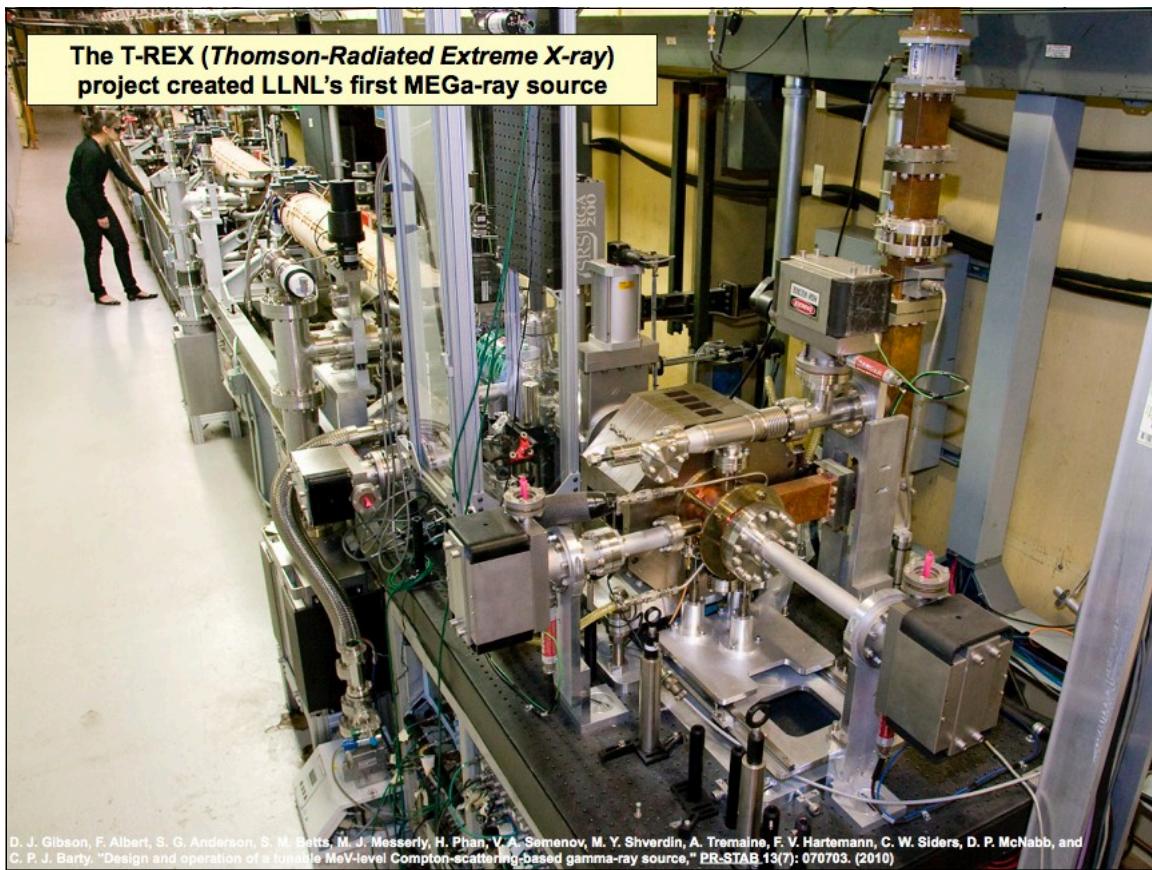
tify an isotope used to get a quick answer. However, there is also a need to measure the signal at relatively low energy to reduce exciting these transitions. This requires a very large and detailed view of cross sections and spectra for many isotopes (e.g., Ref. 2). Nuclear resonance fluorescence (NRF) is a type of powerful tool for interrogating materials by observing nuclear transitions within an object. The most common application for interrogating materials is to determine the isotopic composition of large well-shaped objects. For this reason, these systems have the potential to be used for detection relating to areas that include nonproliferation, waste identification, material characterization, and homeland security.

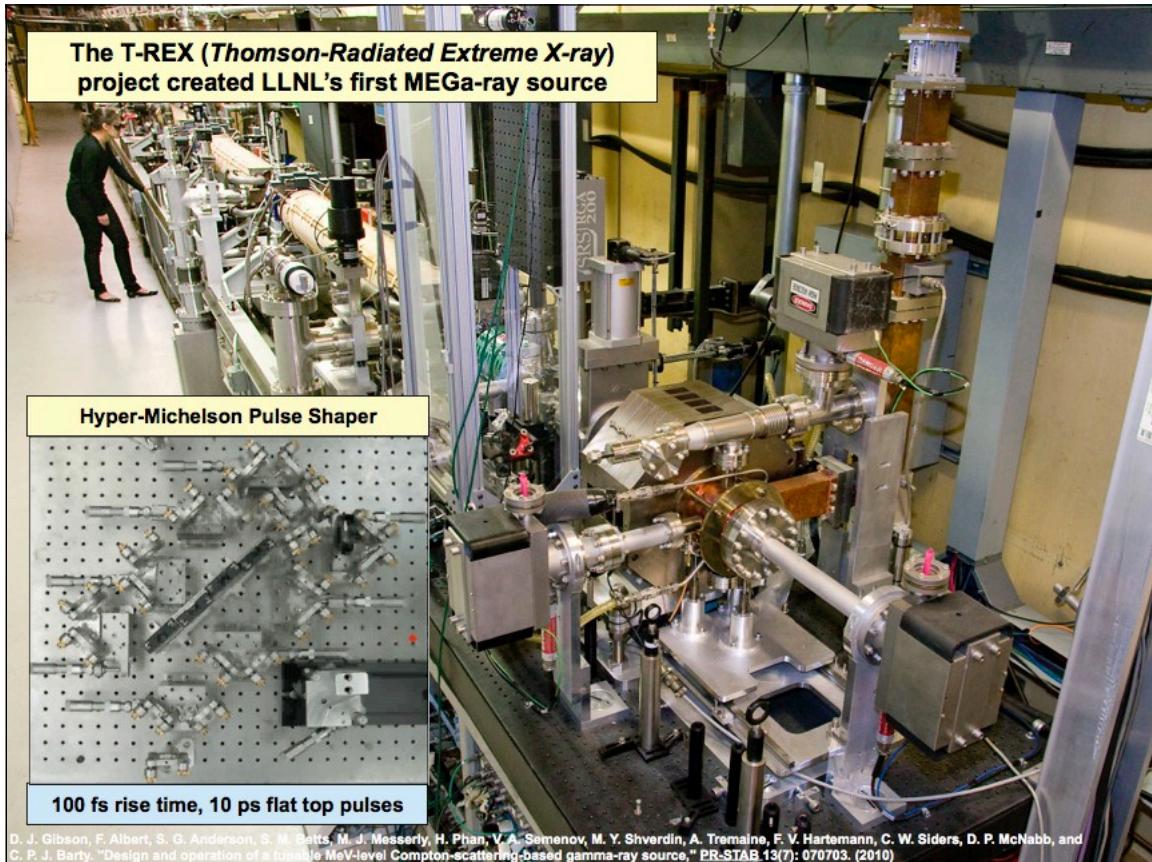
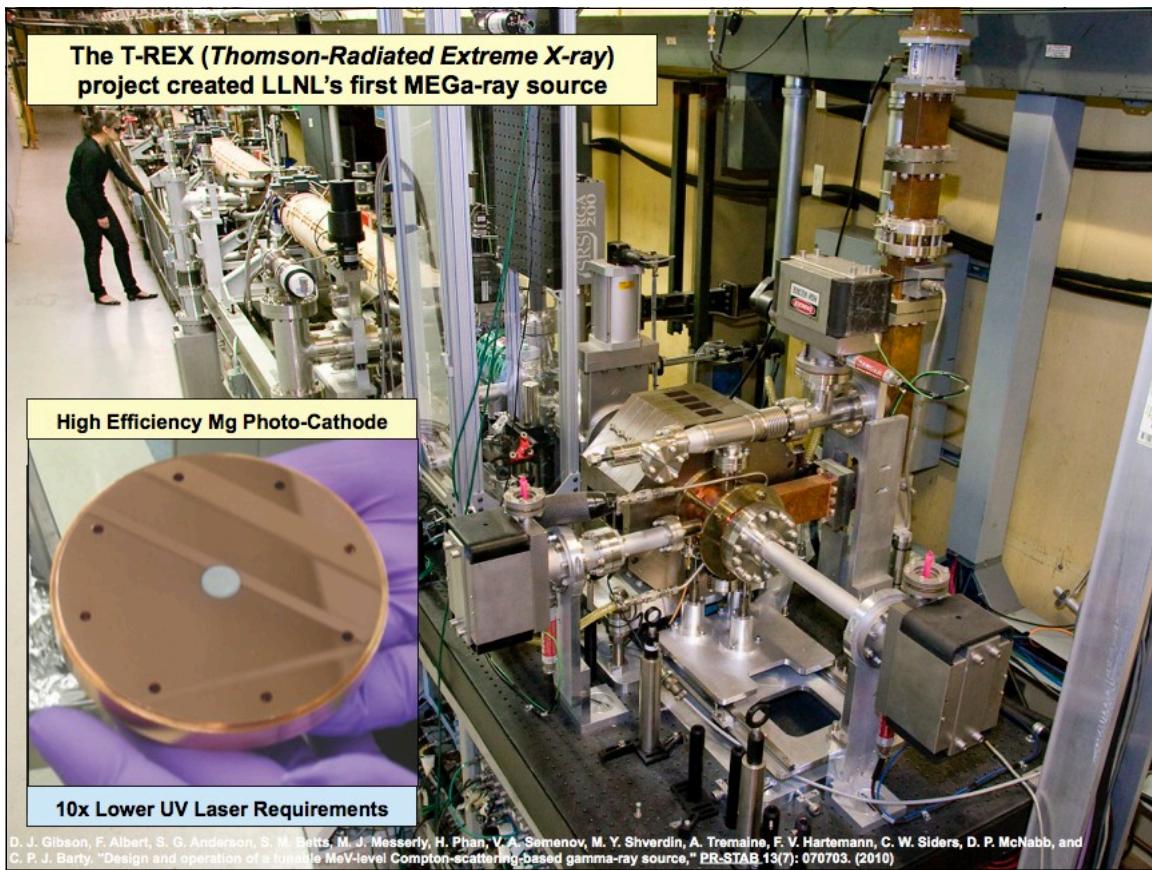
The use of NRF-based detection for identification purposes is currently limited to the use of Monte Carlo simulations. The main problem is that the ideal detector used to record photon events. See text for more details.

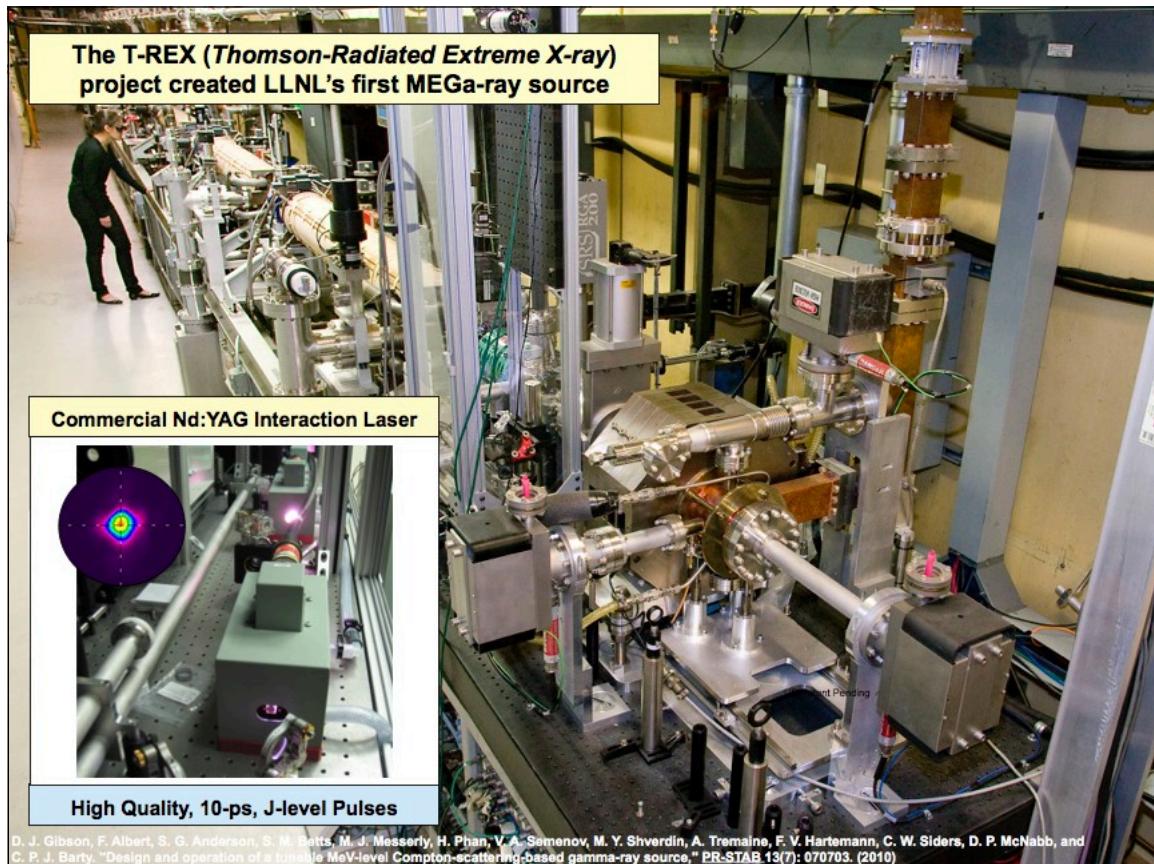
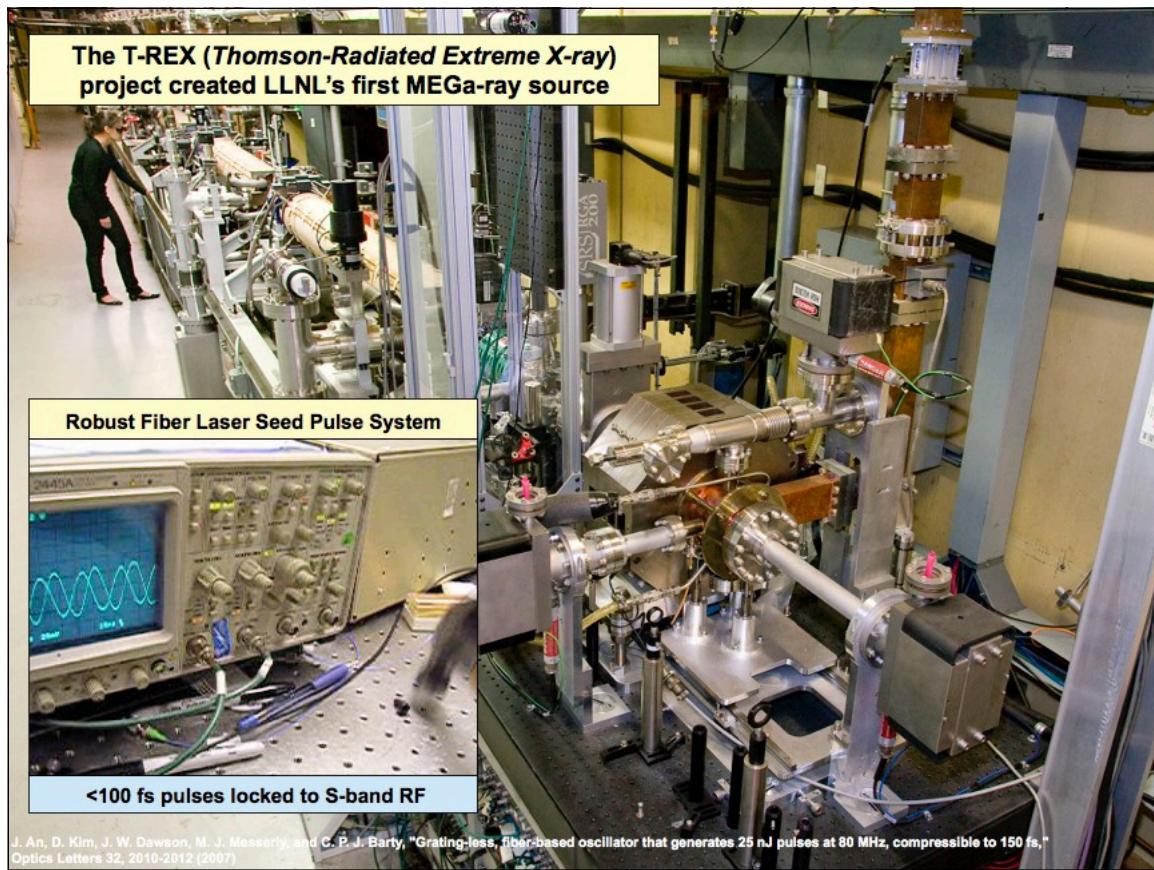
Using MEGA-ray technologies currently under development, identification of a ~5mm thick piece of 235U would take << 1s

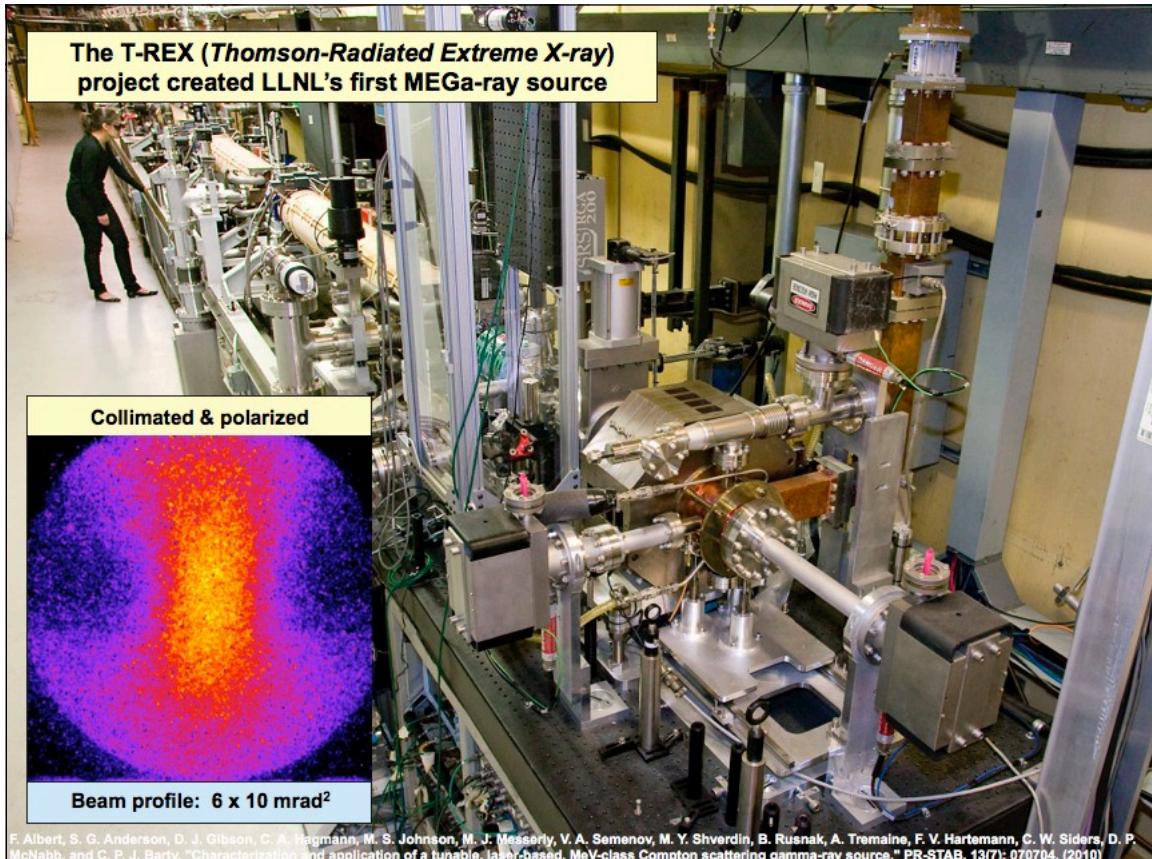
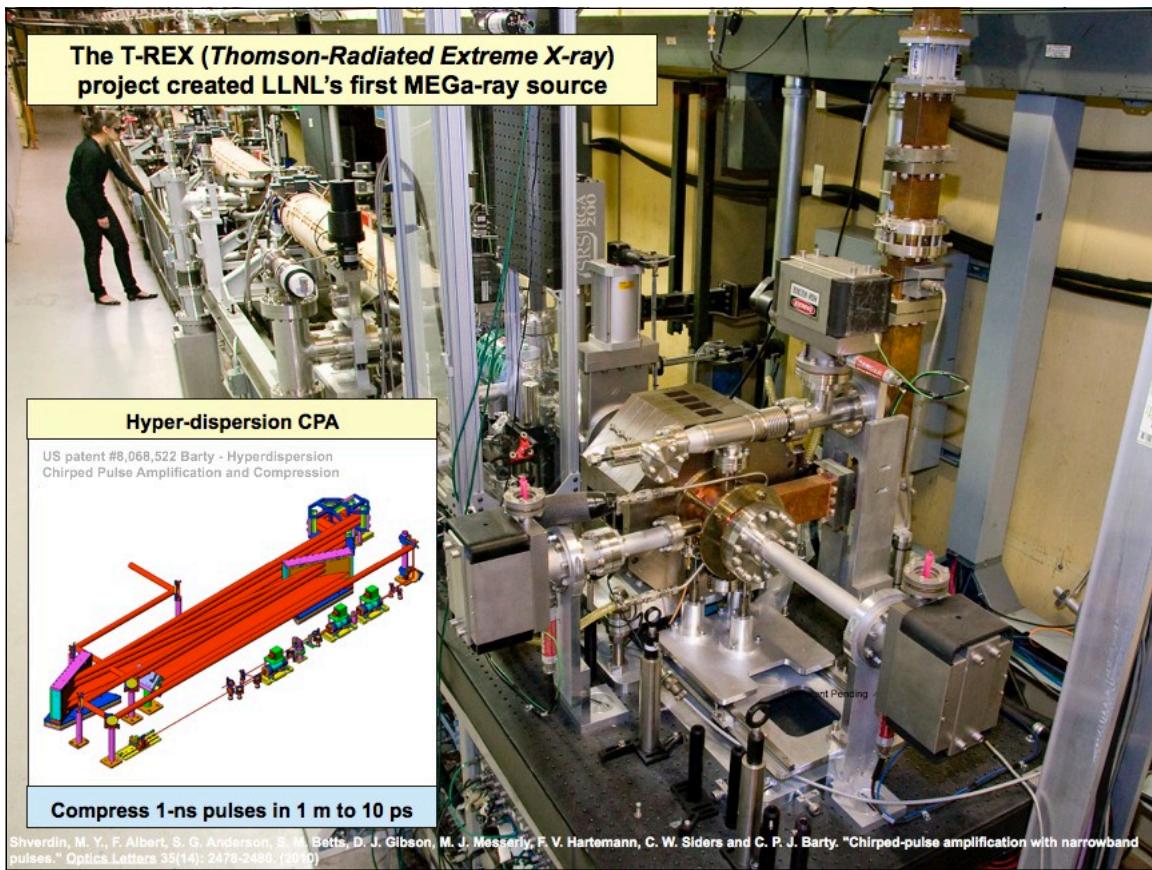


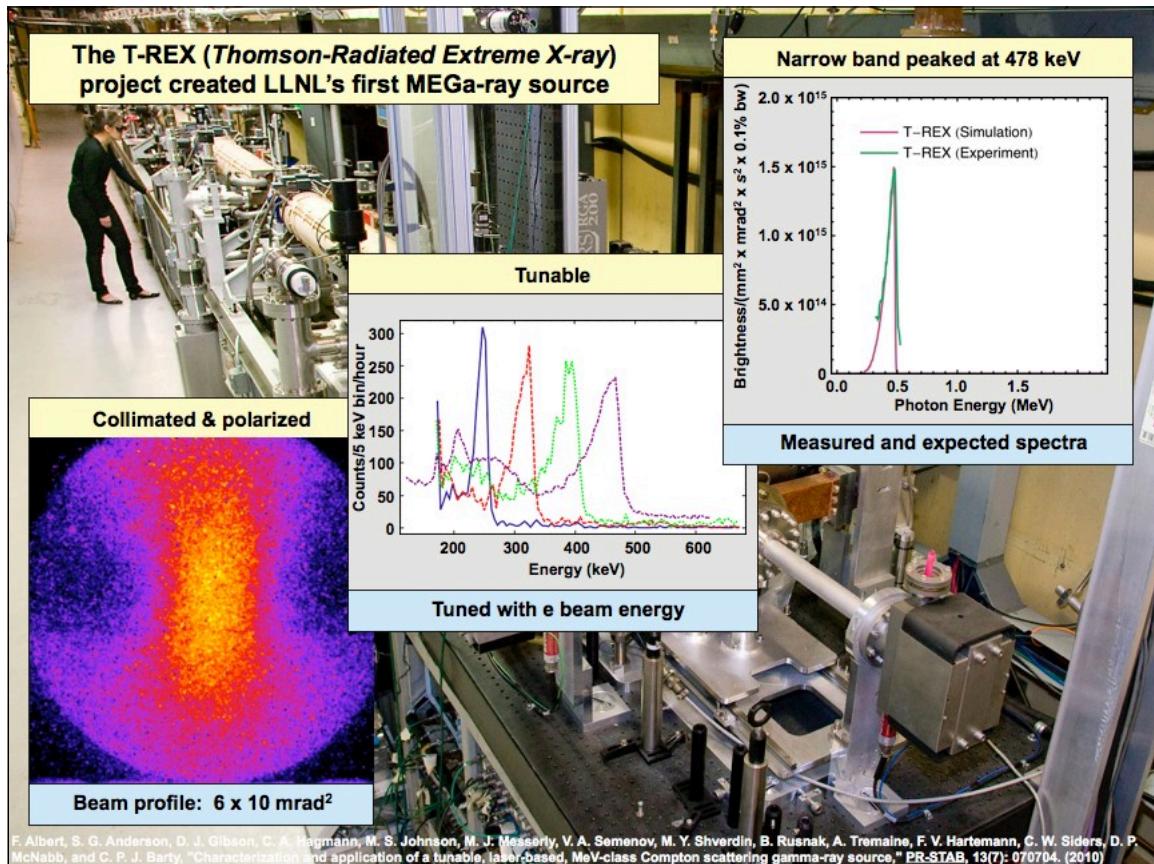
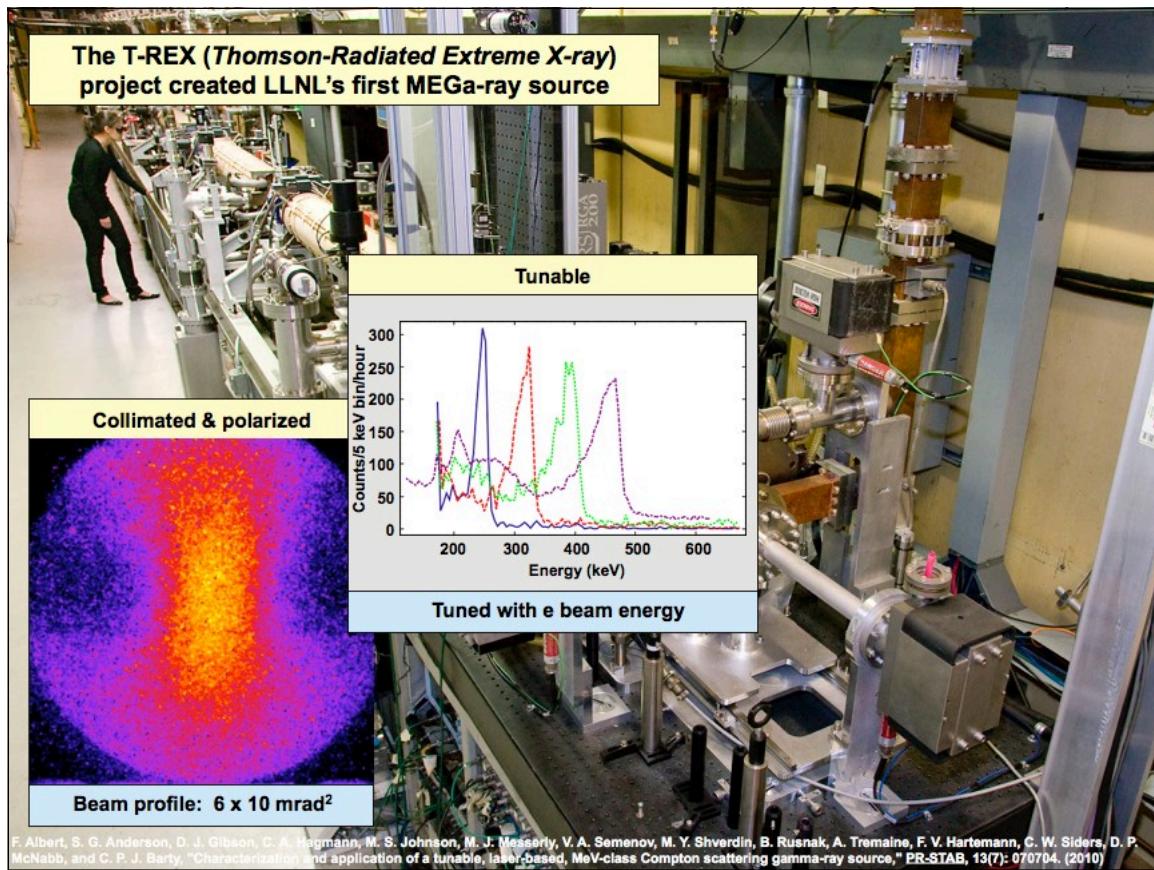
Numerical analysis suggests that active (but not "activating") SNM detection could be accomplished with MEGA-rays at doses that are 100x below allowable human limits

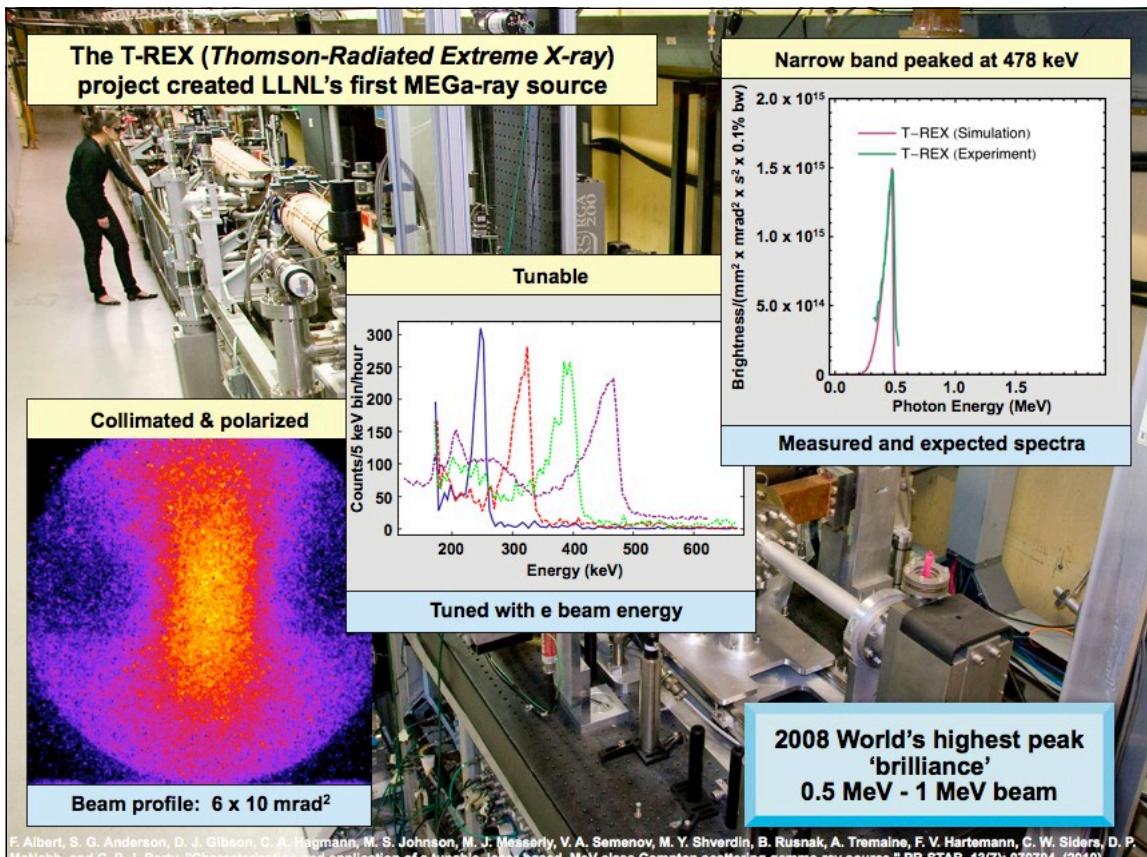




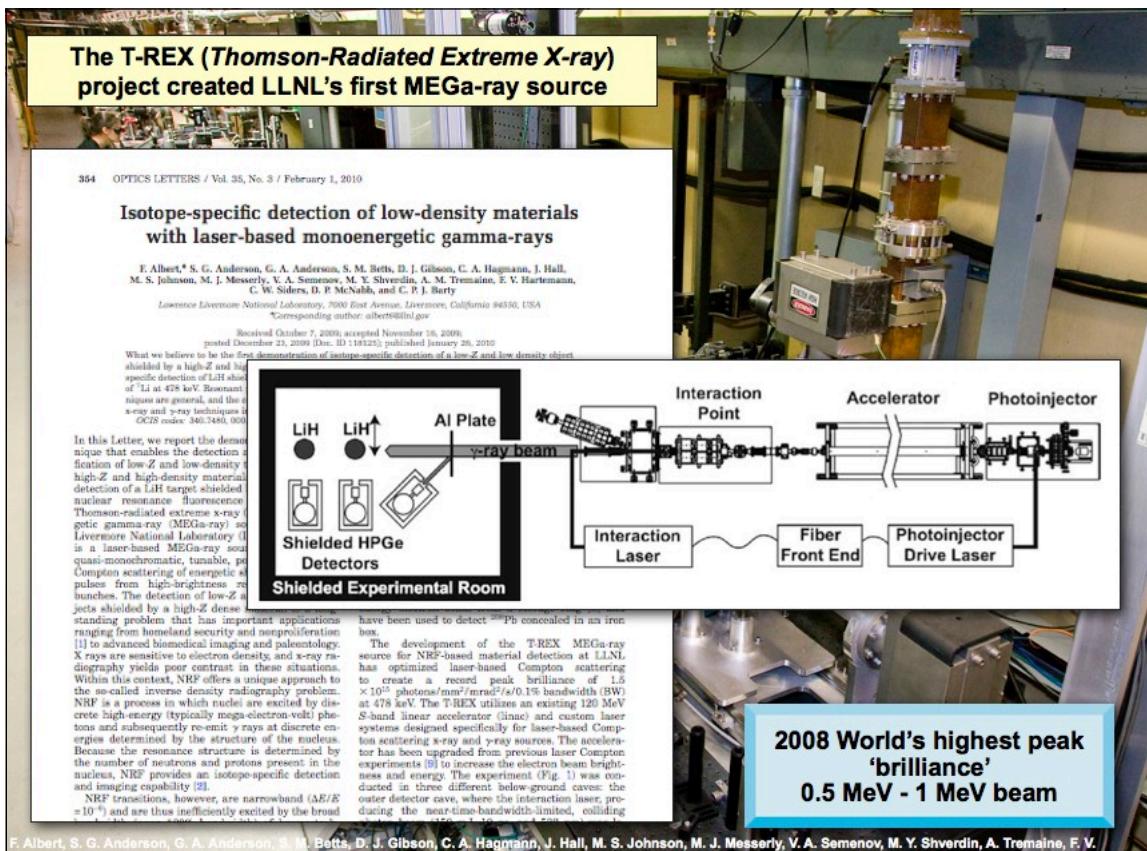






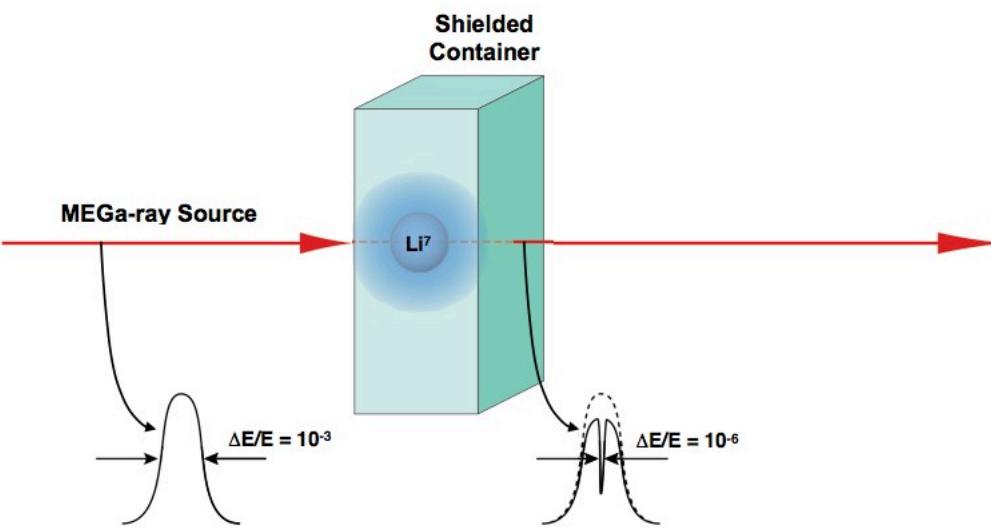


F. Albert, S. G. Anderson, D. J. Gibson, C. A. Hagmann, M. S. Johnson, M. J. Messerly, V. A. Semenov, M. Y. Shverdin, B. Rusnak, A. Tremaine, F. V. Hartemann, C. W. Siders, D. P. McNabb, and C. P. J. Barty, "Characterization and application of a tunable, laser-based, MeV-class Compton scattering gamma-ray source," *PR-STAB*, 13(7): 070704. (2010)

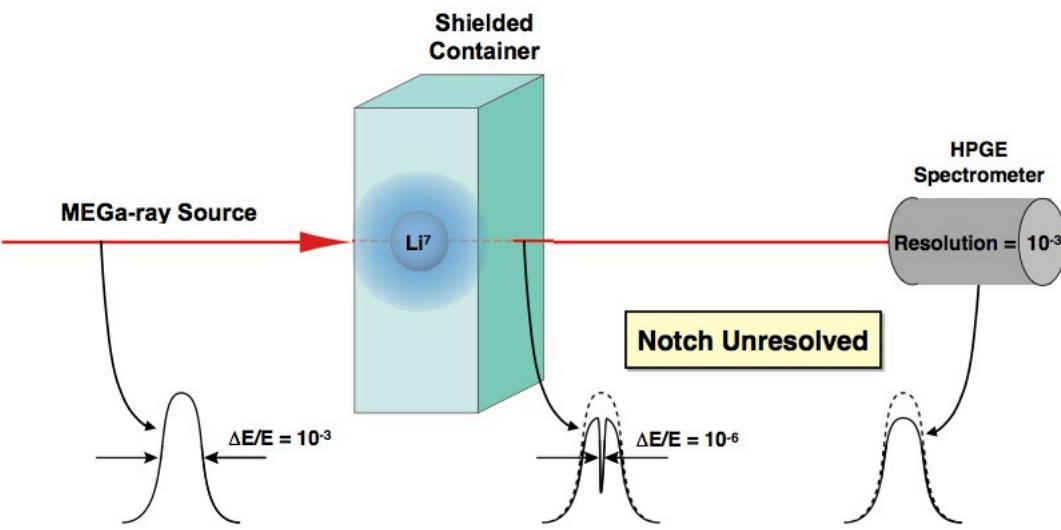


P. J. Barty, "Isotope-specific detection of low-density materials with laser-based monoenergetic gamma-rays," *Opt Lett* 35, (2010)

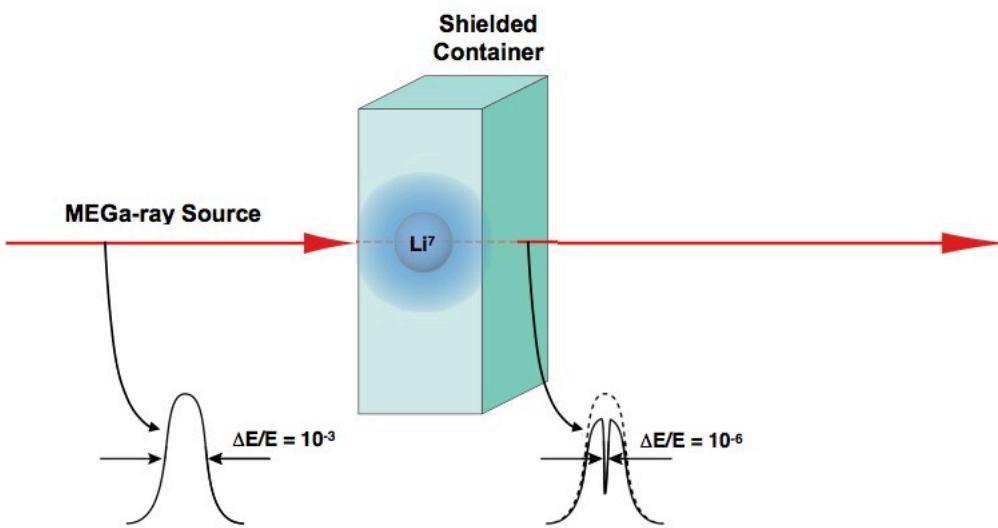
**Transmission-based detection was used for our initial material detection experiments**



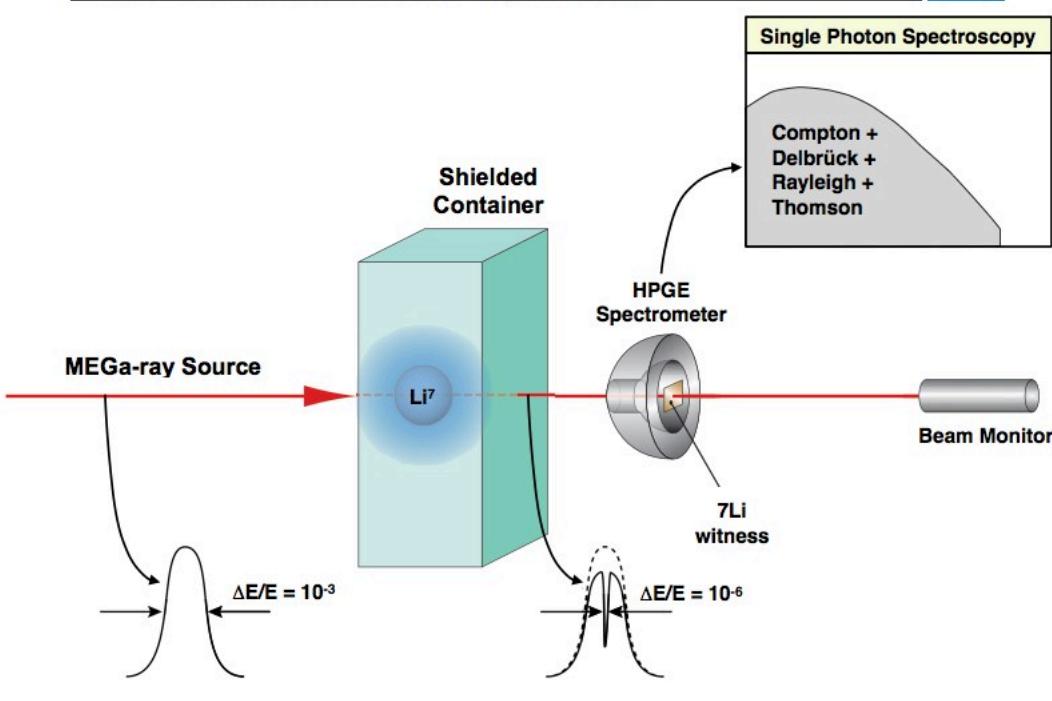
**Transmission-based detection was used for our initial material detection experiments**



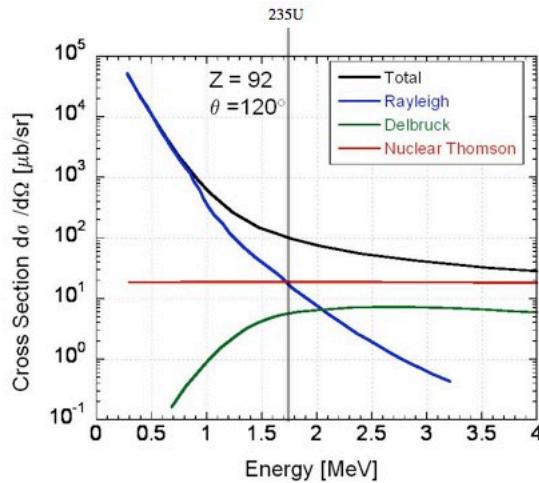
**Transmission-based detection was used for our initial material detection experiments**



**Transmission-based detection was used for our initial material detection experiments**

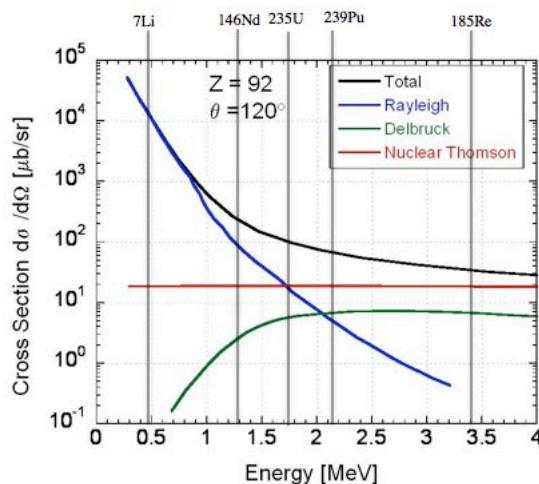


## Nuclear Thomson, Rayleigh and Delbrück coherent scattering channels must also be considered



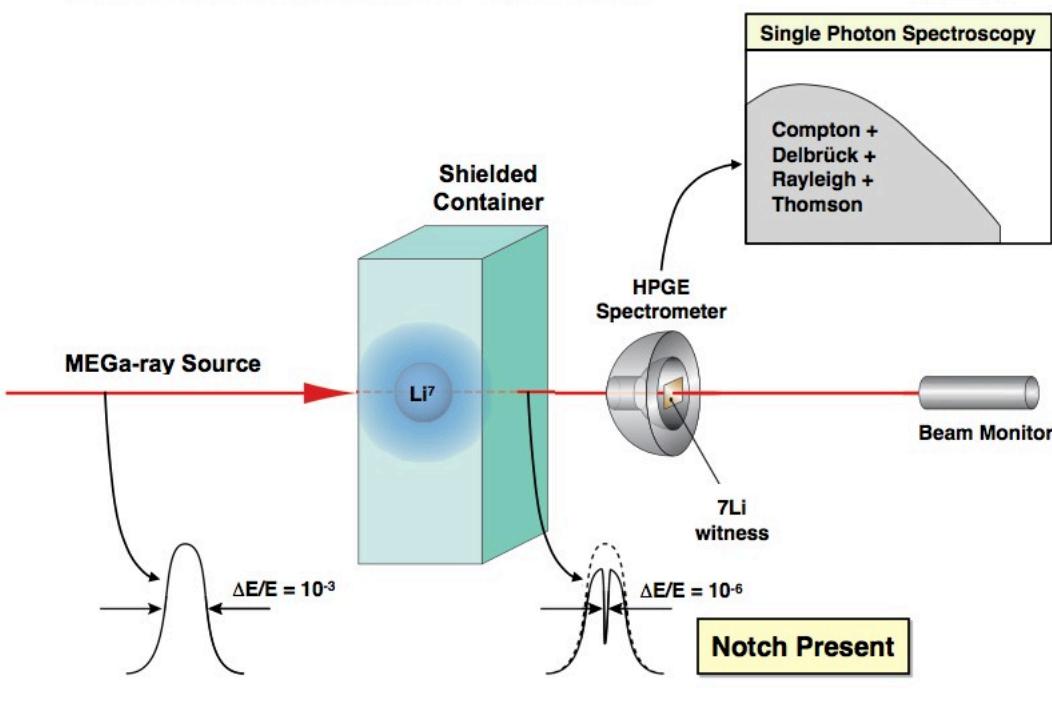
- Magnitude of nuclear Thomson and Delbrück cross sections are important at 1 MeV and comparable to Rayleigh near 2 MeV
- Must be properly treated for MEGa-ray applications and detector design, especially for high  $Z$

## Nuclear Thomson, Rayleigh and Delbrück coherent scattering channels must also be considered

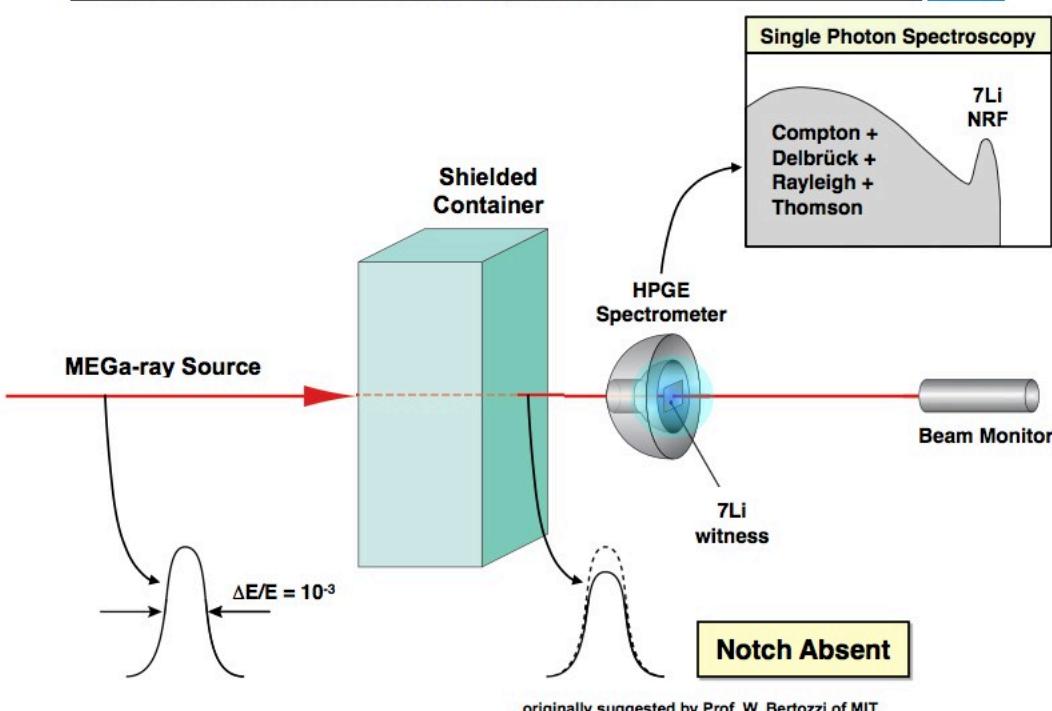


- Magnitude of nuclear Thomson and Delbrück cross sections are important at 1 MeV and comparable to Rayleigh near 2 MeV
- Must be properly treated for MEGa-ray applications and detector design, especially for high  $Z$

**Transmission-based detection was used for our initial material detection experiments**



**Transmission-based detection was used for our initial material detection experiments**

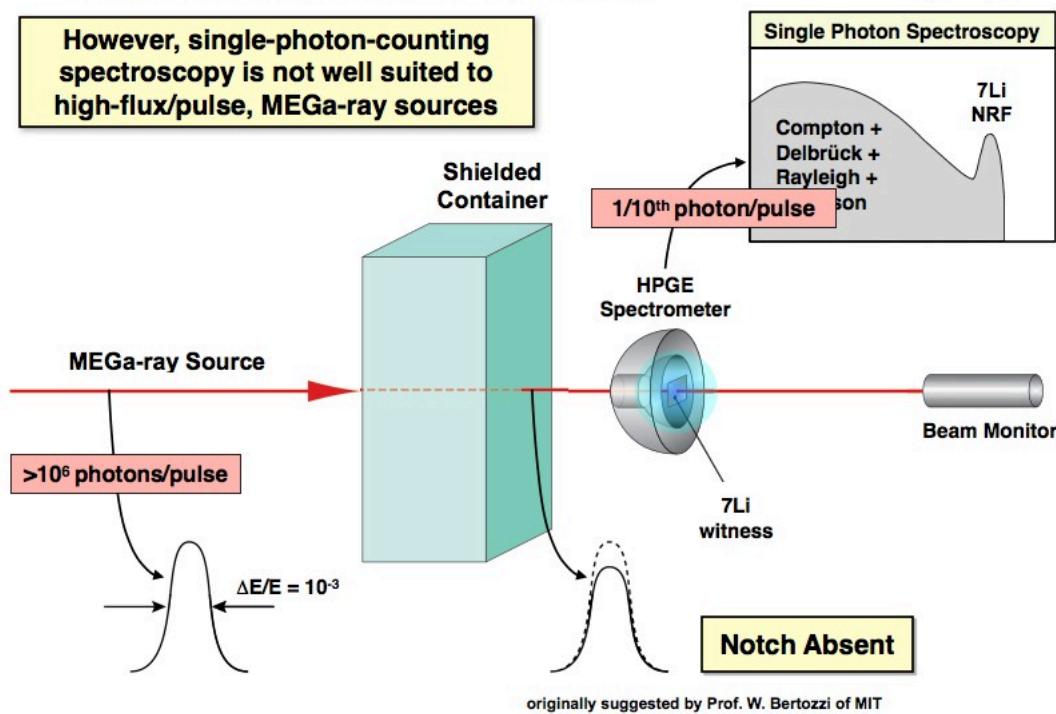


originally suggested by Prof. W. Bertozzi of MIT

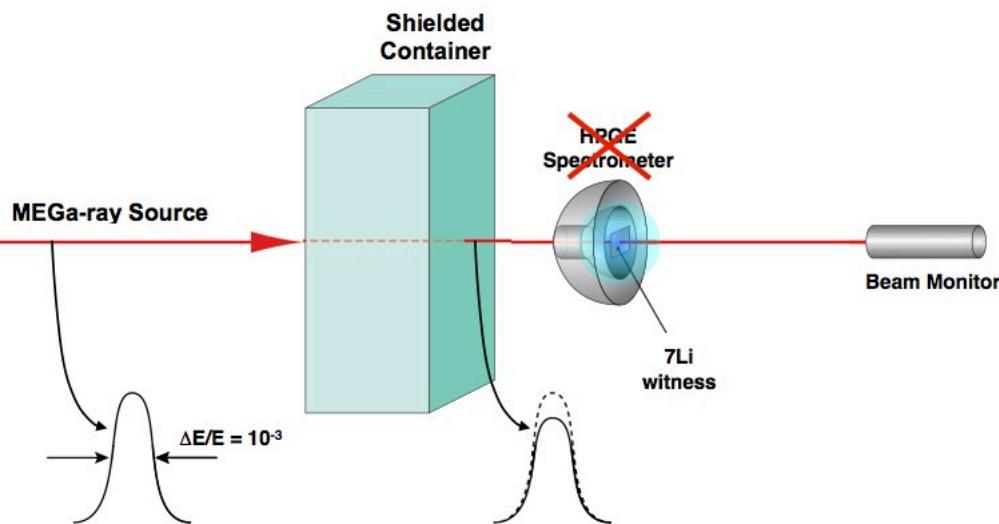
**Transmission-based detection was used for our initial material detection experiments**



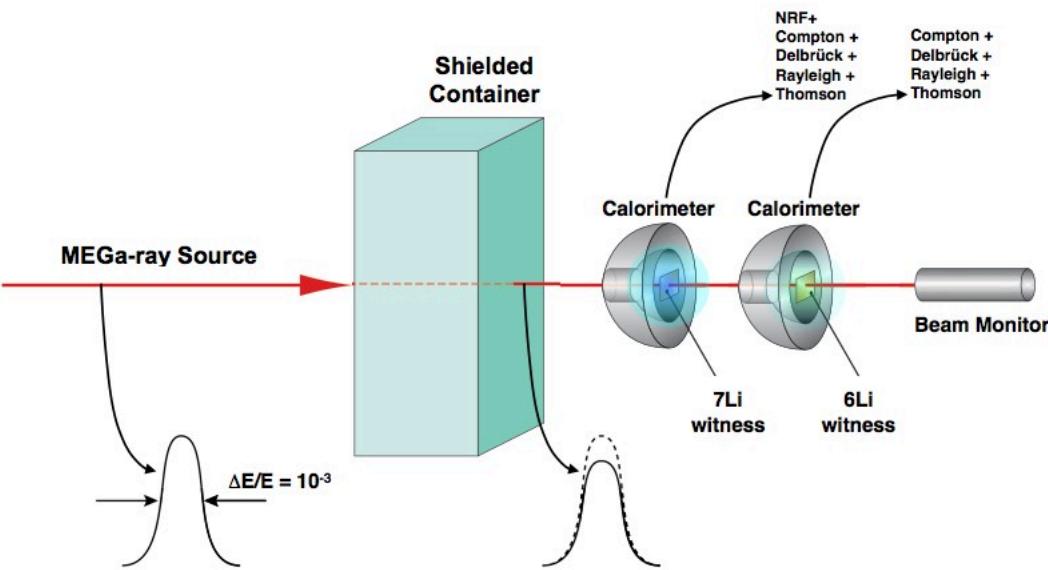
However, single-photon-counting spectroscopy is not well suited to high-flux/pulse, MEGa-ray sources



**Transmission-based detection was used for our initial material detection experiments**

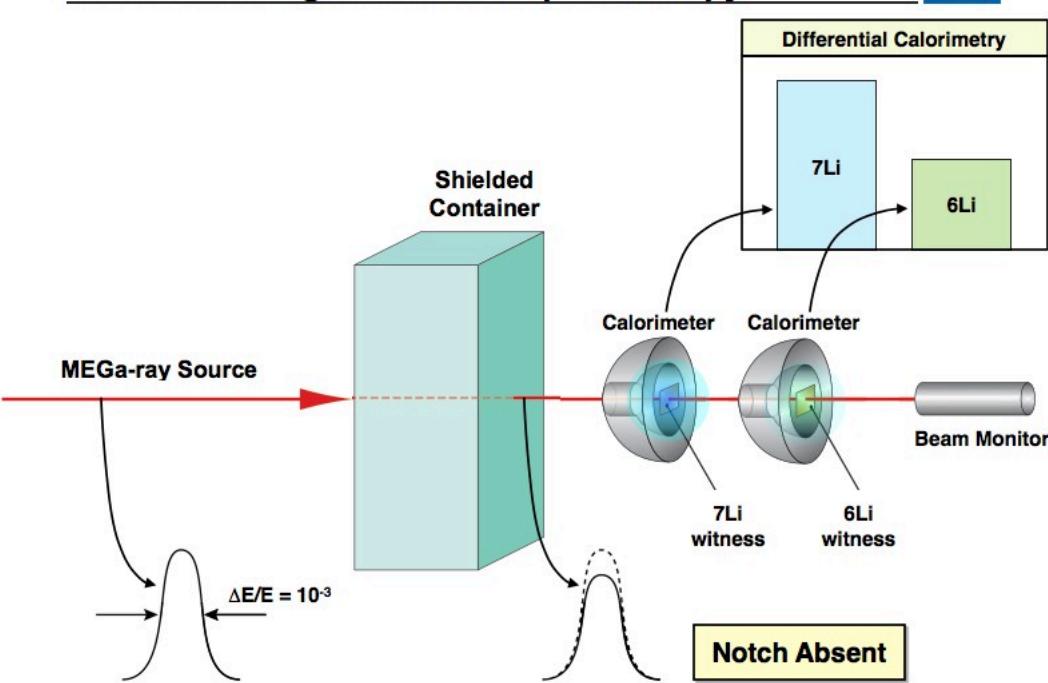


**Dual Isotope Notch Observation (DINO) eliminates  
the need for high resolution spectroscopy**



US patent #8,369,480 Barty, C. P. J. - Dual isotope notch observer for material identification, assay and imaging

**Dual Isotope Notch Observation (DINO) eliminates  
the need for high resolution spectroscopy**

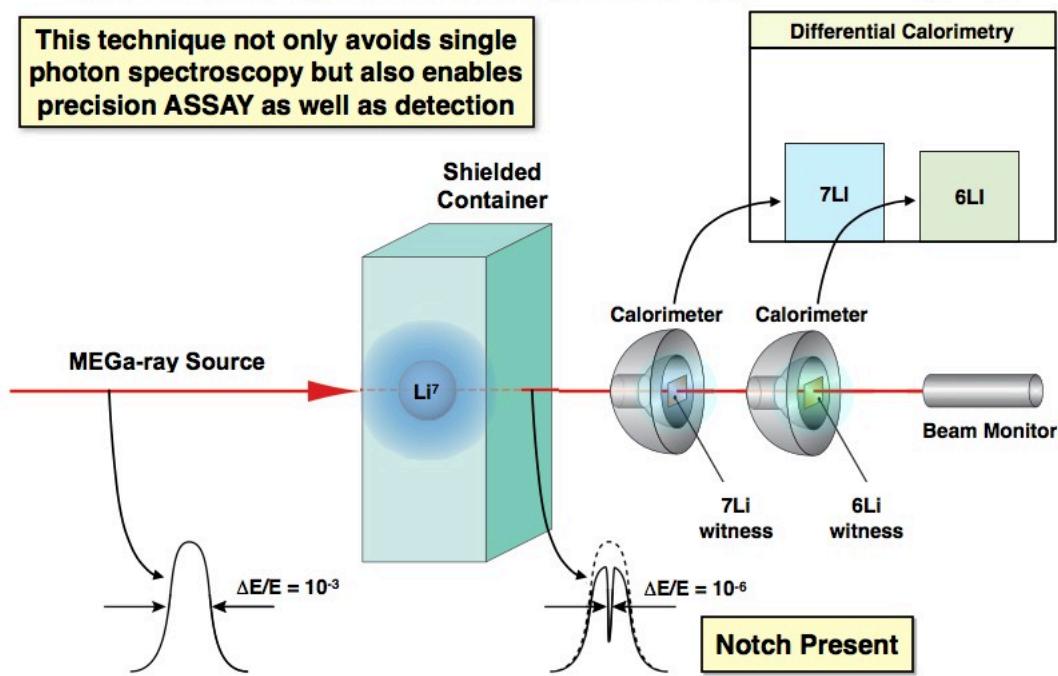


US patent #8,369,480 Barty, C. P. J. - Dual isotope notch observer for material identification, assay and imaging

## Dual Isotope Notch Observation (DINO) eliminates the need for high resolution spectroscopy



This technique not only avoids single photon spectroscopy but also enables precision ASSAY as well as detection

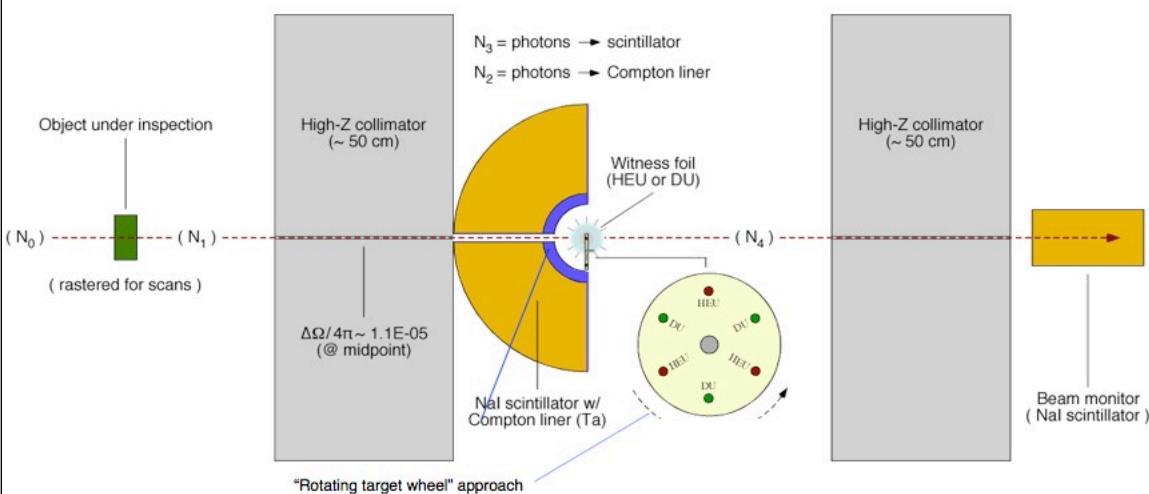


US patent #8,369,480 Barty, C. P. J. - Dual isotope notch observer for material identification, assay and imaging

## We are currently designing and simulating DINO systems for U & Pu detection and assay applications



Strongest  $^{235}\text{U}$  NRF resonance = 1733 keV



\* Note: The total dwell (object inspection) time is assumed to be equally divided between exposures of the HEU & DU witness foils in this case.

MC model of single-stage DINO detector system\* (COG):

**There are numerous design considerations for isotopic-specific applications of DINO detectors**



- Dimensions of the witness “foils” or more accurately the witness “pins”
- Configuration of the witness “pins”
- Solid angle subtended by the calorimeter
- Methods for rejection of non-resonant scattering
- Composition and efficiency of the “calorimeter”
- Collimation of the MEGA-ray illumination source
- Bandwidth of the MEGA-ray illumination source
- “Stability” of the MEGA-ray illumination source
- Choice of DINO “decision metric”
- etc....

**There are numerous design considerations for isotopic-specific applications of DINO detectors**



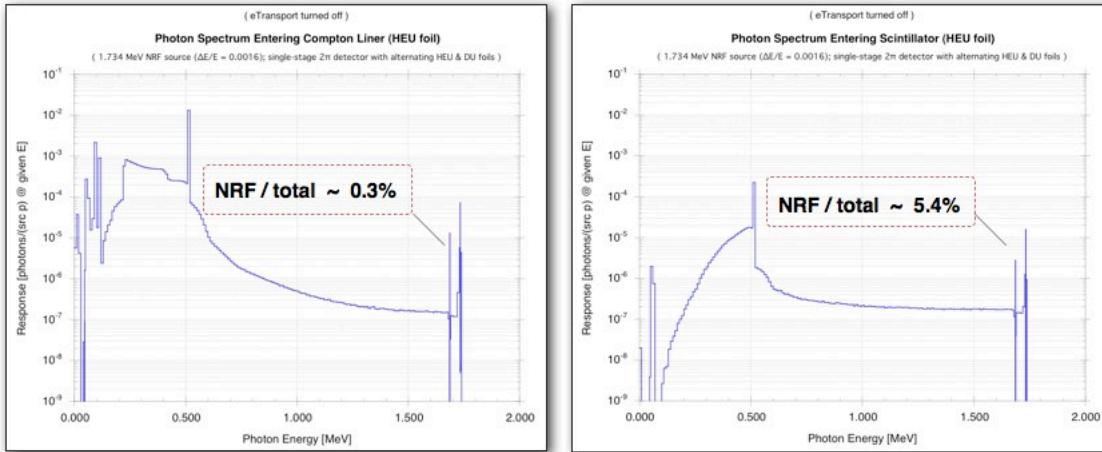
- Dimensions of the witness “foils” or more accurately the witness “pins”
- Configuration of the witness “pins”
- Solid angle subtended by the calorimeter
- Methods for rejection of non-resonant scattering
- Composition and efficiency of the “calorimeter”
- Collimation of the MEGA-ray illumination source
- Bandwidth of the MEGA-ray illumination source
- “Stability” of the MEGA-ray illumination source
- Choice of DINO “decision metric”
- etc....

Monte Carlo simulations using COG



## Design of the Compton liner can greatly increase the ratio of NRF to background in the scintillator

Strongest  $^{235}\text{U}$  NRF resonance = 1733 keV

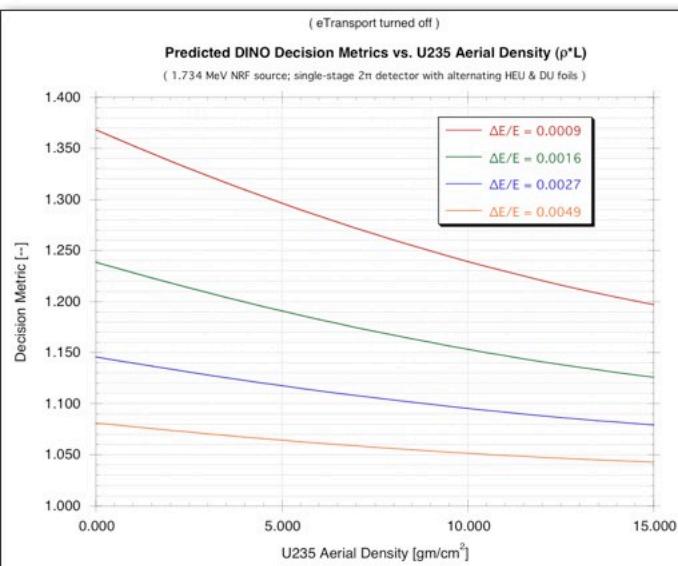


Hall, J., V. Semenov, F. Albert and C. P. J. Barty. "Numerical Simulation of Nuclear Materials Detection, Imaging and Assay with MEGA-rays." Proceedings of the 52nd Annual Meeting of the Institute for Nuclear Materials Management Vol 3. 2673-2682. (2011)

Monte Carlo simulations using COG

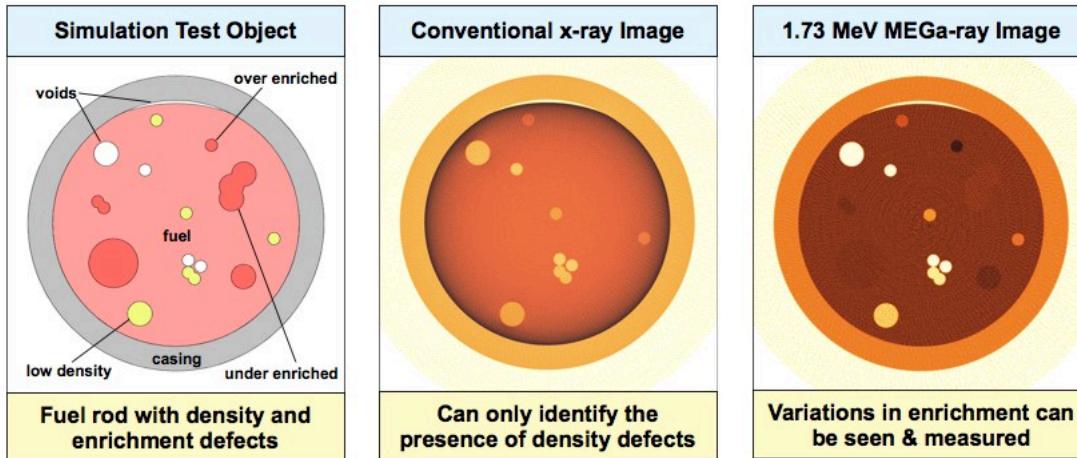


## Sensitivity of the detector decreases rapidly with respect to source bandwidth



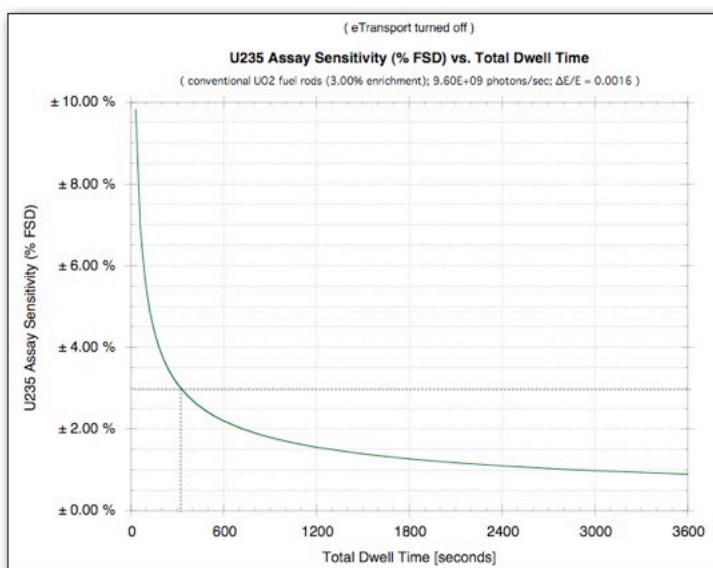
Hall, J., V. Semenov, F. Albert and C. P. J. Barty. "Numerical Simulation of Nuclear Materials Detection, Imaging and Assay with MEGA-rays." Proceedings of the 52nd Annual Meeting of the Institute for Nuclear Materials Management Vol 3. 2673-2682. (2011)

## Precision assay and imaging of nuclear fuel would be possible with MEGa-rays and DINO



Hall, J., V. Semenov, F. Albert and C. P. J. Barty. "Numerical Simulation of Nuclear Materials Detection, Imaging and Assay with MEGa-rays." Proceedings of the 52nd Annual Meeting of the Institute for Nuclear Materials Management Vol 3. 2673-2682. (2011)

## Simulations indicate that the enrichment of fuel rods could be measured to ~3% in 5 to 6 minutes



Hall, J., V. Semenov, F. Albert and C. P. J. Barty. "Numerical Simulation of Nuclear Materials Detection, Imaging and Assay with MEGa-rays." Proceedings of the 52nd Annual Meeting of the Institute for Nuclear Materials Management Vol 3. 2673-2682. (2011)

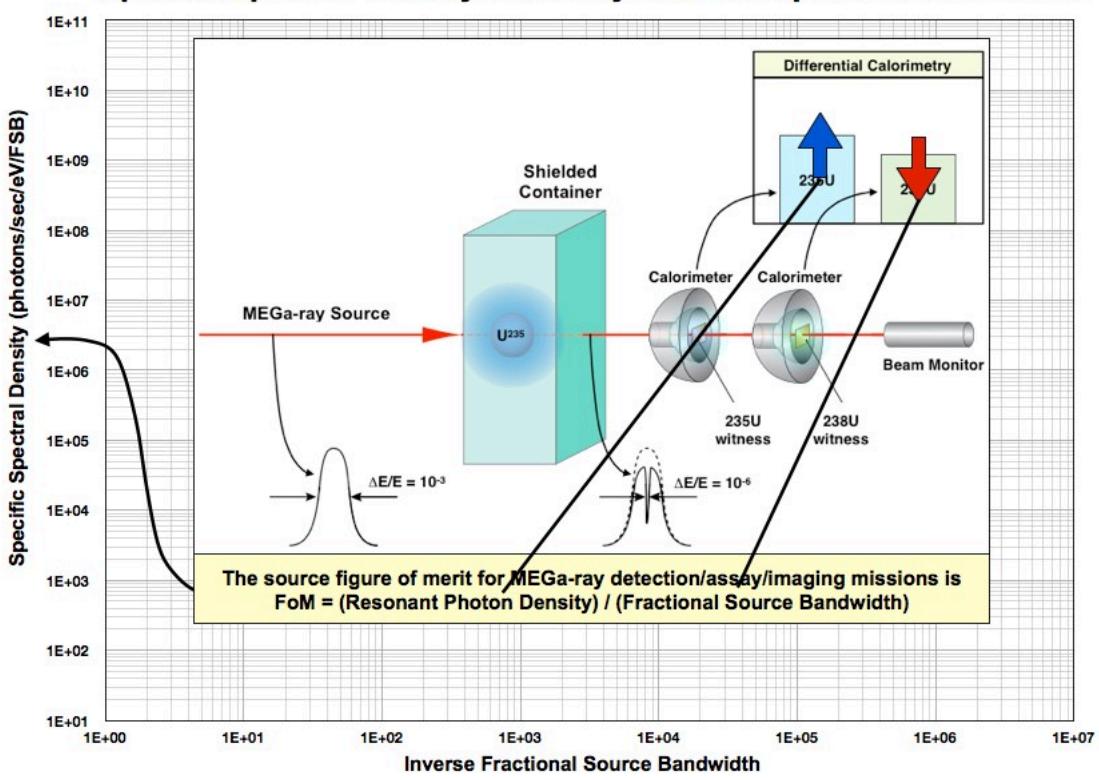
## NRF-based materials evaluation will be much easier elsewhere in the periodic table



atomic number	isotope	oxide	natural isotopic abundance (%)	REE-Oxide weight % at Mountain Pass	REE Atomic % at Mountain Pass	REE Isotope % at Mountain Pass	cost per kg in 2009 (US\$)	NRF gamma energy (keV)	NRF cross section ( barns)	Background cross section ( barns)	NRF to background cross section ratio
92	235U	UO <sub>2</sub>	90.0	na	na	na	\$13.50	1733	26.0	20.0	1.3
39	89Y	Y <sub>2</sub> O <sub>3</sub>	100.0	0.130	0.0092	0.0092	\$13.50	1507	24.5	6.9	3.6
40	90Zr	Lu <sub>2</sub> O <sub>3</sub>	51.5	na	na	na	\$13.50	2186	5.5	6.1	0.9
57	138La	Lu <sub>2</sub> O <sub>3</sub>	0.1	33.784	2.5925	0.0023	\$5.85	no data	no data	no data	no data
57	139La	Lu <sub>2</sub> O <sub>3</sub>	99.9	33.784	2.5925	0.0023	\$5.85	1538	7.2	10.6	0.7
58	136Ce	Ce <sub>2</sub> O <sub>3</sub>	0.19	49.581	3.8096	0.0070	\$4.15	552	5.4	23	0.2
58	138Ce	Ce <sub>2</sub> O <sub>3</sub>	0.25	49.581	3.8096	0.0095	\$4.15	789	6.2	16	0.4
58	140Ce	Ce <sub>2</sub> O <sub>3</sub>	88.45	49.581	3.8096	3.3696	\$4.15	1598	16.9	10.5	1.6
58	142Ce	Ce <sub>2</sub> O <sub>3</sub>	11.11	49.581	3.8096	0.4234	\$4.15	2187	19.3	9.4	2.1
59	141Pr	Pr <sub>2</sub> O <sub>3</sub>	100.0	4.119	0.3168	0.3168	\$15.15	no data	no data	no data	no data
60	142Nd	Nd <sub>2</sub> O <sub>3</sub>	27.2	11.158	0.8610	0.2342	\$15.25	3424	46.8	9.1	5.1
60	143Nd	Nd <sub>2</sub> O <sub>3</sub>	12.2	11.158	0.8610	0.1050	\$15.25	1407	10.9	12	0.9
60	144Nd	Nd <sub>2</sub> O <sub>3</sub>	23.8	11.158	0.8610	0.2049	\$15.25	2186	17.4	9.8	1.8
60	145Nd	Nd <sub>2</sub> O <sub>3</sub>	8.3	na	na	na	na	na	1.8	15	0.1
60	146Nd	Nd <sub>2</sub> O <sub>3</sub>	17.2	na	na	na	na	na	14.9	19	0.8
60	148Nd	Nd <sub>2</sub> O <sub>3</sub>	5.7	na	na	na	na	na	5.4	9.1	0.6
60	150Nd	Nd <sub>2</sub> O <sub>3</sub>	5.6	na	na	na	na	na	1.8	14	0.1
62	144Sm	Sm <sub>2</sub> O <sub>3</sub>	3.1	na	na	na	na	na	55.2	9.5	5.8
62	147Sm	Sm <sub>2</sub> O <sub>3</sub>	15.0	na	na	na	na	na	16.6	11.5	1.4
62	148Sm	Sm <sub>2</sub> O <sub>3</sub>	11.2	na	na	na	na	na	2.2	20	0.1
62	149Sm	Sm <sub>2</sub> O <sub>3</sub>	13.8	na	na	na	na	na	8.4	12	0.7
62	150Sm	Sm <sub>2</sub> O <sub>3</sub>	7.4	na	na	na	na	na	9.2	23	0.1
62	152Sm	Sm <sub>2</sub> O <sub>3</sub>	26.8	na	na	na	na	na	29.1	13.6	1.7
62	152Sm	Sm <sub>2</sub> O <sub>3</sub>	26.8	na	na	na	na	na	18.0	9.7	1.9
62	154Sm	Sm <sub>2</sub> O <sub>3</sub>	22.8	0.850	0.0960	0.0150	\$4.50	921	70.5	15	4.7
62	154Sm	Sm <sub>2</sub> O <sub>3</sub>	22.8	0.850	0.0960	0.0150	\$4.50	1440	95.1	16	5.9
63	151Eu	Eu <sub>2</sub> O <sub>3</sub>	47.8	0.105	0.0082	0.0099	\$465.00	1008	3.8	12.1	0.3
63	153Eu	Eu <sub>2</sub> O <sub>3</sub>	52.2	0.105	0.0082	0.0043	\$465.00	no data	16.7	60	0.3
64	152Gd	Gd <sub>2</sub> O <sub>3</sub>	0.2	0.210	0.0164	0.0000	\$6.50	344	5.0	50	0.1
64	154Gd	Gd <sub>2</sub> O <sub>3</sub>	2.2	0.210	0.0164	0.0004	\$6.50	1241	735.3	13.8	53.3
64	155Gd	Gd <sub>2</sub> O <sub>3</sub>	14.8	0.210	0.0164	0.0024	\$6.50	615	0.2	90	0.0
64	156Gd	Gd <sub>2</sub> O <sub>3</sub>	20.5	0.210	0.0164	0.0034	\$6.50	1243	33.1	137	2.4
64	156Gd	Gd <sub>2</sub> O <sub>3</sub>	20.5	0.210	0.0164	0.0034	\$6.50	1367	32.4	13	2.5
64	156Gd	Gd <sub>2</sub> O <sub>3</sub>	20.5	0.210	0.0164	0.0034	\$6.50	2745	16.1	102	1.6
64	157Gd	Gd <sub>2</sub> O <sub>3</sub>	15.7	0.210	0.0164	0.0026	\$6.50	3071	41.2	10	4.1
64	158Gd	Gd <sub>2</sub> O <sub>3</sub>	24.8	0.210	0.0164	0.0041	\$6.50	1264	60.9	13.7	4.4
64	158Gd	Gd <sub>2</sub> O <sub>3</sub>	24.8	0.210	0.0164	0.0041	\$6.50	3201	25.0	9.9	2.5
64	160Gd	Gd <sub>2</sub> O <sub>3</sub>	21.9	0.210	0.0164	0.0036	\$6.50	1224	56.5	13.7	4.1
65	159Tb	Tb <sub>2</sub> O <sub>3</sub>	100.0	0.016	0.0013	0.0013	\$350.00	58	192.7	3800	0.1
65	159Tb	Tb <sub>2</sub> O <sub>3</sub>	100.0	0.016	0.0013	0.0013	\$350.00	581	4.4	28	0.2

The ratio of the NRF to background cross section for many materials is significantly greater than that for the actinides

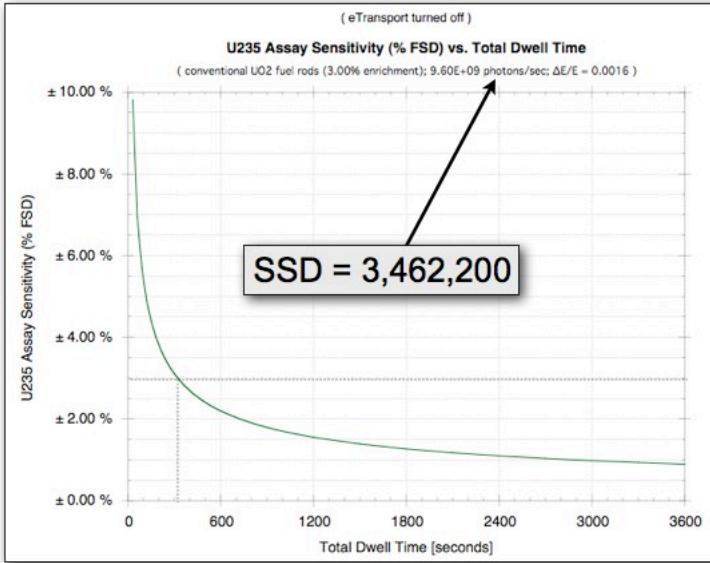
## Specific Spectral Density is the key laser-Compton source metric



Monte Carlo simulations using COG

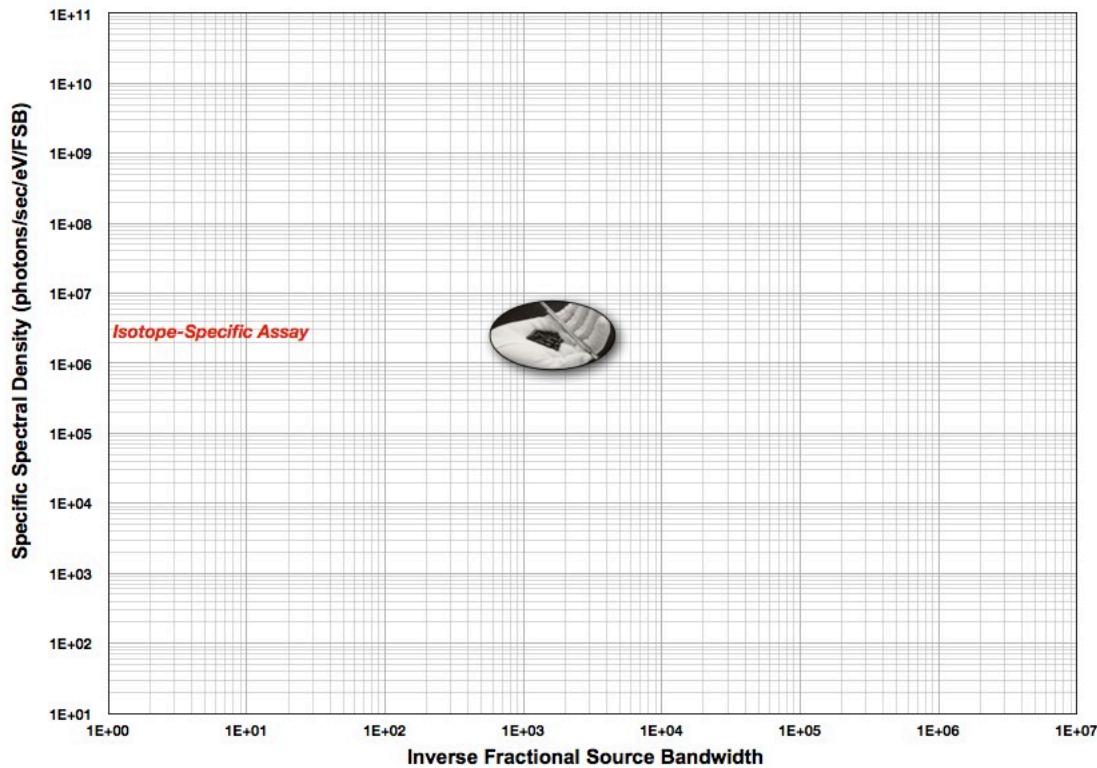


**Simulations indicate that the enrichment of fuel rods could be measured to ~3% in 5 to 6 minutes**

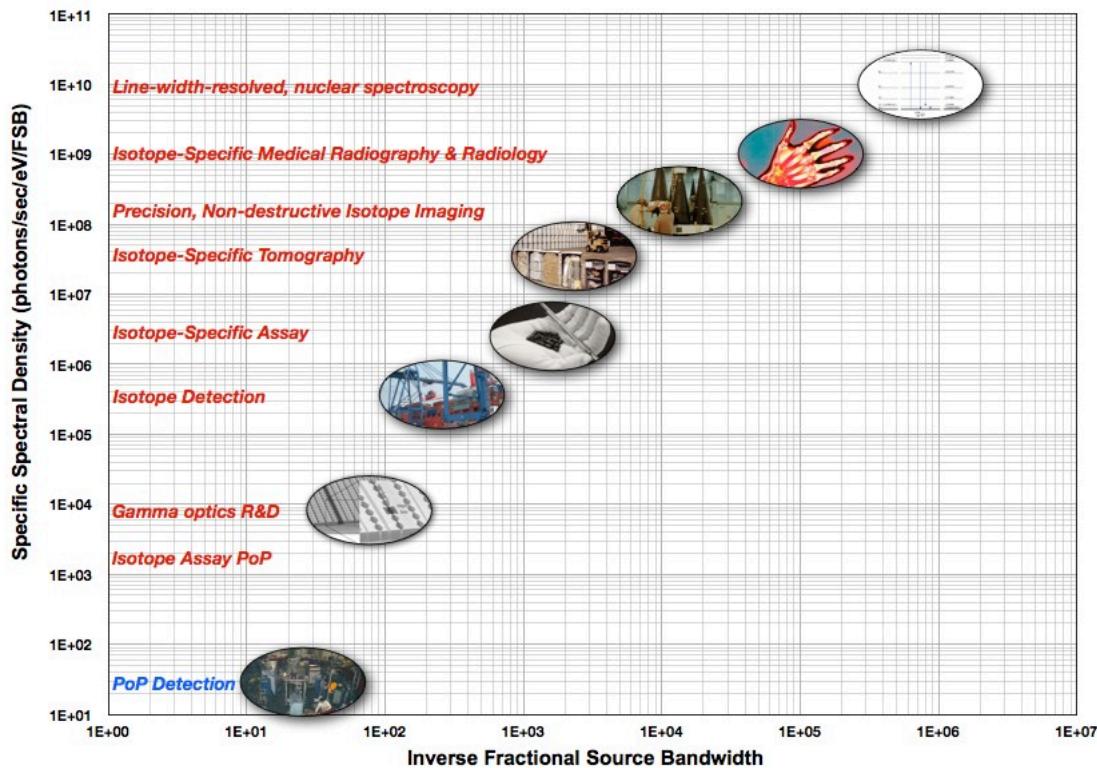


Hall, J., V. Semenov, F. Albert and C. P. J. Barty. "Numerical Simulation of Nuclear Materials Detection, Imaging and Assay with MEGa-rays." Proceedings of the 52nd Annual Meeting of the Institute for Nuclear Materials Management Vol 3. 2673-2682. (2011)

### Specific Spectral Density is the key source metric

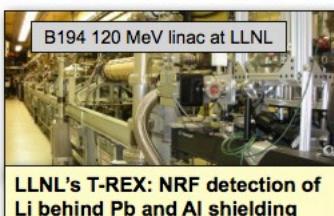


## New applications become viable with increasing SSD



US patent #7,564,241 Barty, Hartemann, McNabb & Pruet - detection, assay and imaging with laser-Compton gamma-rays

**Three laser-Compton gamma sources have been built from existing hardware & used NRF studies**



## Three laser-Compton gamma sources have been built from existing hardware & used NRF studies



- Tunable 0.5-0.9 MeV  
- 12 % bandwidth  
-  $10^6$  photons/s  
- 10 Hz

F. Albert et al, Opt. Lett. 2010  
D.G. Gibson et al, PRSTAB 2010  
F. Albert et al, PRSTAB 2010

- 5.7 MeV  
- 7 % bandwidth  
-  $10^5$  photons/s  
- 20 kHz

N. Kikuzawa et al, Appl. Phys. Expr 2009

- Tunable 1-97 MeV in 2011  
- 4 % bandwidth  
-  $10^6 - 10^7$  photons/s 1-3 MeV  
-  $10^8$  photons/s 3-20 MeV  
-  $10^7$  photons/s 21-100 MeV  
- 5.58 MHz  
Y. K. Wu, PAC 2011  
C. A. Hagmann et al, JAP 2009

## Three laser-Compton gamma sources have been built from existing hardware & used NRF studies



- Tunable 0.5-0.9 MeV  
- 12 % bandwidth  
-  $10^6$  photons/s  
- 10 Hz  
**SSD ~ 17**

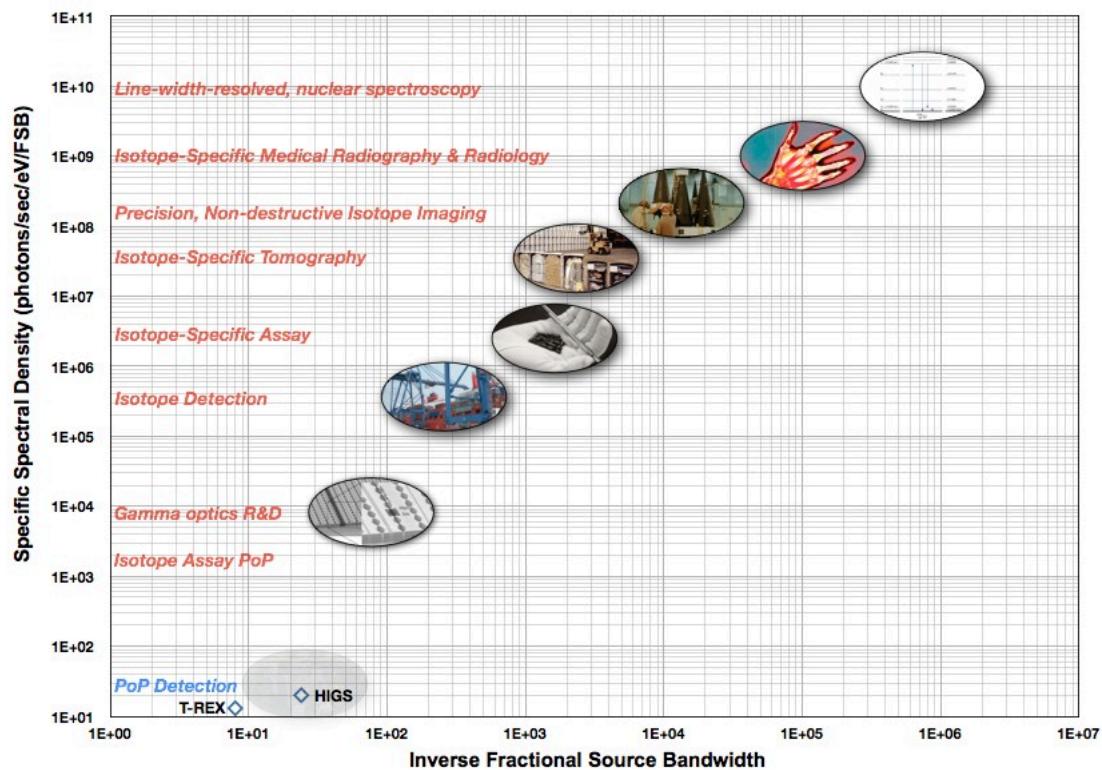
F. Albert et al, Opt. Lett. 2010  
D.G. Gibson et al, PRSTAB 2010  
F. Albert et al, PRSTAB 2010

- 5.7 MeV  
- 7 % bandwidth  
-  $10^5$  photons/s  
- 20 kHz  
**SSD = 0.3**

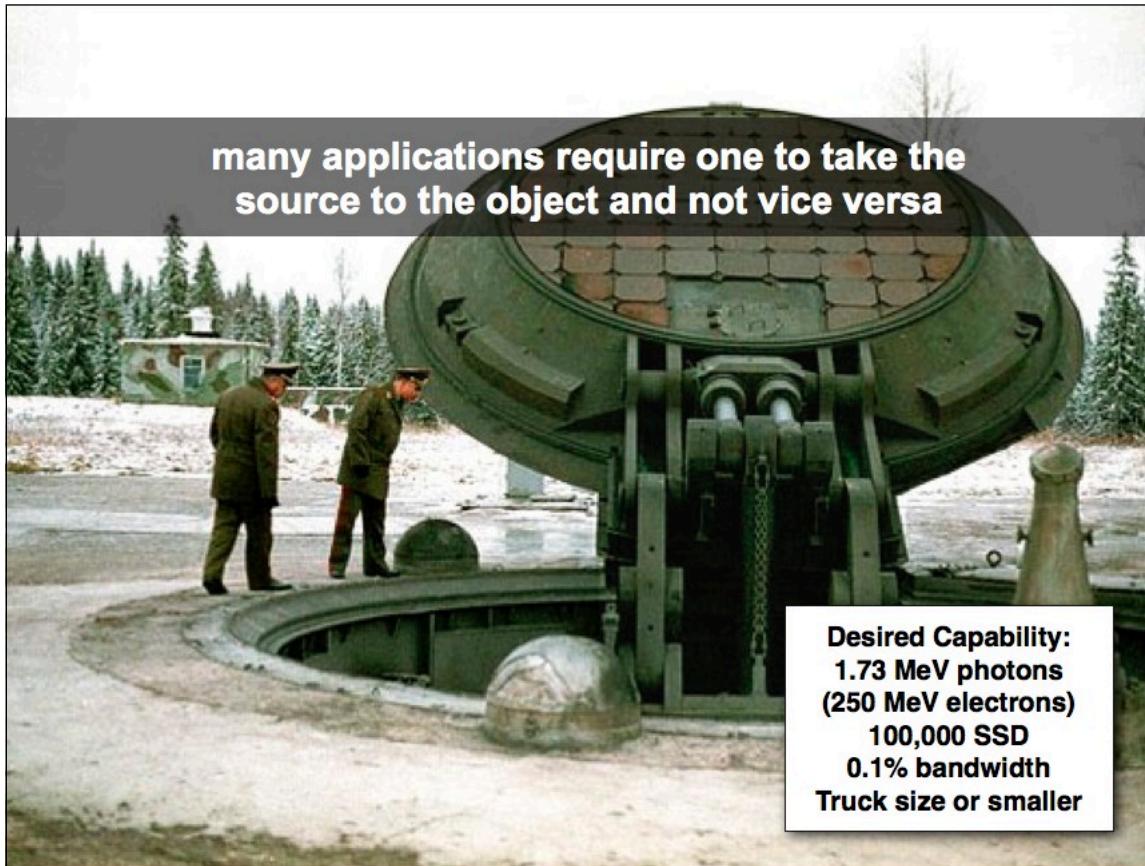
N. Kikuzawa et al, Appl. Phys. Expr 2009

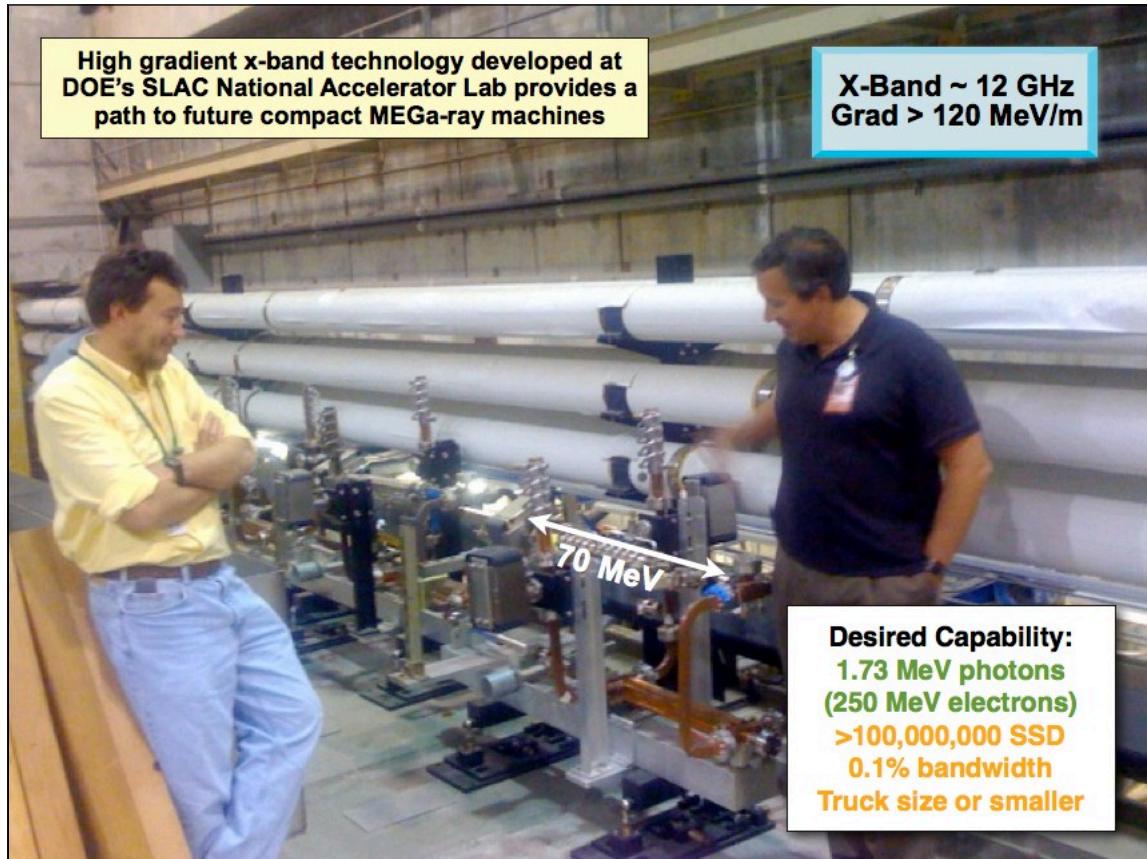
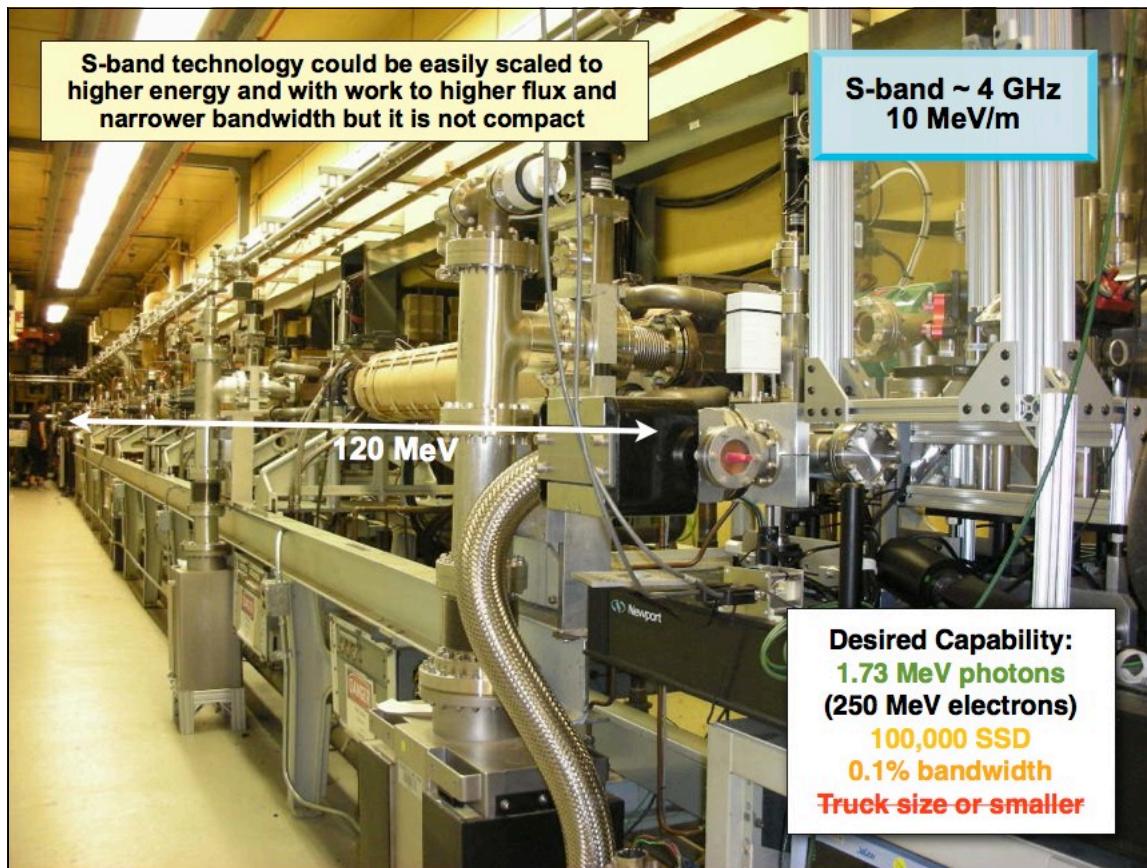
- Tunable 1-97 MeV in 2011  
- 4 % bandwidth  
-  $10^6 - 10^7$  photons/s 1-3 MeV  
- **SSD = 20 - 100**  
- 5.58 MHz  
Y. K. Wu, PAC 2011  
C. A. Hagmann et al, JAP 2009

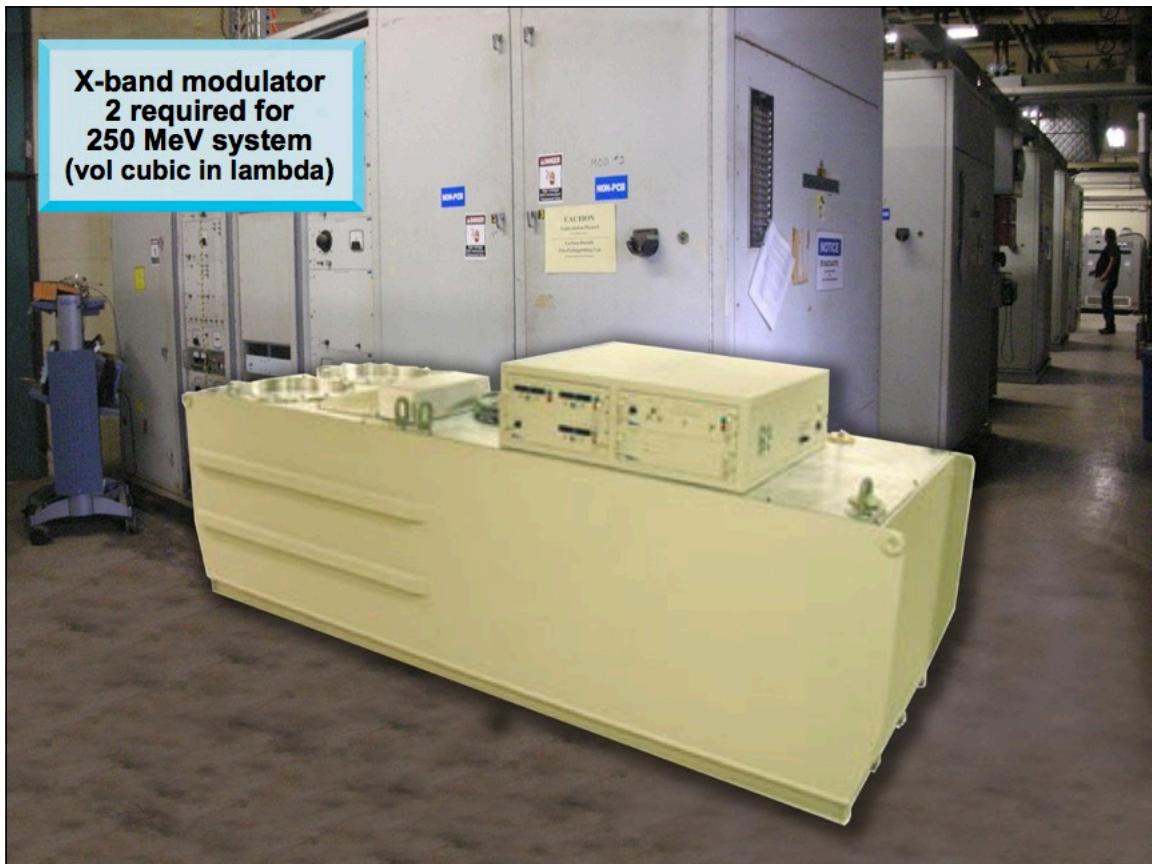
## New applications become viable with increasing SSD



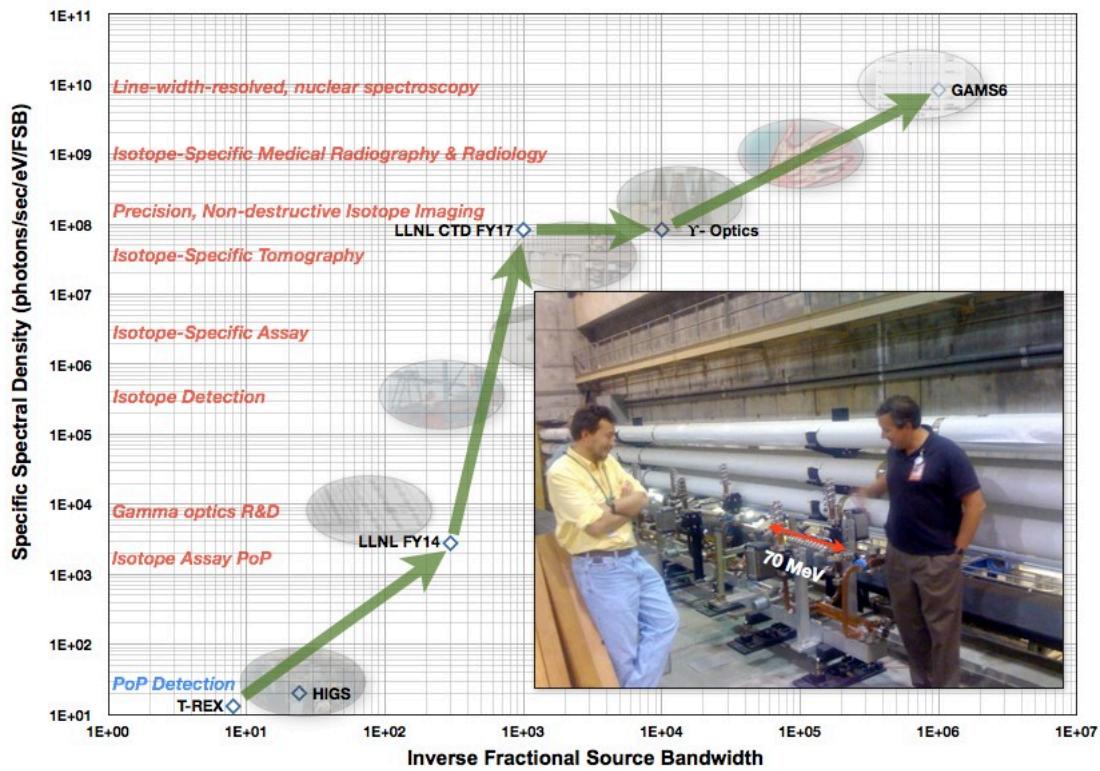
many applications require one to take the source to the object and not vice versa



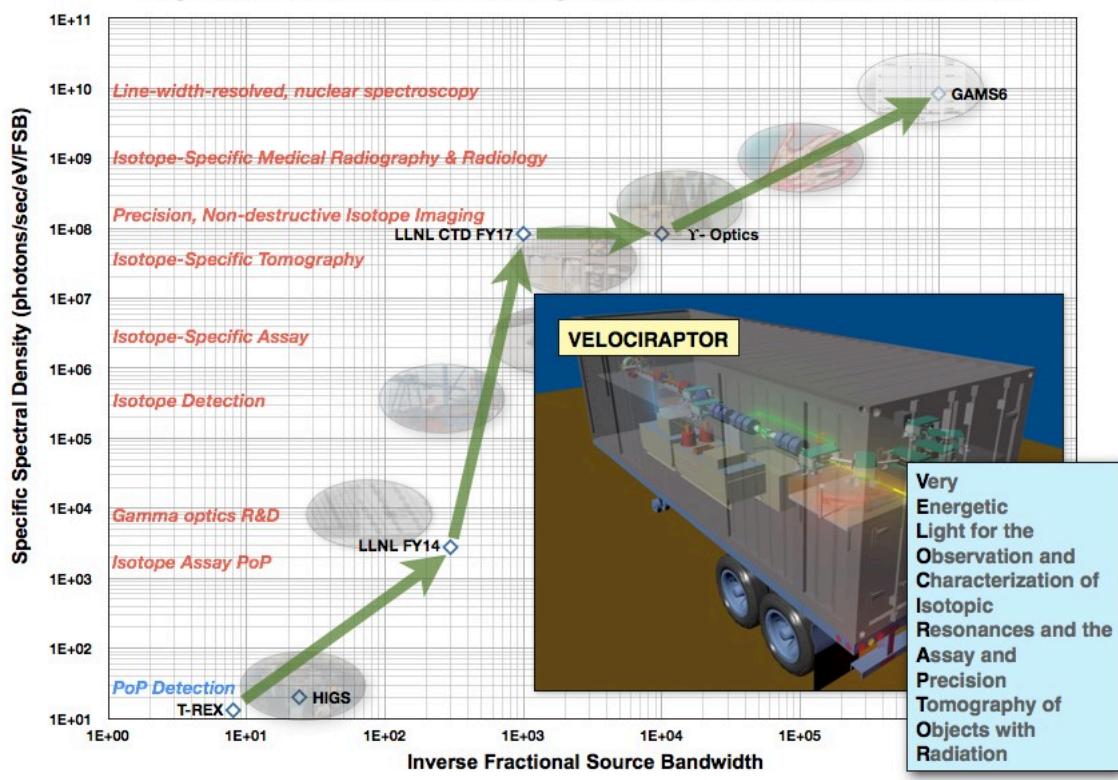




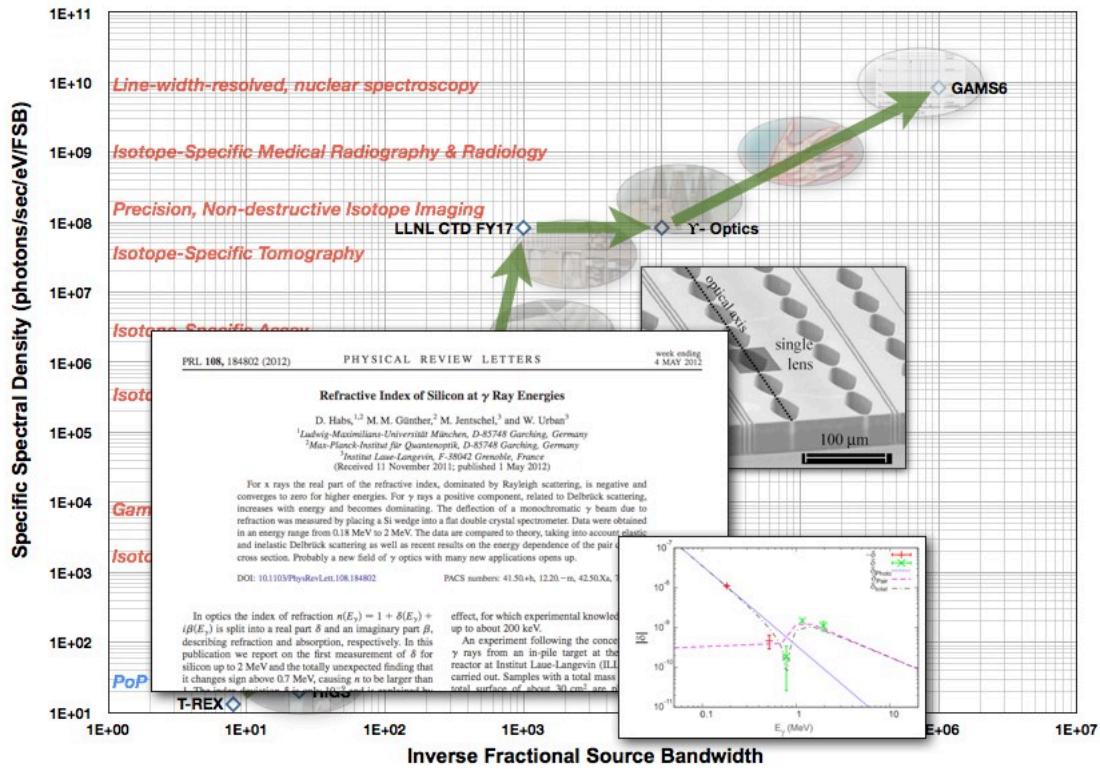
## A path to extreme MEGA-ray capability has been defined at LLNL



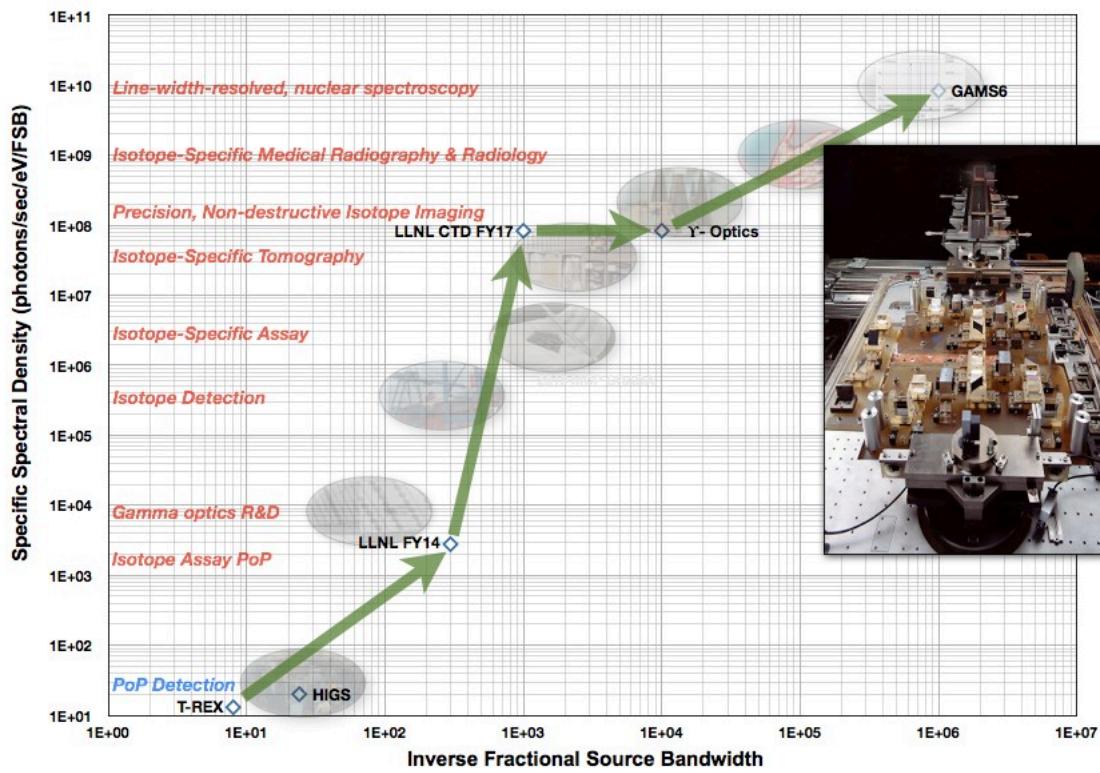
## A path to extreme MEGA-ray SSD has been defined at LLNL

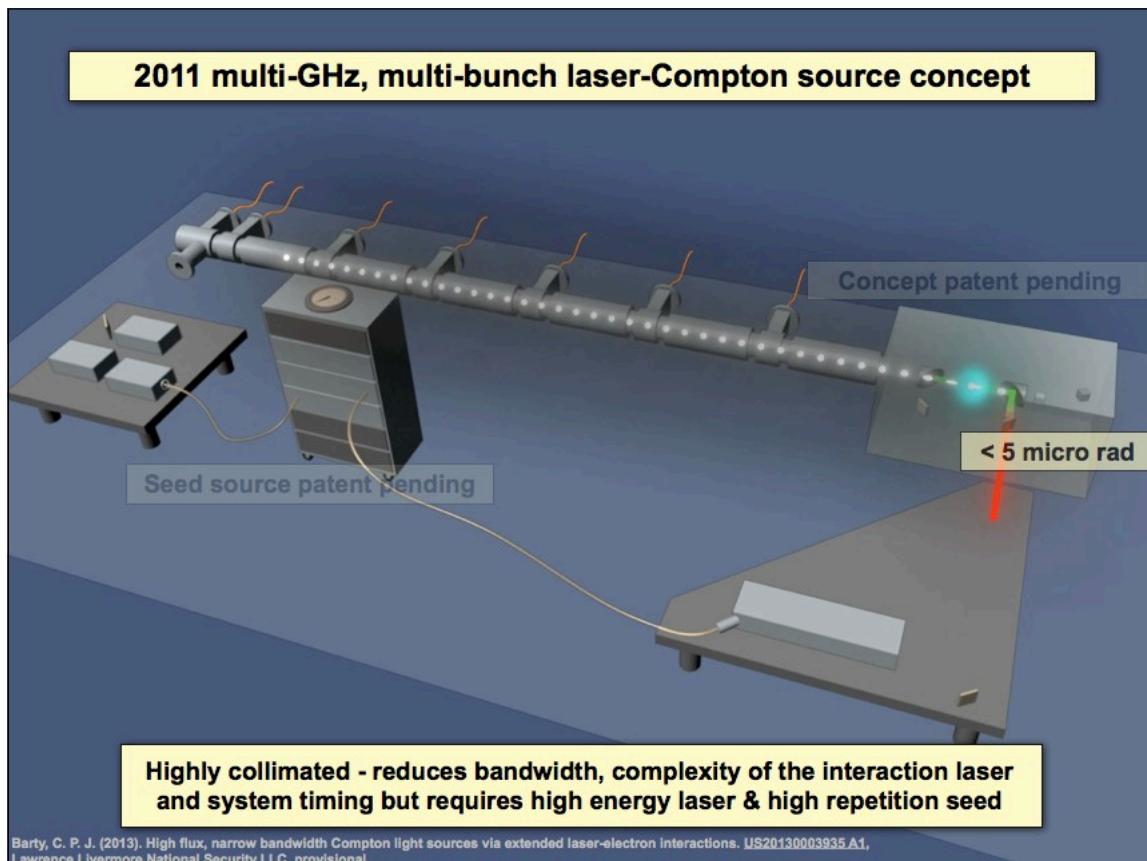


## A path to extreme MEGA-ray capability has been defined at LLNL

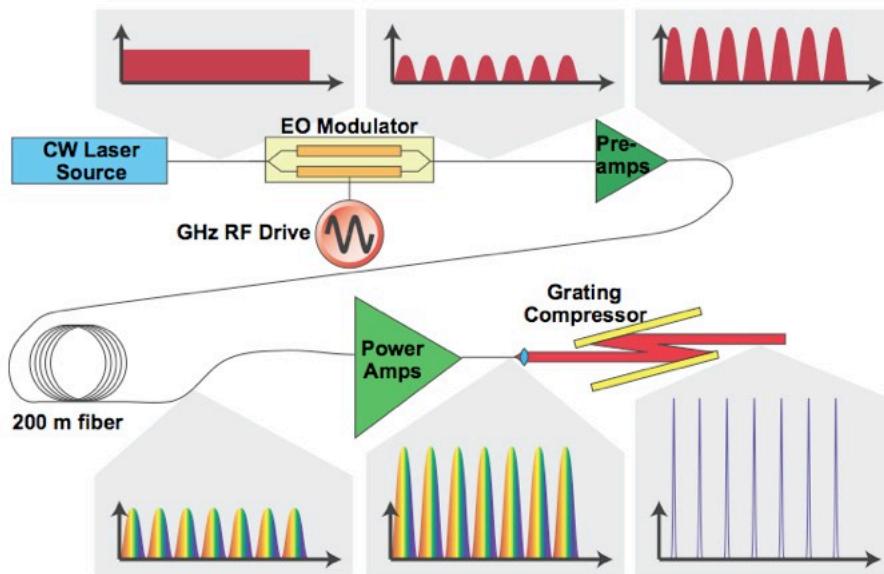


## A path to extreme MEGA-ray capability has been defined at LLNL



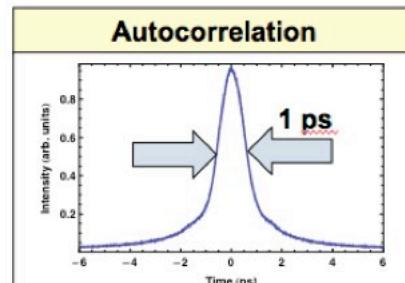
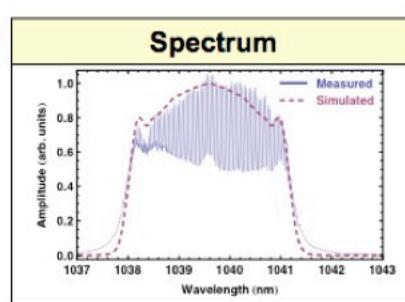
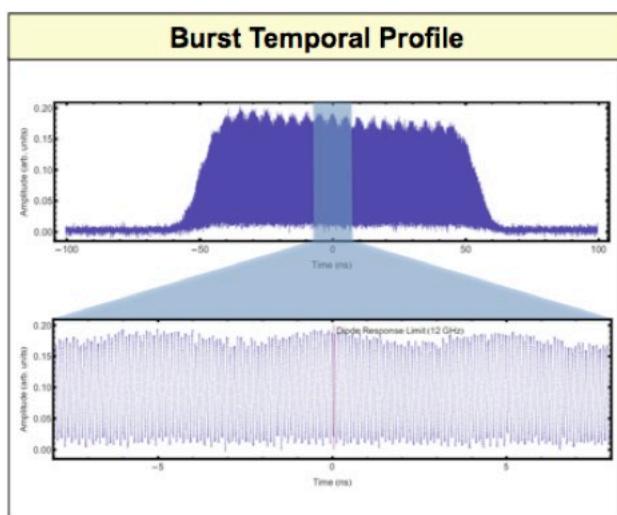


**"CW" method for generation of 11.424 GHz,  
synchronized train of picosecond IR seed pulses**



Barty, C. P. J., M. J. Messerly, J. W. Dawson, D. J. Gibson, M. A. Prantil, Eric Cormier (2012), Directly driven source of multi-gigahertz, sub-picosecond optical pulses. WO2013040041 A3, Lawrence Livermore National Security LLC, provisional

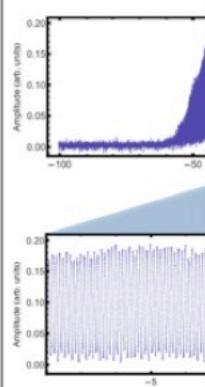
**"CW" method for generation of 11.424 GHz,  
synchronized train of picosecond IR seed pulses**



# **“CW” method for generation of 11.424 GHz, synchronized train of picosecond IR seed pulses**



Bur



## Widely tunable 11 GHz femtosecond fiber laser based on a non-mode-locked source

Matthew A. Prantil,<sup>1</sup>\* Eric Cormier,<sup>1</sup> Jay W. Dawson,<sup>1</sup> David J. Gibson,<sup>1</sup> Michael J. Messerly,<sup>1</sup> Christopher P. J. Barty,<sup>1</sup>

<sup>1</sup>Lawrence Livermore National Laboratory, Livermore, California 94550

\*Corresponding author: [prantil@llnl.gov](mailto:prantil@llnl.gov)

Received Month X, XXXX; revised Month X, XXXX; accepted Month X, XXXX

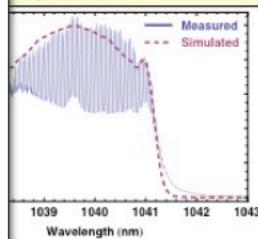
An 11 GHz fiber laser built on a modulated CW platform is described and characterized. This compact, vibration-tolerant system generates a synchronized train of picosecond pulses at 11.424 GHz with a repetition rate of 83 kHz. The RF signal directly, and is tunable over a wide range of drive frequencies. The demonstration system when operated at 1040 nm is capable of 50 fs bursts of 275 micro-pulses produced at a micro-pulse rate of 83 kHz where the macro-pulse and micro-pulse energies are 1.8 mJ and 3.2 nJ respectively. In micro-pulses are compressed to a duration of 850 fs.

OCIS Codes: 060.1510, 160.3990

Very-short-pulse laser sources (in the range of picoseconds to femtoseconds) with high repetition rates (100 MHz to tens of GHz) are needed to drive short-wave-length lasers, to generate high-intensity pulses for nonlinear optical parametric interactions, and as photo-cathode illuminators to create photo-electrons in high frequency particle accelerators. Other applications of these ultrafast pulsed lasers include materials processing, 3-D lithography, high-resolution fiber communication and remote sensing systems. The system architecture described here is wavelength compatible with high energy fiber technology, is driven directly by an RF source (and thus sidesteps synchronization issues), and allows for direct access to materials processing parameters. Moreover, the ability to modify electronically the temporal pulse shape, in amplitude and phase [1, 2], offers the possibility of controlling various complex photochemical processes and quantum control of interactions on nanosecond time scales.

We demonstrate a laser pulse train by operating a continuous-wave (cw) laser with an RF source. Several groups [2, 3] have converted cw lasers to subps, high frequency pulse trains; however, these groups have relied on “time-lens” techniques [4] to generate the level bandwidth that we used to accomplish at 1550 nm in specially optimized fibers to generate further bandwidth while simultaneously compressing the pulses. At the 1030–1070 nm wavelengths where we wish to work, the soliton compression scheme is not feasible because there is no inelastic scattering. However, opposite sign dispersion at 1550 nm – in the demonstration of this cw-modulation concept reported here, we rely on self-phase modulation (SPM) [5] to generate 3.2 nm of bandwidth and compress the pulse with a grating-based compressor. To reduce aliasing

## Spectrum



## Autocorrelation

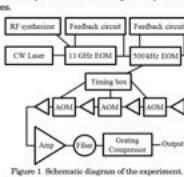
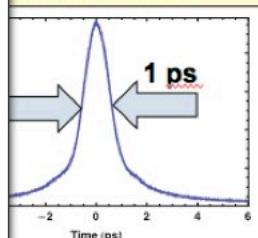
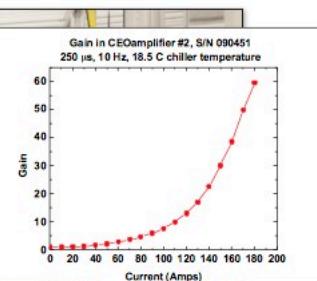
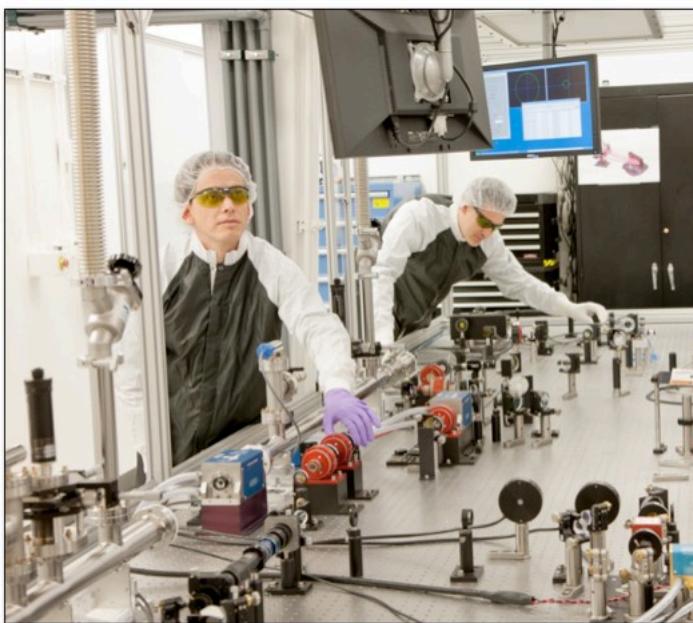


Figure 1 Schematic diagram of the experiment.

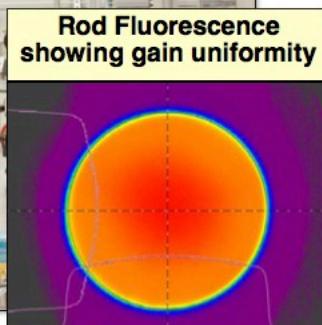
The architecture, shown in Figure 1, begins with a cw laser at 1040 nm. Velocity has been reduced to 1030 nm beam. The beam is sent through an EOSPACE-brand, Z-cut, 20 GHz, dual drive Mach-Zehnder electro-optic modulator (EOM) modulated by a control circuit (YY Labs, Inc.) to keep the modulator milli-based – that is, to move the light when the drive is driven. The modulator is driven with 20.1 dBs of 5.7 GHz RF power. Because of the mill bias, the RF creates an 11.4 GHz laser pulse train with 44 ps pulse length and no cw component as shown in Figure 2; the latter prevents stimulated Brillouin scattering from damaging subsequent fiber amplifiers. A second EOM temporally

Prantil, M. A., E. Cormier, J. W. Dawson, D. J. Gibson, M. J. Messerly and C. P. J. Barty. "Widely tunable 11 GHz femtosecond fiber laser based on a nonmode-locked source." *Optics Letters* 38(17): 3216-3218 (2013)

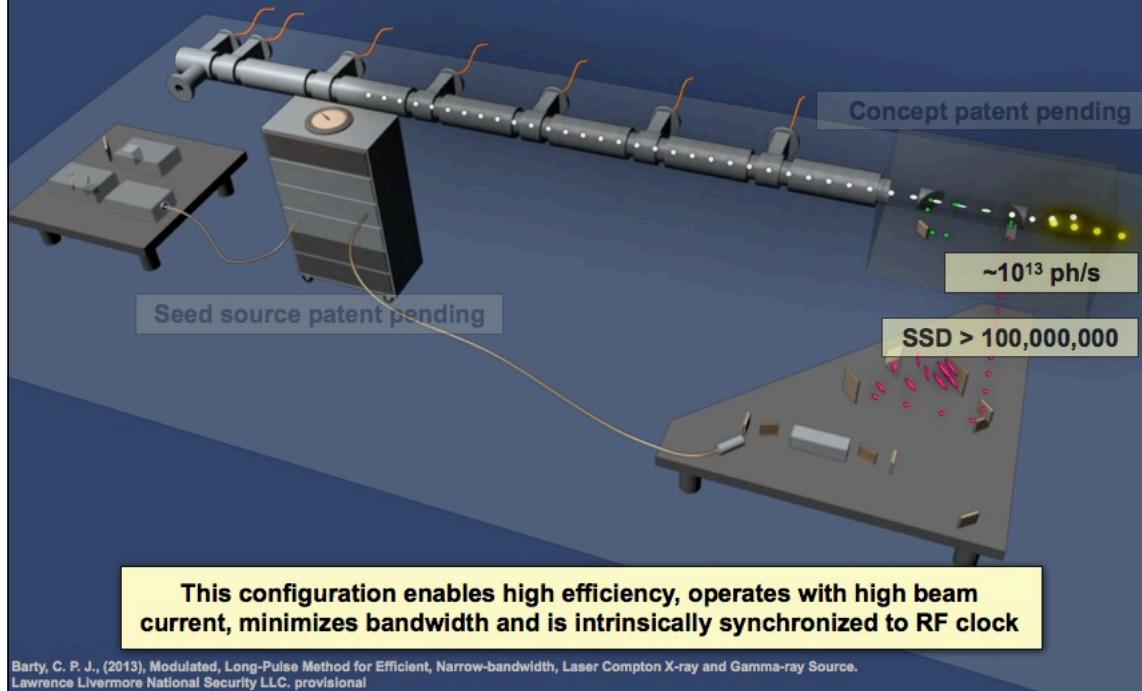
# A new custom, diode-pumped solid state laser architecture can generate > 1J per pulse @ 120 Hz



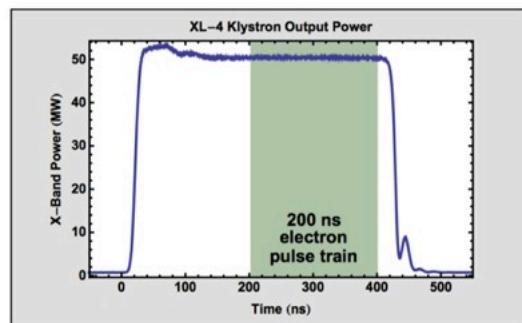
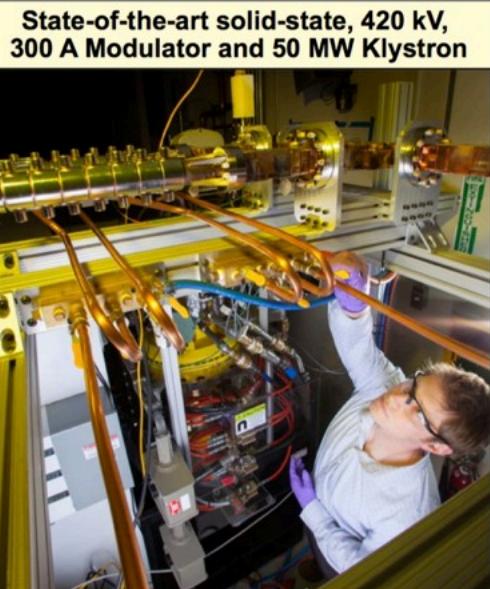
## Rod Fluorescence showing gain uniformity



## LLNL's "Picket Fence" multi-GHz, laser-Compton source concept



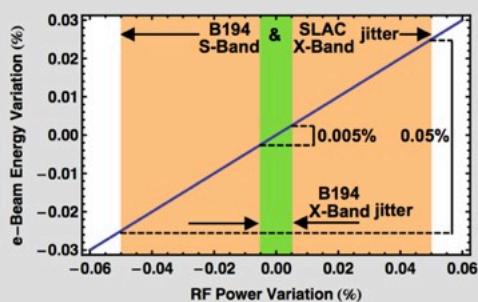
## RF Power combines the best of SLAC klystron technology & commercial solid state modulators



## Performance of the XL4 klystron and ScandiNova modulator exceed all of our requirements

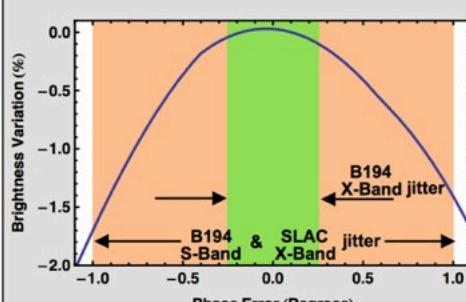


E-Beam Energy vs. RF Power



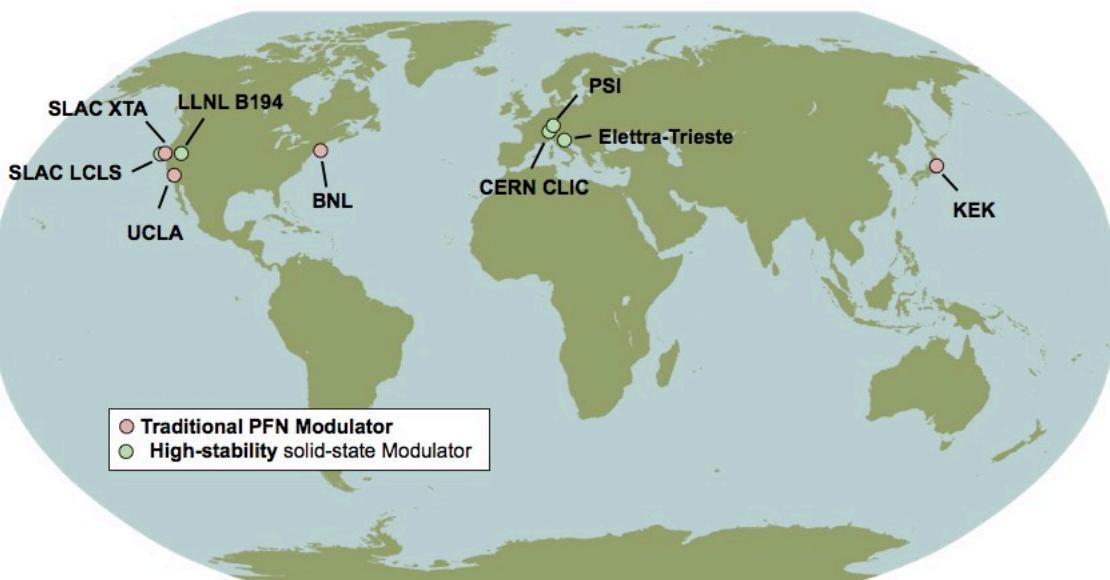
This system is 10x more stable in amplitude than existing SLAC and LLNL sources.

E-Beam Brightness vs. RF Phase



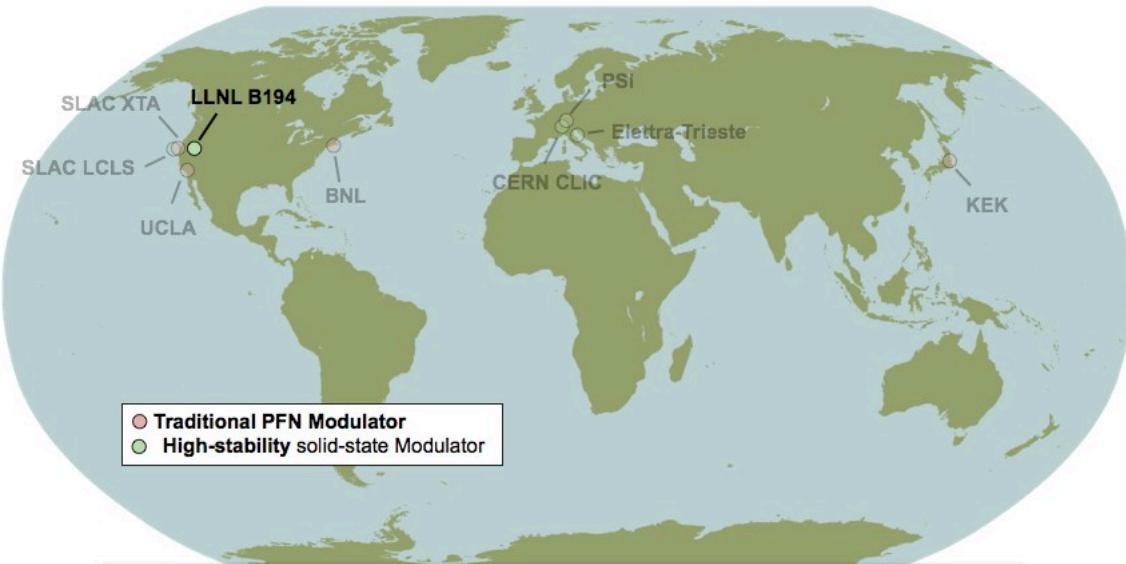
4x improvement in RF phase stability corresponds to 10x improvement in brightness stability.

## Worldwide high power x-band sources





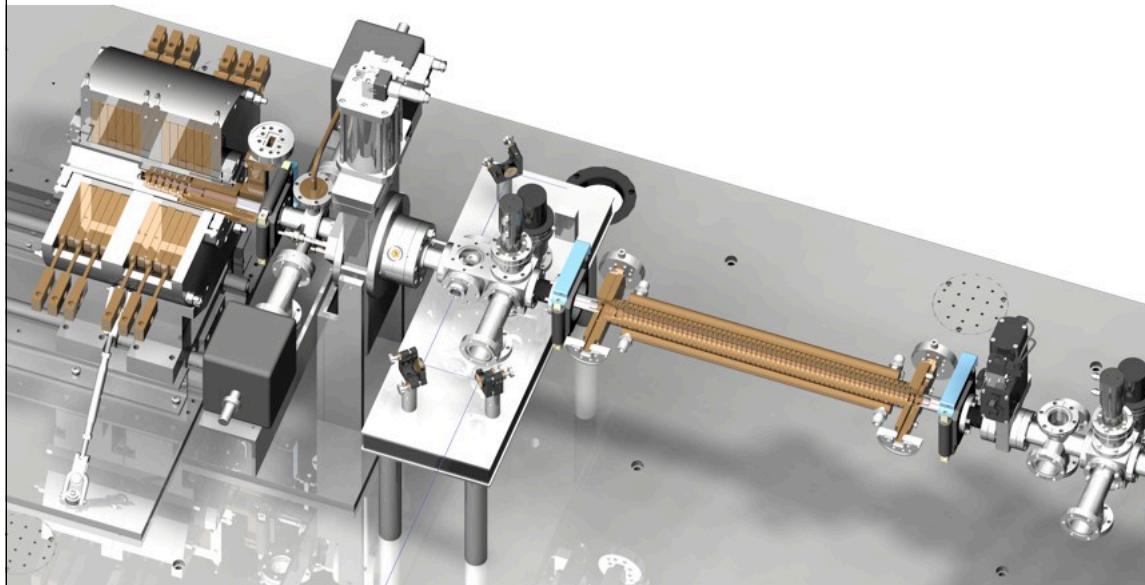
## Worldwide high power x-band sources



**LLNL set up is currently the only facility where high quality x-band RF is coupled with state-of-the-art structures to produce beam**



## Photo-gun and first section



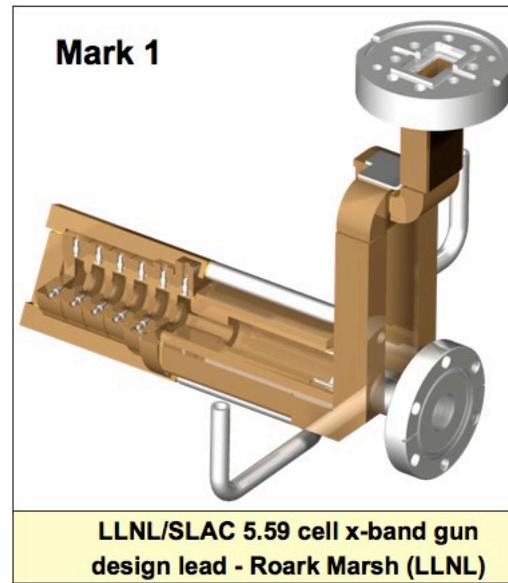


## X-band photo-gun evolution



## X-band photo-gun evolution

- Longer Half cell for lower final emittance
- Better mode separation for less mode beating on cathode surface
- Elliptical irises for lower peak surface electric field
- Dual feed racetrack coupler for minimized RF quadrupole kick
- Optimized beta for a balance of fast gun fill time and low pulsed heating

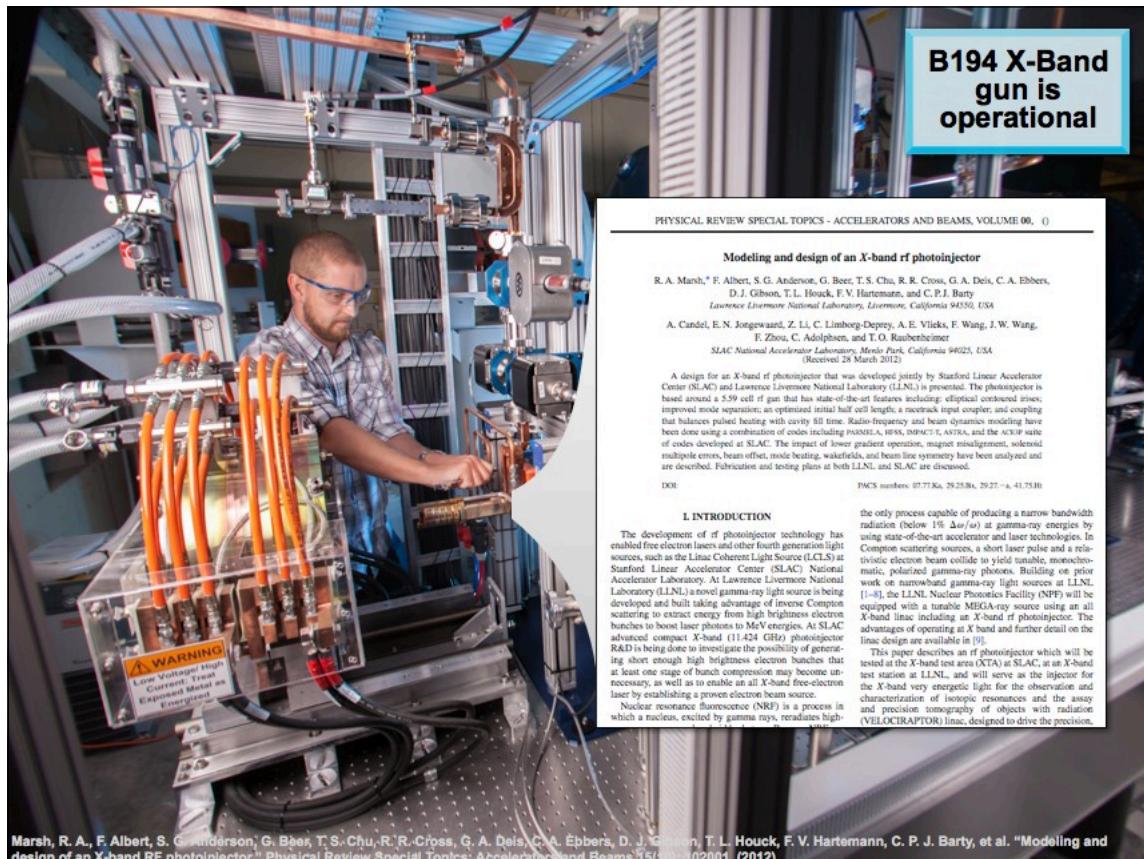
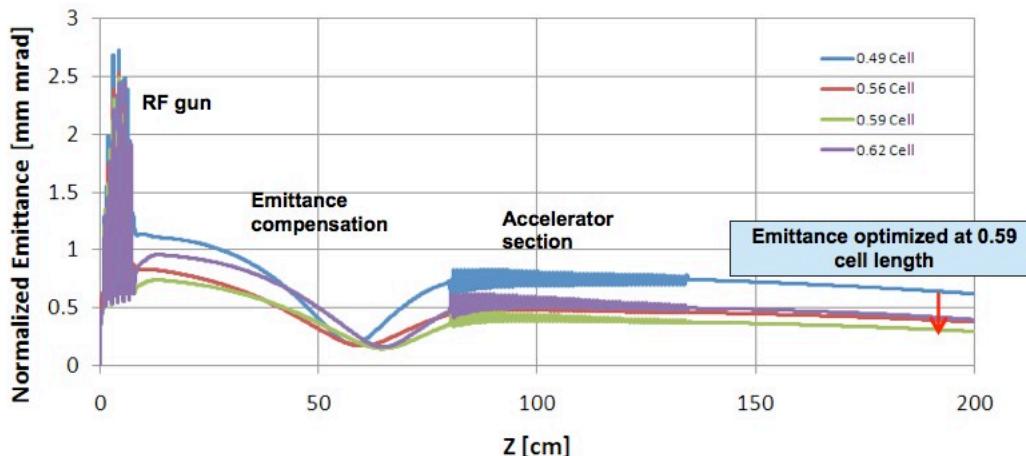


## PARMELA results

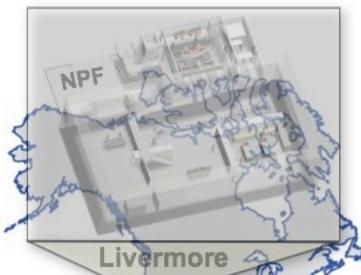
### Redesigned longer half cell optimizes the electron beam brightness and reduces Compton bandwidth



Optimized launch phase and solenoid strength  
Beam parameters:  $Q = 250 \text{ pC}$ ,  $\tau\varphi = 10 \text{ deg}$ .  
 $200 \text{ MV/m}$  cathode field



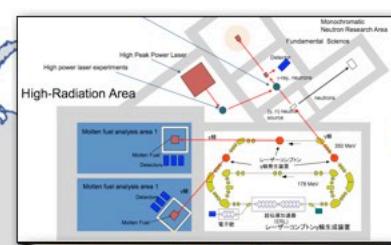
## Other next generation laser-Compton gamma-ray projects are emerging to pursue isotope science



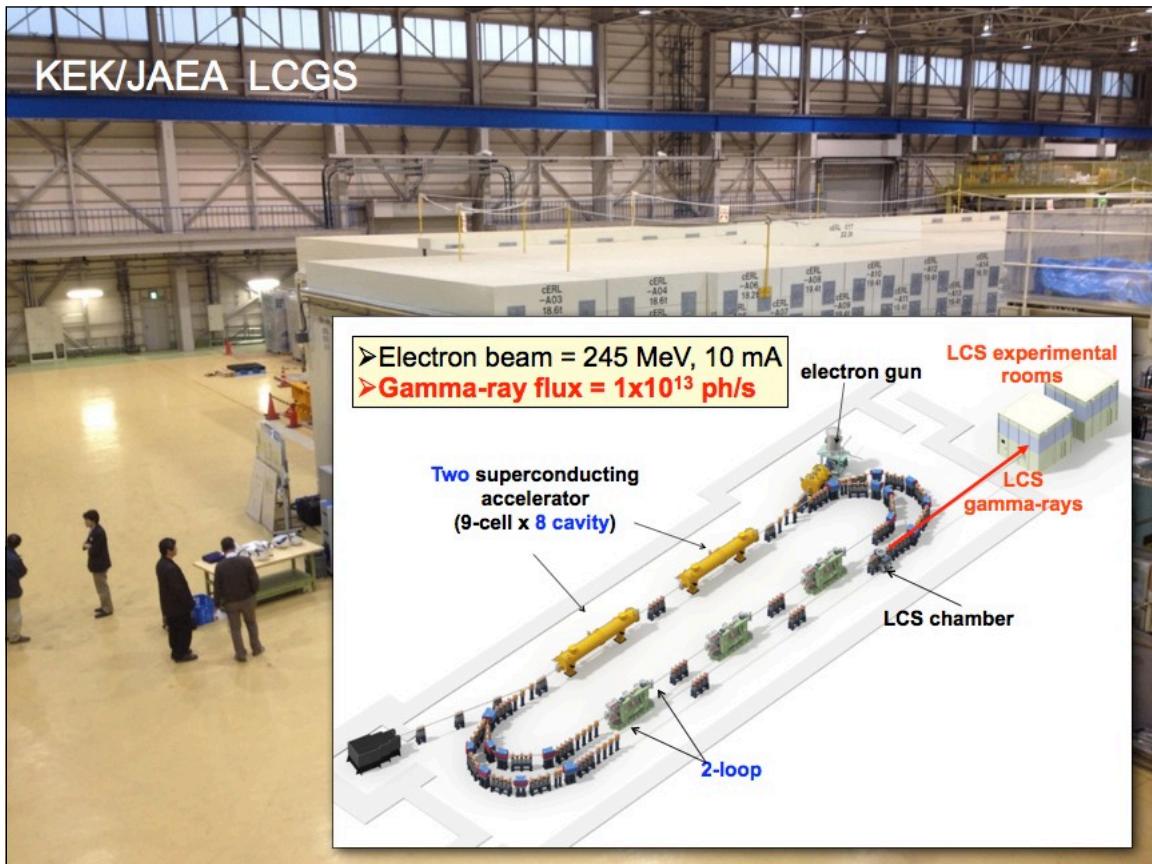
Bucharest



## Other next generation laser-Compton gamma-ray projects are emerging to pursue isotope science



Bucharest







**el**  
extreme light infrastructure

**Infrastructure Producing High Intensity Gamma Rays  
for ELI Nuclear Physics Bucharest-Magurele,  
Romania**

The ELI-Gamma Source working group

**Editors:**  
Dietrich Habs, Ludwig-Maximilians-Universität München  
Marian Toma, National Institute for Laser, Plasma and Radiation Physics  
Dan Cutoiu, National Institute for Physics and Nuclear Engineering "Horia Hulubei"

**Authors**  
G. Wormser<sup>1</sup>, R. Hajima<sup>2</sup>, C. Barty<sup>3</sup>

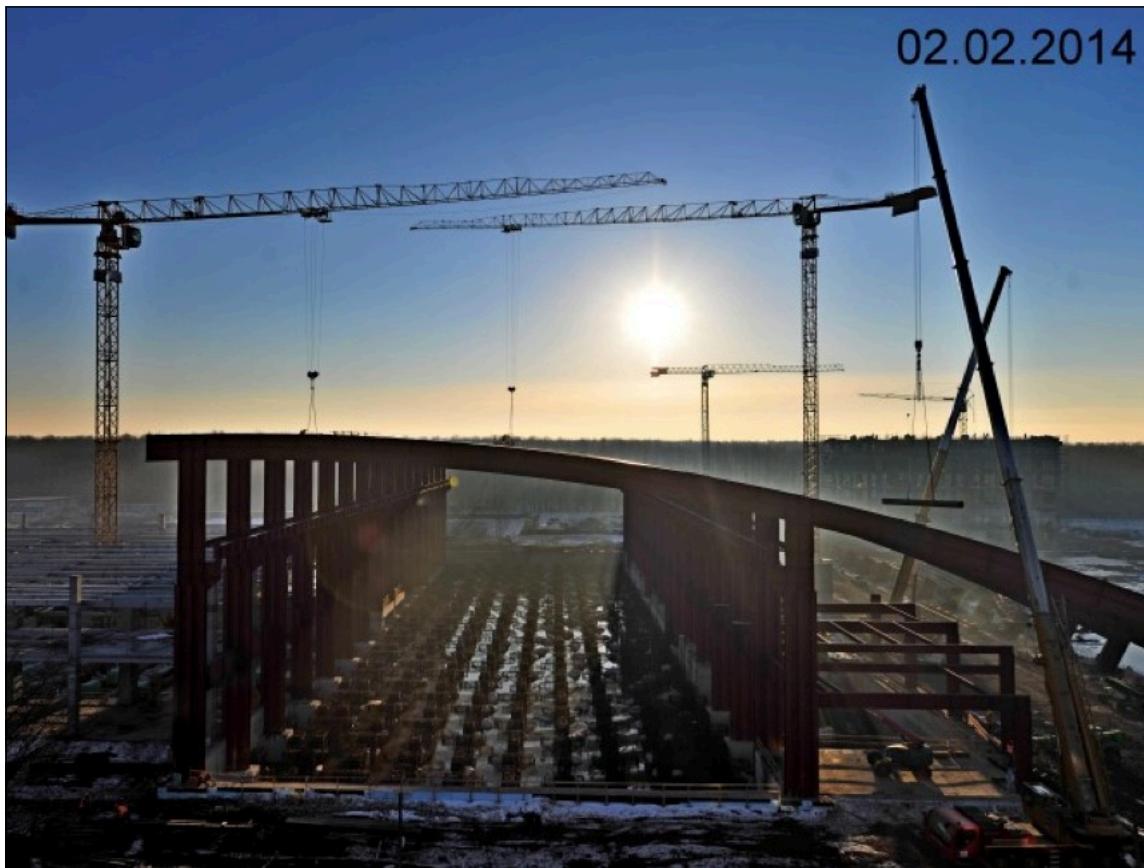
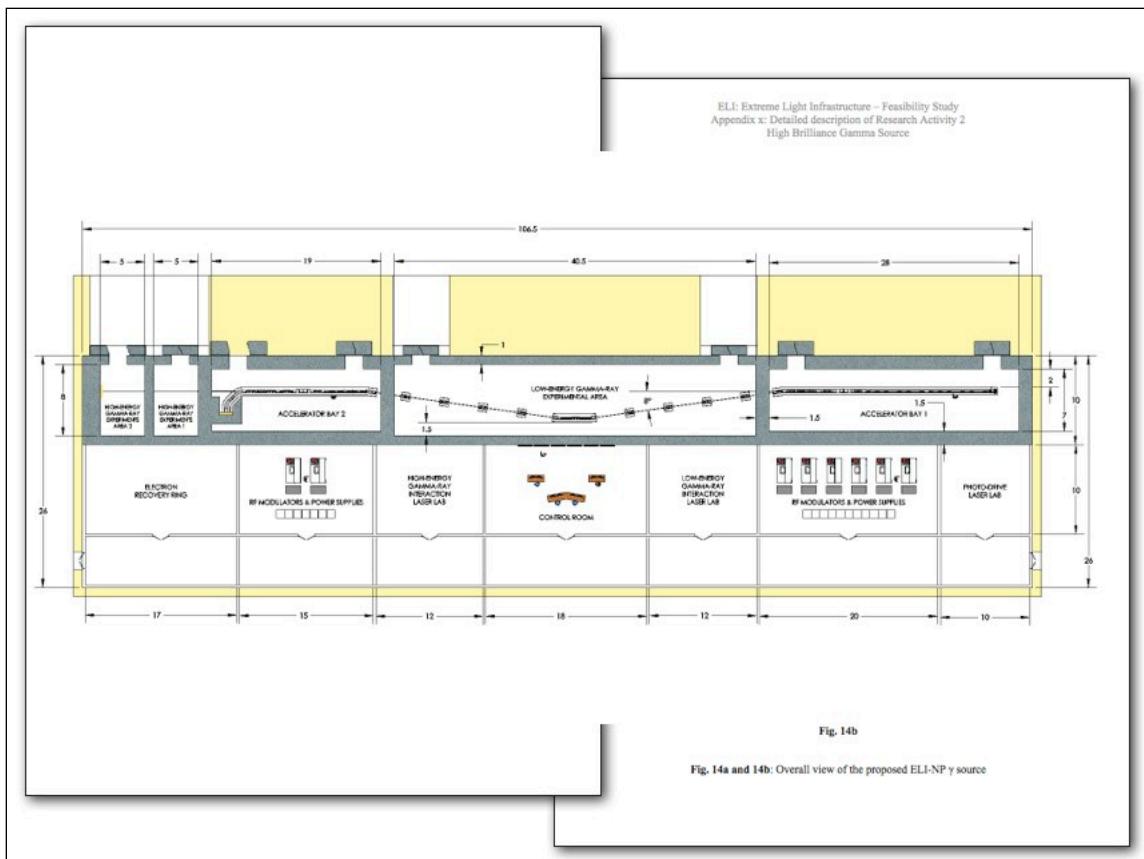
**Affiliations**  
<sup>1</sup> Linear Accelerator Laboratory, Orsay, France  
<sup>2</sup> ERL Development Group, Japan Atomic Energy Agency  
<sup>3</sup> Lawrence Livermore National Laboratory, USA

ELI: Extreme Light Infrastructure – Feasibility Study  
Appendix x: Detailed description of Research Activity 2  
High Brilliance Gamma Source

**Fig. 14a**

**Fig. 14b**

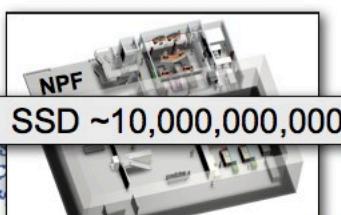
**Fig. 14a and 14b: Overall view of the proposed ELI-NP  $\gamma$  source**



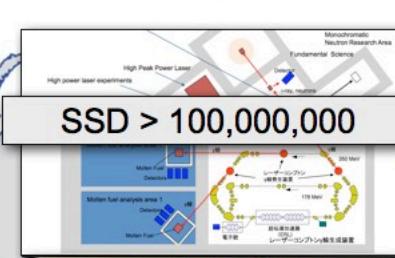
2017



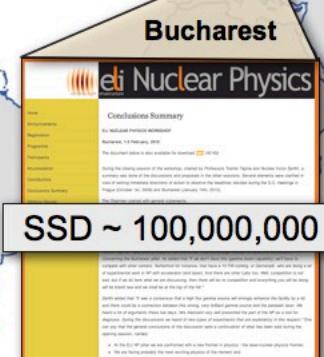
**Other next generation MEGa-ray projects are now emerging to pursue isotope science & applications**



Livermore



Fukushima



B194 X-Band  
x-ray  
Test Station

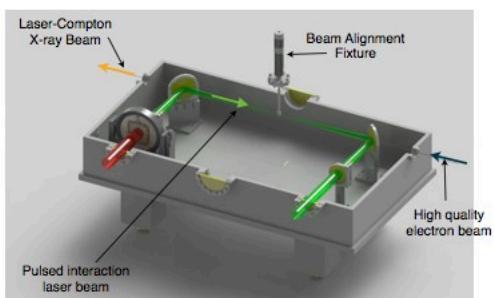
Scale of a clinical x-ray machine

LLNL's compact laser-Compton gamma-ray technology is also ideal for production of bright, high-flux, tunable, mono-energetic X-RAYS and thus opens new possibilities

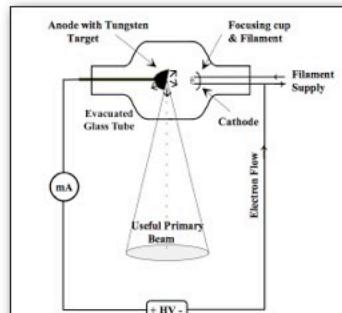
### Optimized laser-Compton x-ray technology could revolutionize medical radiography & radiology



- Mono-energetic (reduces dose for all x-ray radiographic procedures)

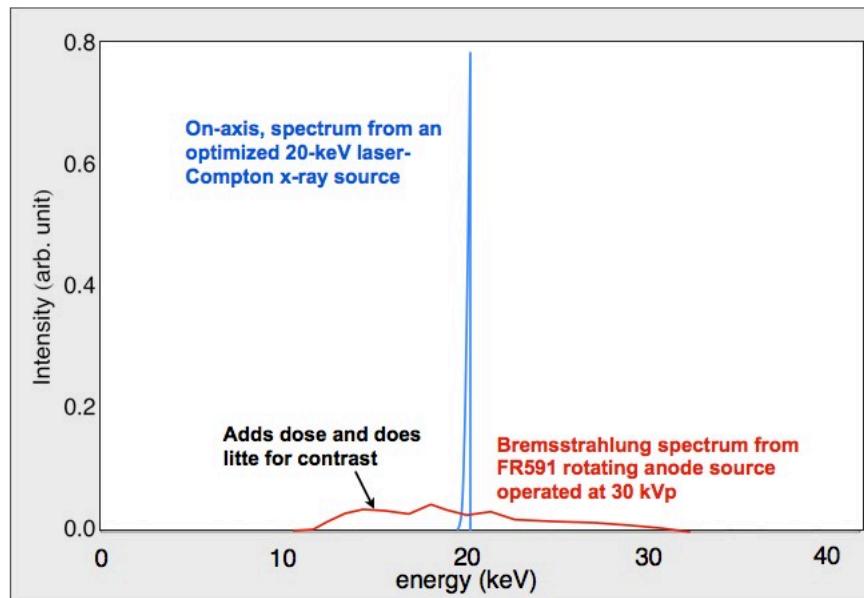


Laser-Compton X-ray Source



Rotating Anode X-Ray Source

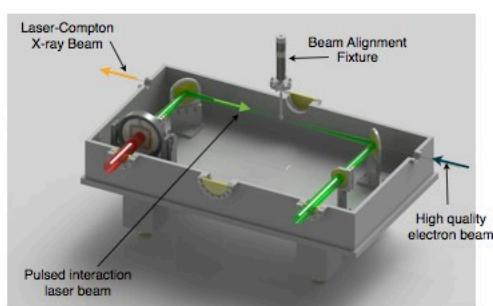
## The mono-energetic spectrum of laser-Compton source vs. rotating anode enables dose reduction



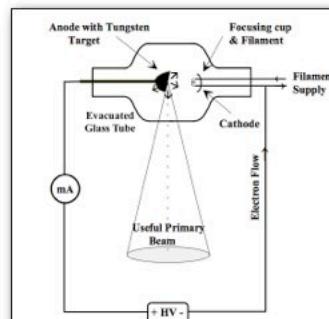
## Optimized laser-Compton x-ray technology could revolutionize medical radiography & radiology



- Mono-energetic (reduces dose for all x-ray radiographic procedures)
- Highly collimated (enables multi-angle radiology & dose localization)
- Superior spatial resolution (enables small subject radiography)
- Instantaneously adjustable output (enables minimum dose radiography)
- High flux (enables fast imaging and treatment times)
- Easily tunable (enables k-edge imaging and Auger-cascade therapy)

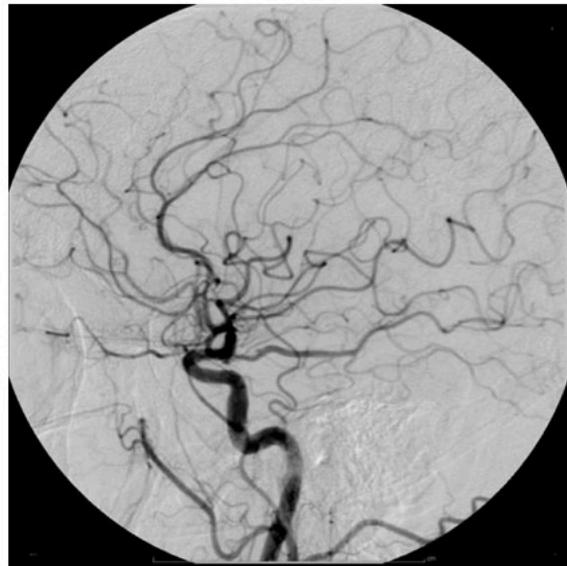


Laser-Compton X-ray Source



Rotating Anode X-Ray Source

**Two color subtraction imaging has been studied  
for coronary angiography**



**Mono-energetic imaging has been investigated  
using the output of synchrotron sources**

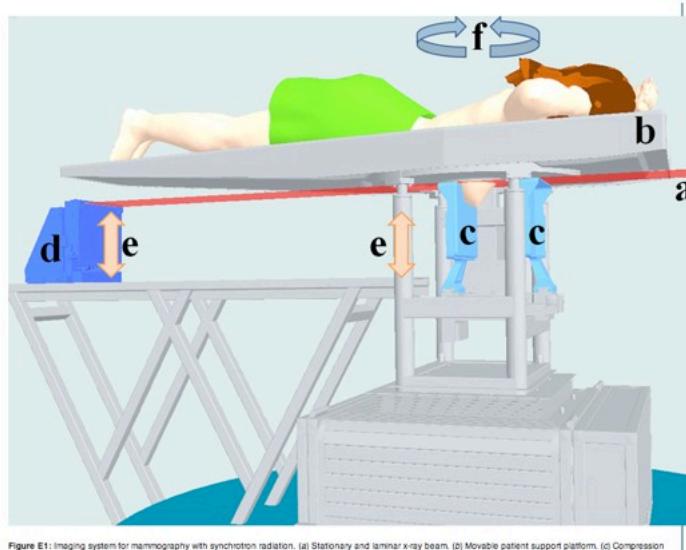
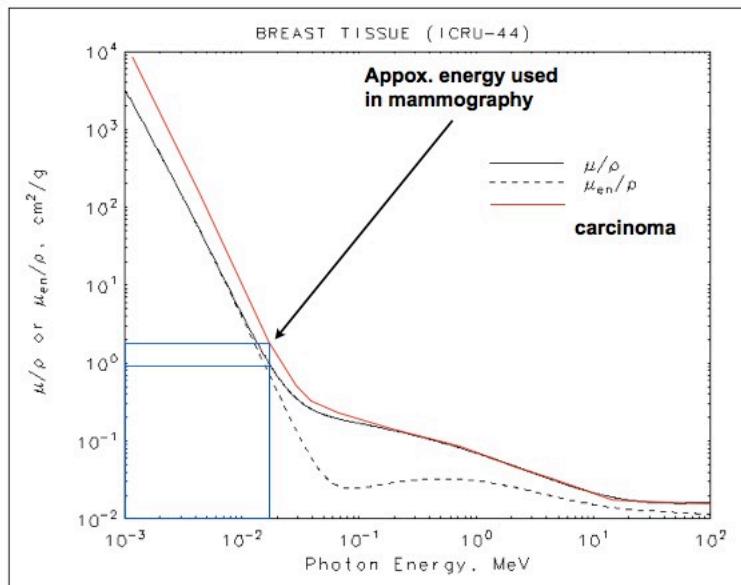


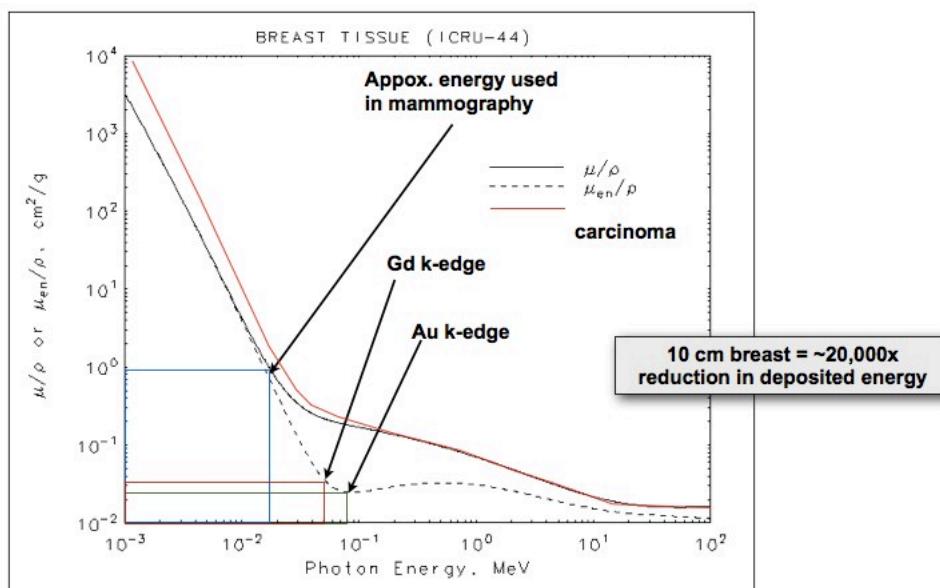
Figure E1: Imaging system for mammography with synchrotron radiation. (a) Stationary and laminar x-ray beam. (b) Movable patient support platform. (c) Compression system. (d) Holder for screen-film cassette. (e) Vertical movement of patient and screen-film system through the beam, which facilitates the acquisition of planar radiographic images. (f) Rotational movement of patient, which facilitates the acquisition of craniocaudal and mediolateral oblique views.

**From Italian mammography study with synchrotron radiation. Note  
this geometry does not require breast compression**

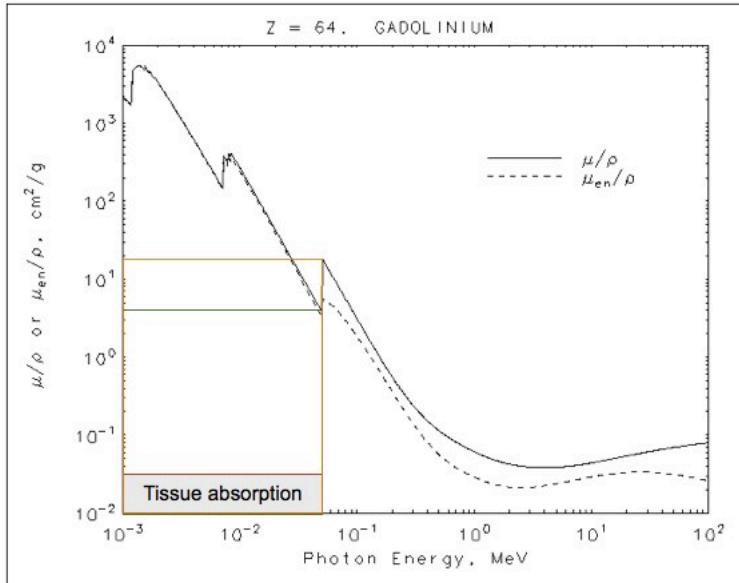
## Conventional breast tissue imaging is a compromise between contrast and dose



## Conventional breast tissue imaging is a compromise between contrast and dose



**Gd has excellent k-edge properties and is already an approved contrast agent for breast MRI**



### **Optimized laser-Compton x-ray technology could revolutionize medical radiography & radiology**

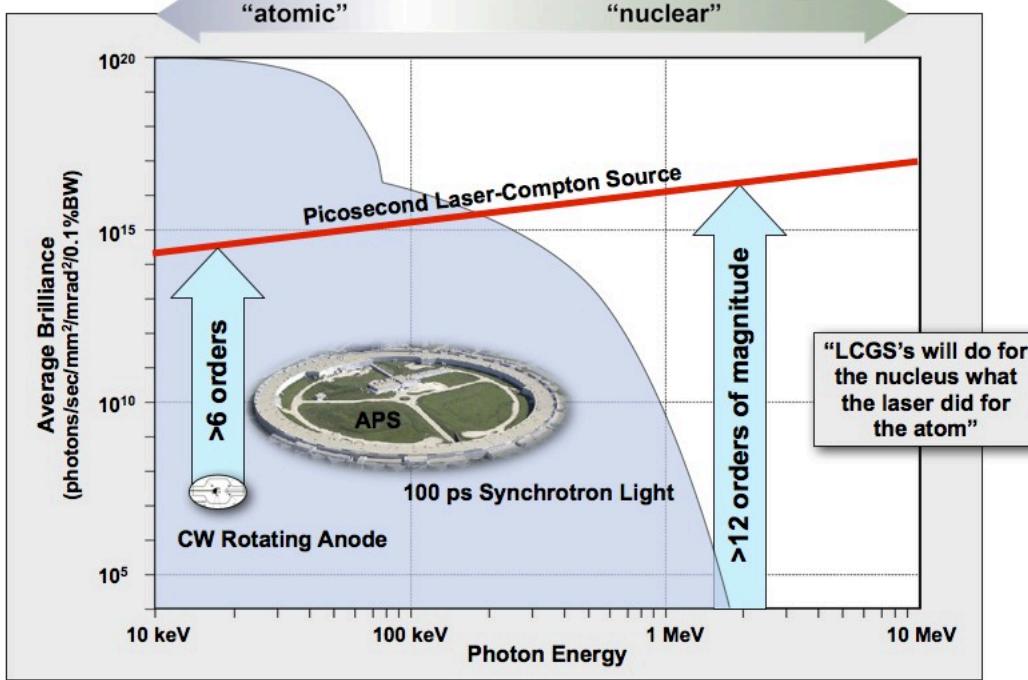


- Mono-energetic (reduces dose for all x-ray radiographic procedures)
- Highly collimated (enables multi-angle radiology & dose localization)
- Superior spatial resolution (enables small subject radiography)
- Instantaneously adjustable output (enables minimum dose radiography)
- High flux (enables fast imaging and treatment times)
- Easily tunable (enables k-edge imaging and Auger-cascade therapy)

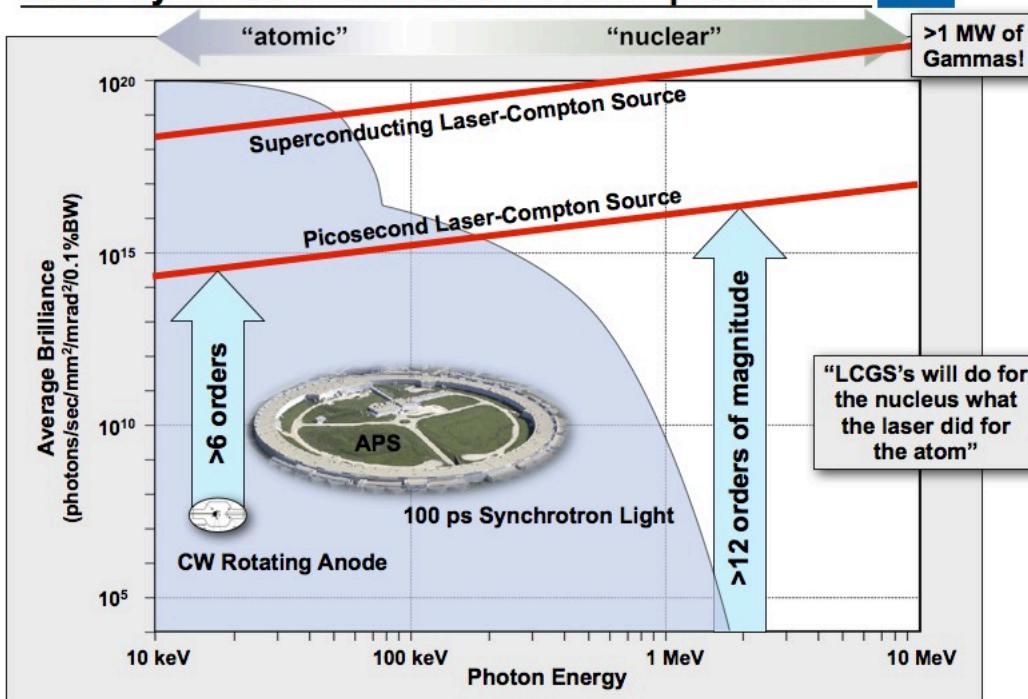
Attribute	LLNL Picket Fence c-2013	LLNL RING c-2013	LLNL T-REX c-2009	MXIS/Vanderbilt c-2007	Rotating Anode	Units
bandwidth	~0.1%	<0.5%	8%	10%	100%	$\Delta E/E$
collimation	~0.5	1	1	2	524	mrad
source size	5	15	30	35	150	microns
average brilliance	$>1E+13^*$	$>1E+13^*$	$8.89E+10$	$3.60E+06$	$6.00E+07$	$ph/s/mm^2/mrad^2/0.1\%BW$
e-beam current	6.00E-03	6.00E-04	1.50E-06	2.08E-09	100	mA
laser power	1200	120	2.5	0.0036	n/a	W

- These attributes are important to many other applications besides medicine

**Optimized laser-Compton sources will enable “lab-scale” synchrotron science & “nuclear” photonics**



**Optimized laser-Compton sources will enable “lab-scale” synchrotron science & “nuclear” photonics**



The Laser-Compton and Nuclear Photonics efforts described in  
this presentation represent contributions from **12** institutions

Marvin	Adams	TAMU
Chris	Adolphsen	SLAC
Felicie	Albert	LLNL
Gerry	Anderson	LLNL
Scott	Anderson	LLNL
Paul	Armstrong	LLNL
Chris	Barty	LLNL
Andy	Bayramian	LLNL
Bret	Beck	LLNL
Glenn	Beer	LLNL
Shawn	Betts	LLNL
Dave	Boyle	TAMU
Patrick	Brantley	LLNL
Eugene	Brooks	LLNL
Arno	Candel	SLAC
Bill	Charlton	TAMU
Sam	Chu	SLAC
Eric	Cormier	UBordeaux
Rick	Cross	LLNL
Dan	Cutiou	ELI-NP
Gary	Deis	LLNL
Bob	Demaret	LLNL
Shawn	Densberger	LLNL
Valery	Dolgashev	SLAC
Chris	Ebbers	LLNL
Mike	Fazio	SLAC
Diana	George	LLNL
David	Gibson	LLNL
Marc	Gunther	LMU
Dietrich	Habs	LMU
Chris	Hagmann	LLNL
Ryoichi	Hajima	JAEA
James	Hall	LLNL
Fred	Hartemann	LLNL
Corrine	Izak	CEA
Michael	Jentschel	ILL
Micah	Johnson	LLNL
Ed	Jones	LLNL
Erik	Jongewaard	SLAC
Zenghai	Li	SLAC
Cecile	Limborg-Deprey	SLAC
Roark	Marsh	LLNL
Scott	McKinley	LLNL
Dennis	McNabb	LLNL
Jim	Morel	TAMU
Ed	Morse	UCB
Kaila	O'Neil	LLNL
Henry	Phan	LLNL
Norbert	Pietralla	GSI
John	Post	LLNL
Matt	Prantil	LLNL
Cesar	Pruneda	LLNL
Sofia	Quaglioni	LLNL
Tor	Raubenheimer	SLAC
Vladimir	Semenov	LLNL
Michio	Seya	JAEA
Rich	Shuttleworth	LLNL
David	Stevens	LLNL
Sami	Tantawi	SLAC
Peter	Thiorlf	LMU
Arnold	Vlieks	SLAC
Faya	Wang	SLAC
Juwen	Wang	SLAC
Caroline	Winters	LLNL
Sheldon	Wu	LLNL
Victor	Zamfir	ELI-NP
Feng	Zhou	SLAC

